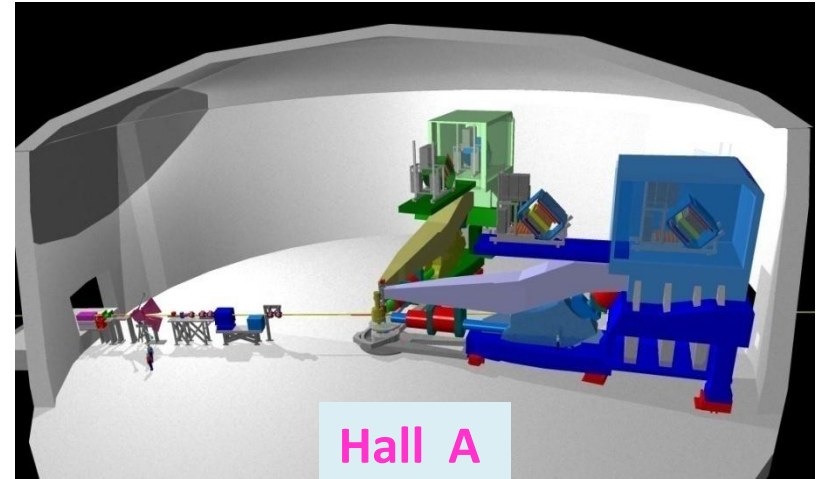


Parity-Violating Measurements of the Weak Charge of ^{208}Pb (PREX) & ^{48}Ca (CREX)

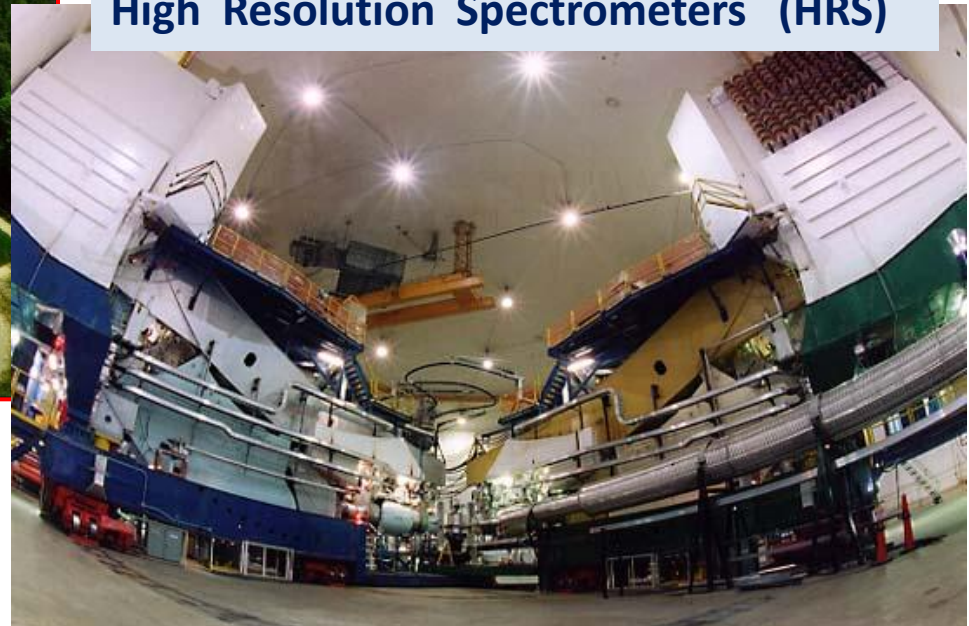


.... and possible future measurements

Hall A at Jefferson Lab



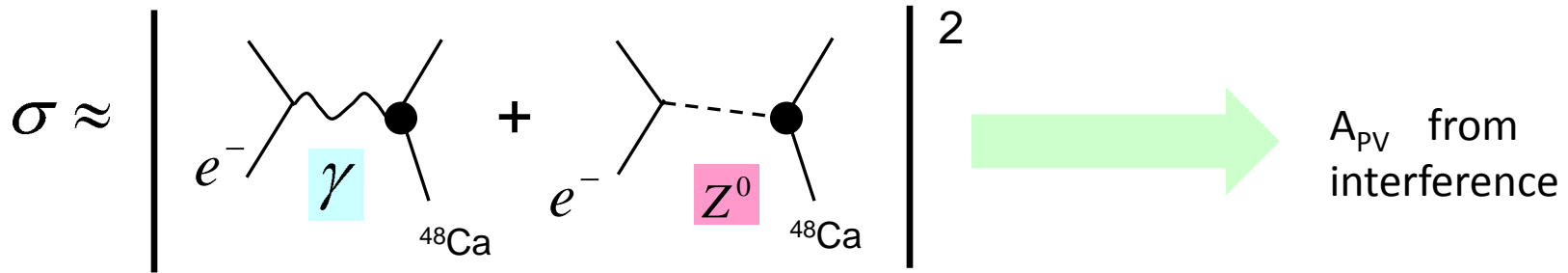
High Resolution Spectrometers (HRS)



Parity Violating Asymmetry

Elastic scattering of electrons

$$A_{PV} = \frac{\sigma_R - \sigma_L}{\sigma_R + \sigma_L} \sim 10^{-4} \times Q^2$$



Parameter	PREX-II (Pb)	CREX (Ca)
Beam Energy (GeV)	1.0	2.2
Angle	5°	4°
Asymmetry (ppm)	0.6	2.2
Asy Stat. Error	3%	2.4%
Error in R_N (fm)	0.06	0.02
Beam Time (days)	35	45

septum magnet

1% polarimetry

CREX / PREX : Z^0 of weak interaction : sees the **neutrons**

	proton	neutron
Electric charge	1	0
Weak charge	0.08	1

T.W. Donnelly, J. Dubach, I. Sick
Nucl. Phys. A 503, 589, 1989

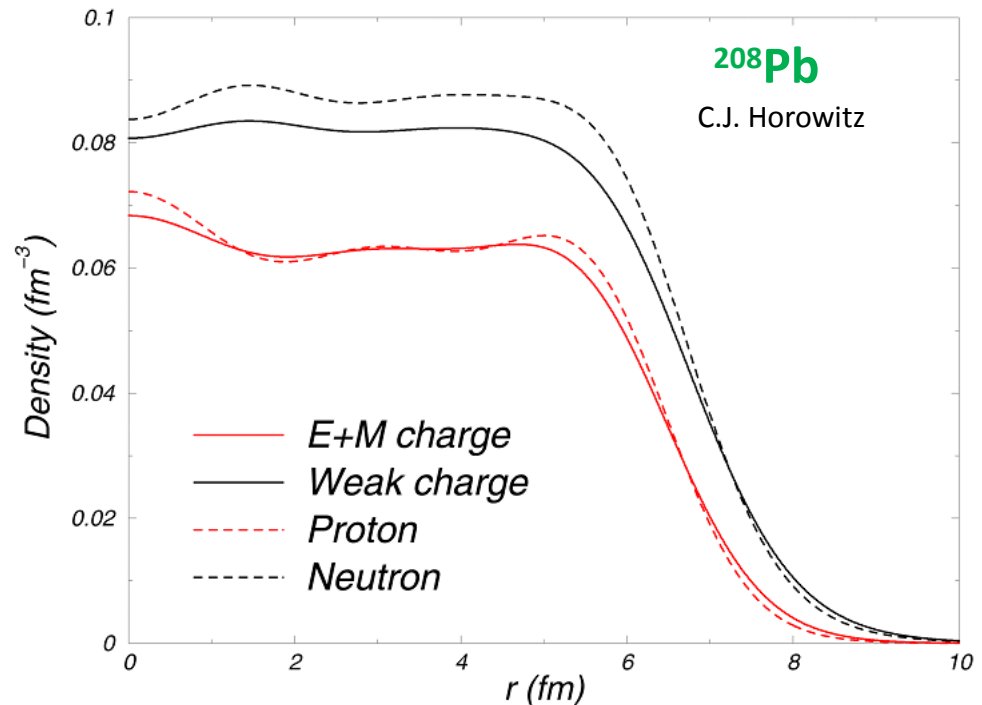
C. J. Horowitz, S. J. Pollock, P.
A. Souder, R. Michaels
Phys. Rev. C 63, 025501, 2001

Neutron form factor

$$F_N(Q^2) = \frac{1}{4\pi} \int d^3r j_0(qr) \rho_N(r)$$

Parity
Violating
Asymmetry

$$A = \frac{G_F Q^2}{2\pi\alpha\sqrt{2}} \left[\underbrace{1 - 4\sin^2\theta_W}_{\approx 0} - \frac{F_N(Q^2)}{F_P(Q^2)} \right]$$



PREX II ²⁰⁸Pb and CREX ⁴⁸Ca

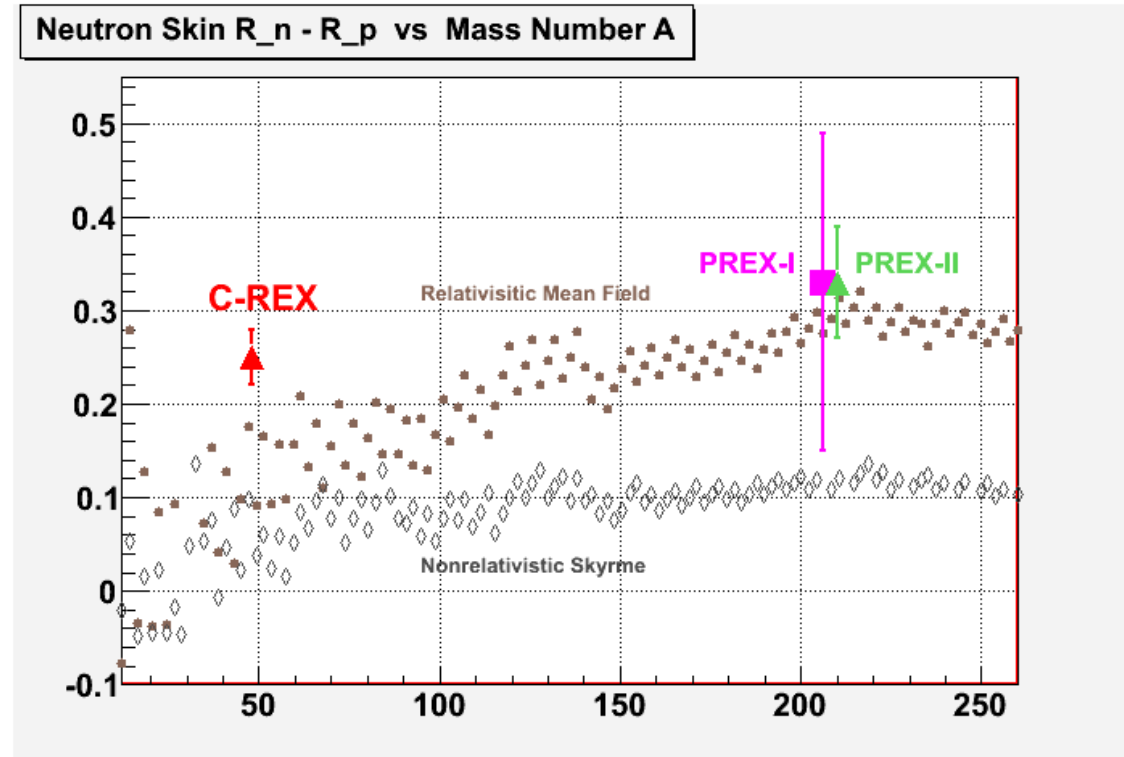
²⁰⁸Pb skin approximates infinite nuclear matter ~ small neutron star

⁴⁸Ca smaller, optimizes at higher Q² (a natural 11 GeV expt).



Lots of neutrons, sensitive to 3N force

Tests microscopic models for neutrons.



Want doubly-magic nuclei: little coupling to inelastic -- “clean”

Want neutrons: weak interaction couples to neutrons

C-REX Workshop

<http://www.jlab.org/conferences/crex/program.html>

at JLab March 17 – 20, 2013

40 Participants from several fields 19 talks & discussion



C-REX Workshop

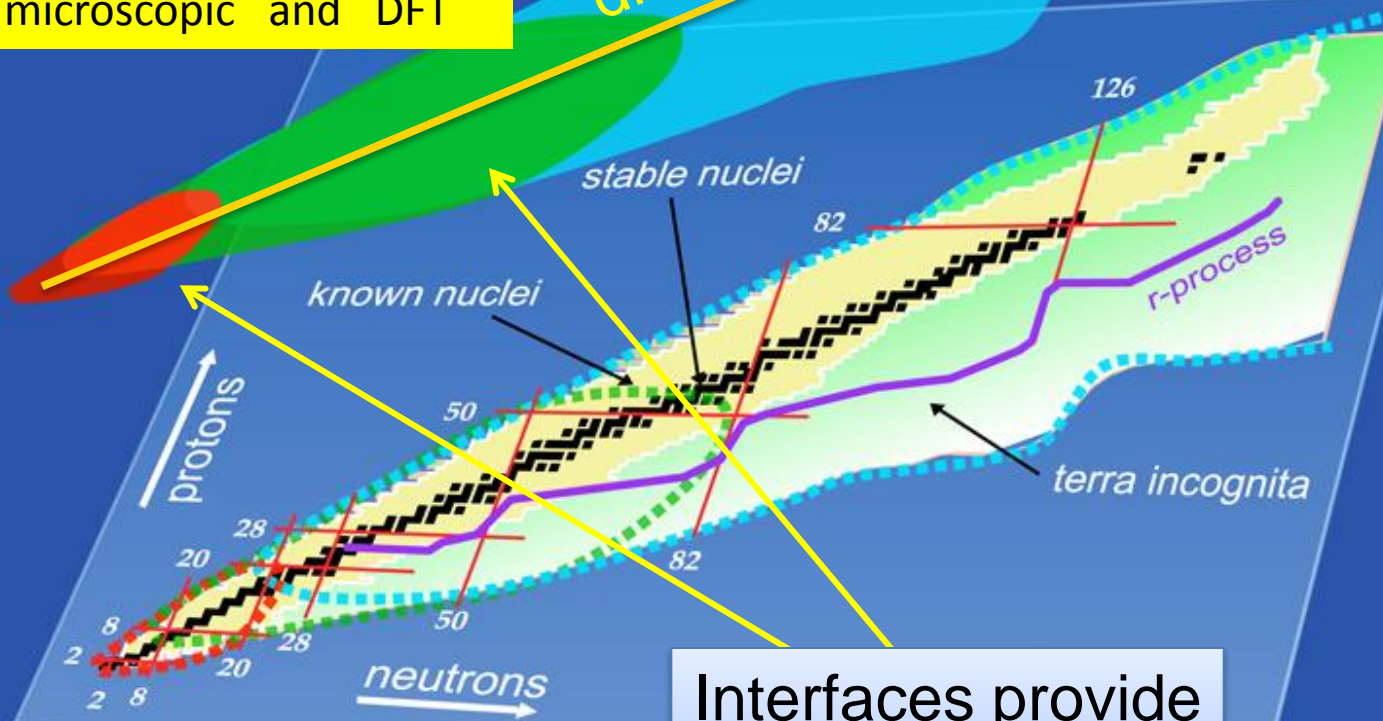
Jefferson Lab
March 17-19, 2013

Nuclear Landscape

- Ab initio
- Configuration Interaction
- Density Functional Theory

^{48}Ca : bridge between microscopic and DFT

dimension of the problem

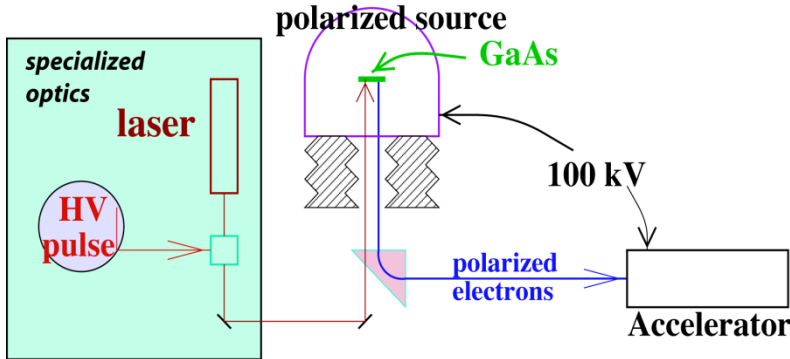


Interfaces provide crucial clues



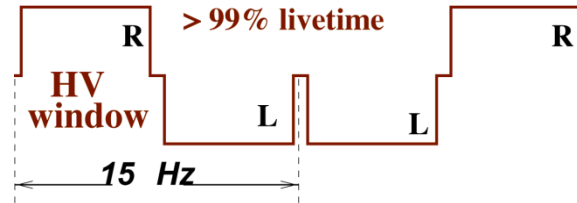
How to do a Parity Experiment

(integrating method)



rapid, random, helicity flipping

Rapid, Random Helicity Flips



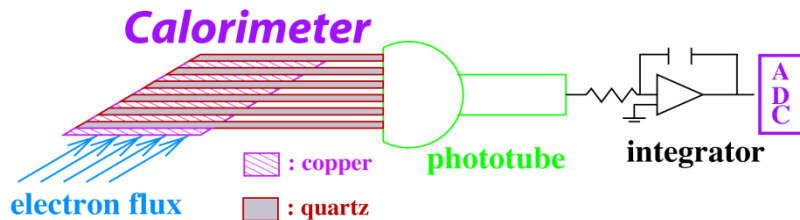
Measure flux F for each window

$$A_{\text{window pair}} = \frac{F_R - F_L}{F_R + F_L}$$

Flux Integration Technique:

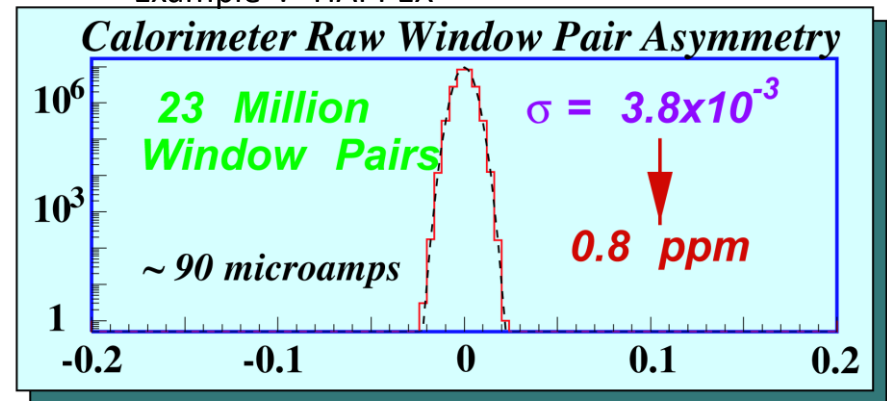
HAPPEX: 2 MHz

PREX: 500 MHz



Signal Average N Windows Pairs: $A \pm \frac{\sigma(A)}{\sqrt{N_{\text{windows}}}}$

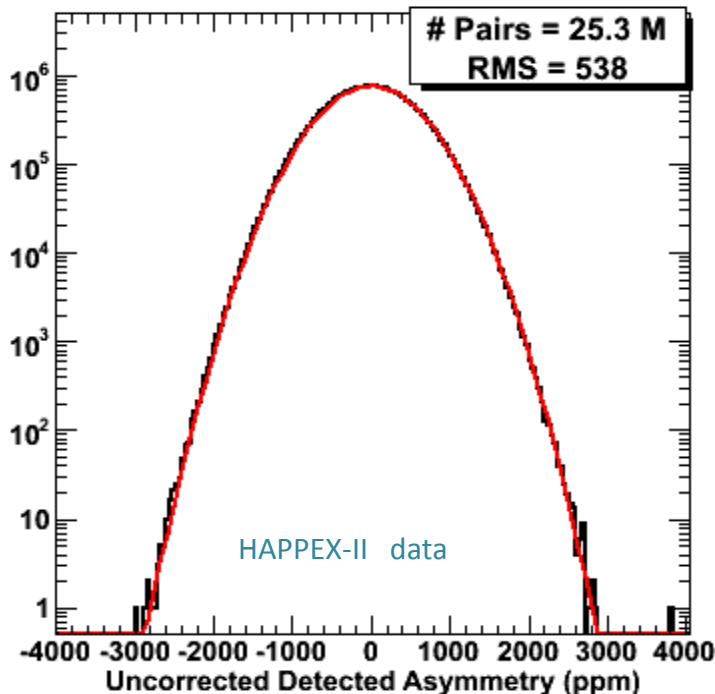
Example : HAPPEX



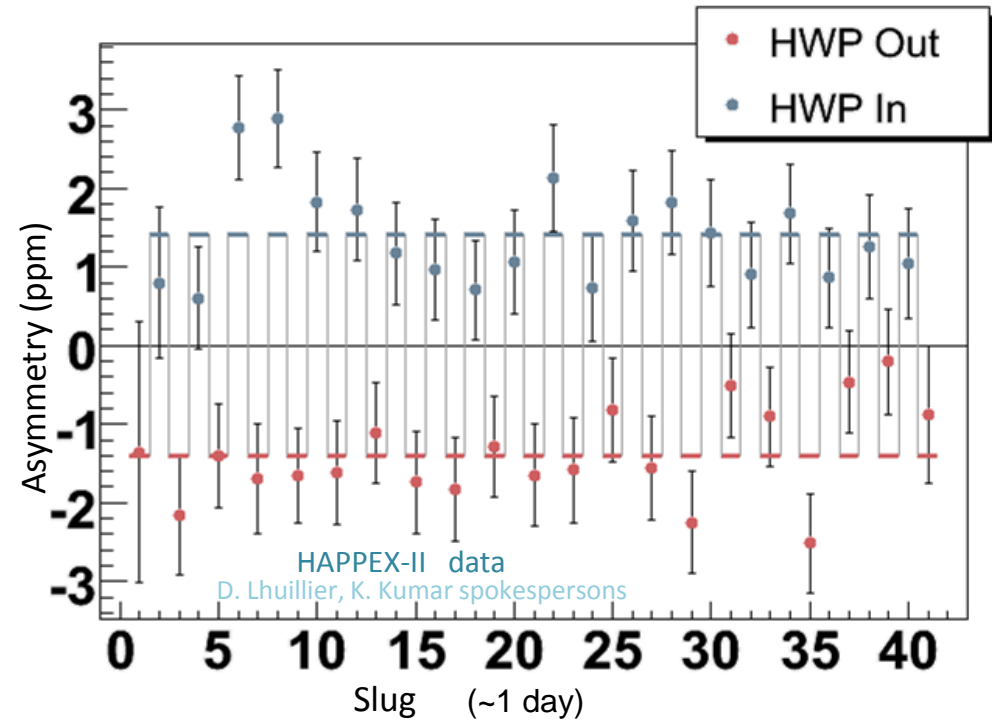
No non-gaussian tails to $\pm 5\sigma$

Small beam-related Systematics -- thanks to Jlab beam quality

- Offline asymmetries nearly identical to online.
- Corrections tiny (here, 3 ppb)
- Errors are statistical only



Parity Violating Asymmetry



$$A_{\text{raw}} = -1.58 \text{ ppm} \pm 0.12 \text{ (stat)} \pm 0.04 \text{ (syst)}$$

(HWP = optical element used to flip beam helicity, helps suppress some systematics)

Parity Quality Beam : Unique Strength of JLab

Helicity – Correlated Position Differences

$$\langle X_R - X_L \rangle \text{ for helicity } L, R$$

$$A_{\text{raw}} = A_{\text{det}} - A_Q + \alpha \Delta_E + \sum \beta_i \Delta x_i$$

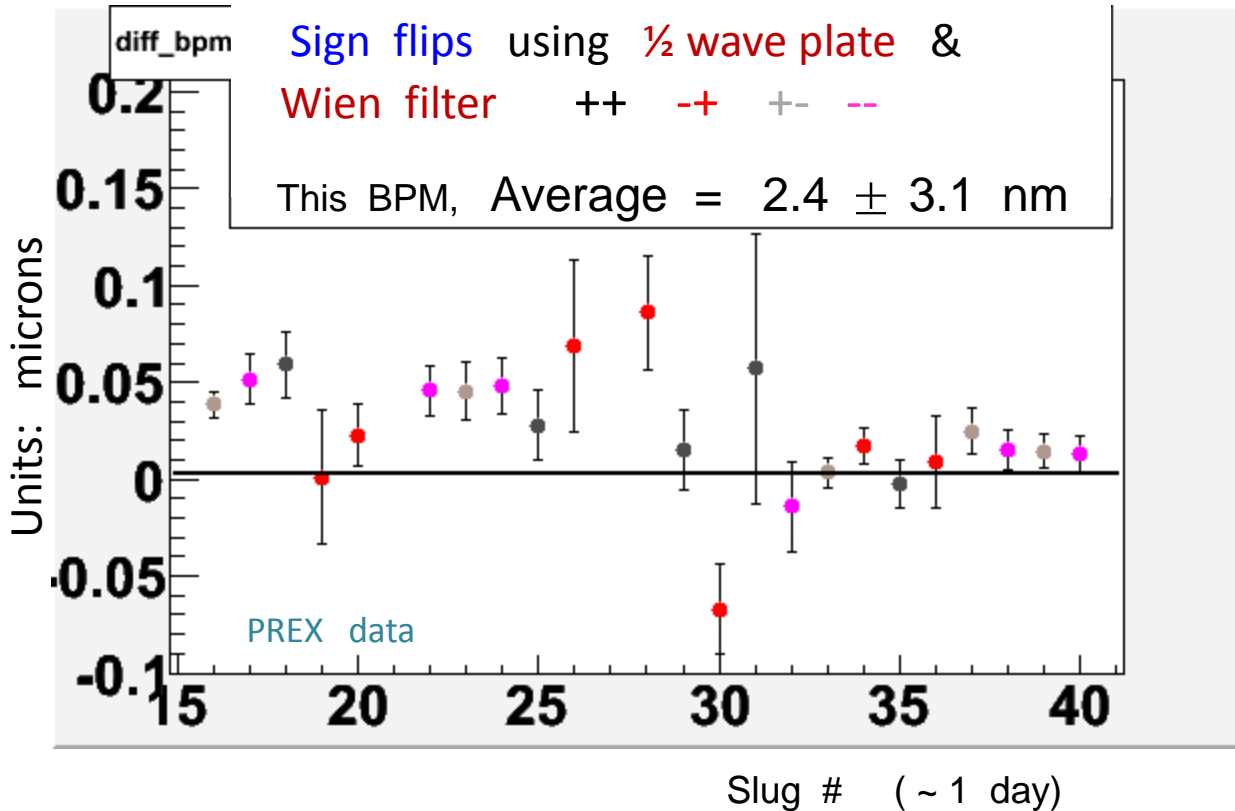
Plotted below

Measured separately

Points: Not sign-corrected. 20-50 nm diffs. with pol. source setup & feedback

Sign flips provide further suppression : Average with signs = what experiment feels

achieved
< 5 nm



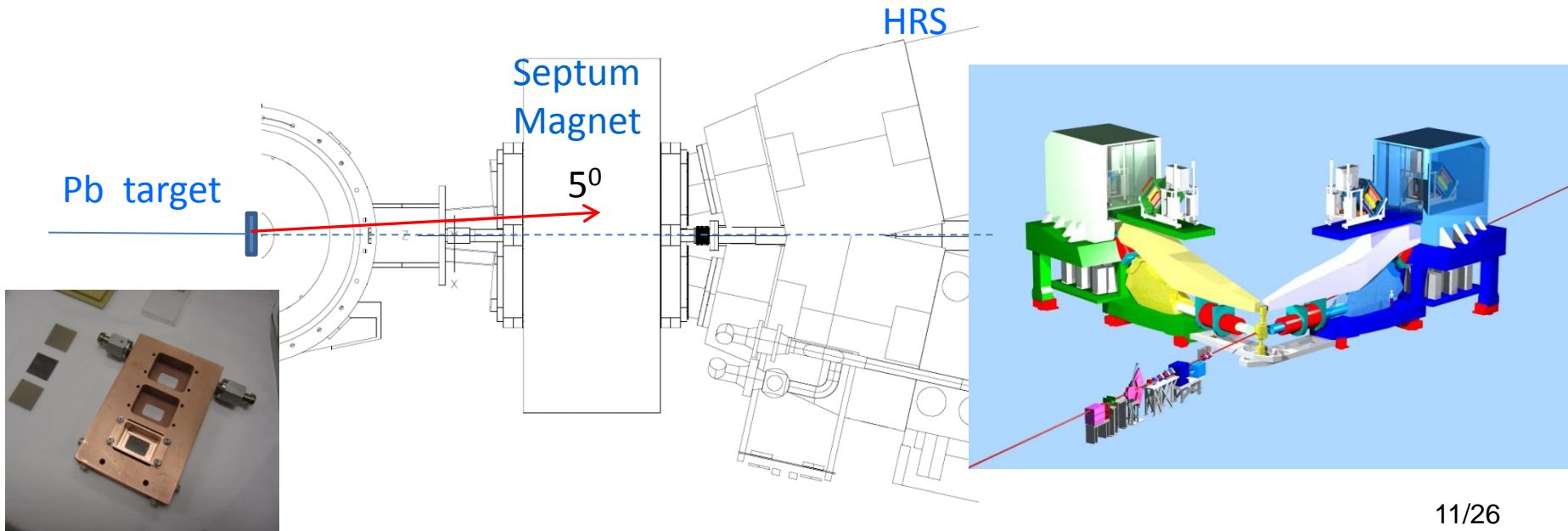
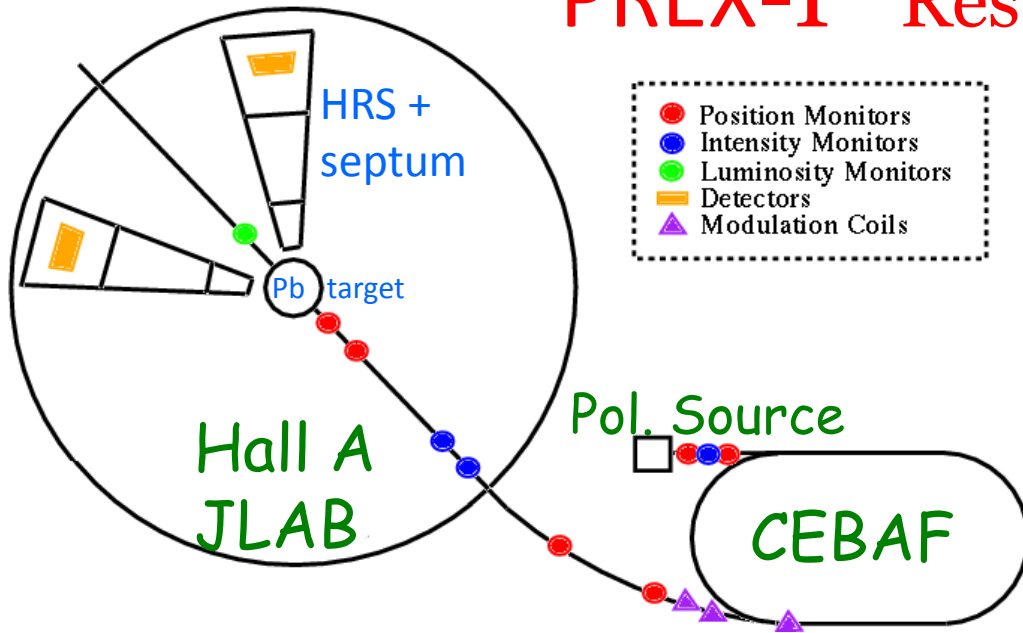
PREX-I Results

PRL 108 (2012) 112502

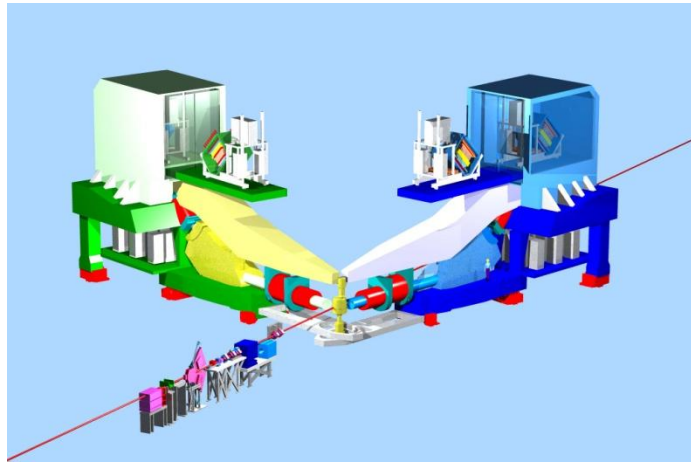
Physics Asymmetry

$$A = 0.656 \text{ ppm} \pm 0.060 \text{ (stat)} \pm 0.014 \text{ (syst)}$$

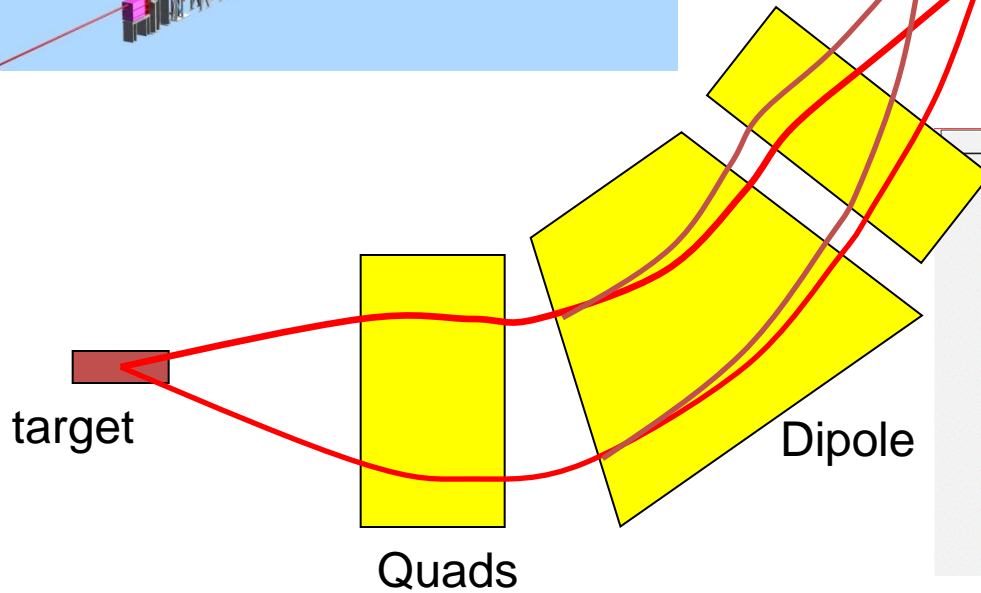
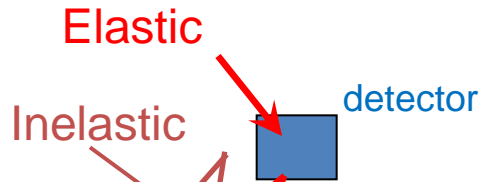
- Statistics limited (9%)
- Systematic error goal achieved! (2%)



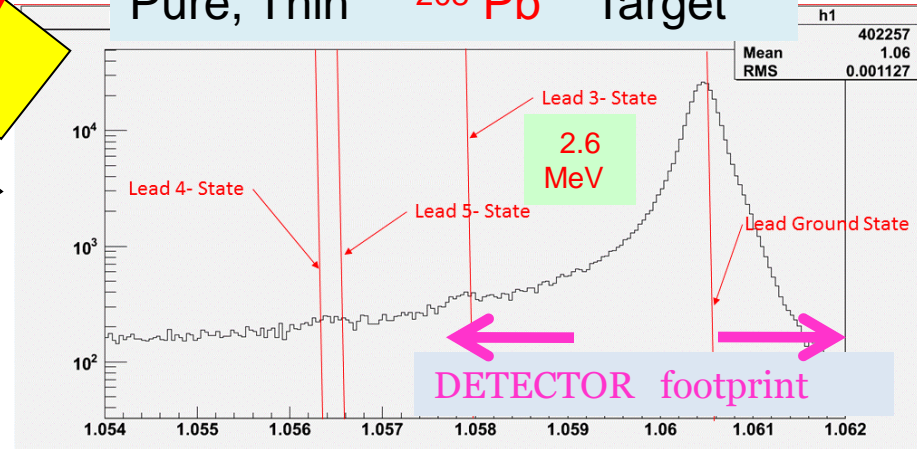
Hall A High Resolution Spectrometers



- Resolve Elastic Scattering
- Discriminate Excited States



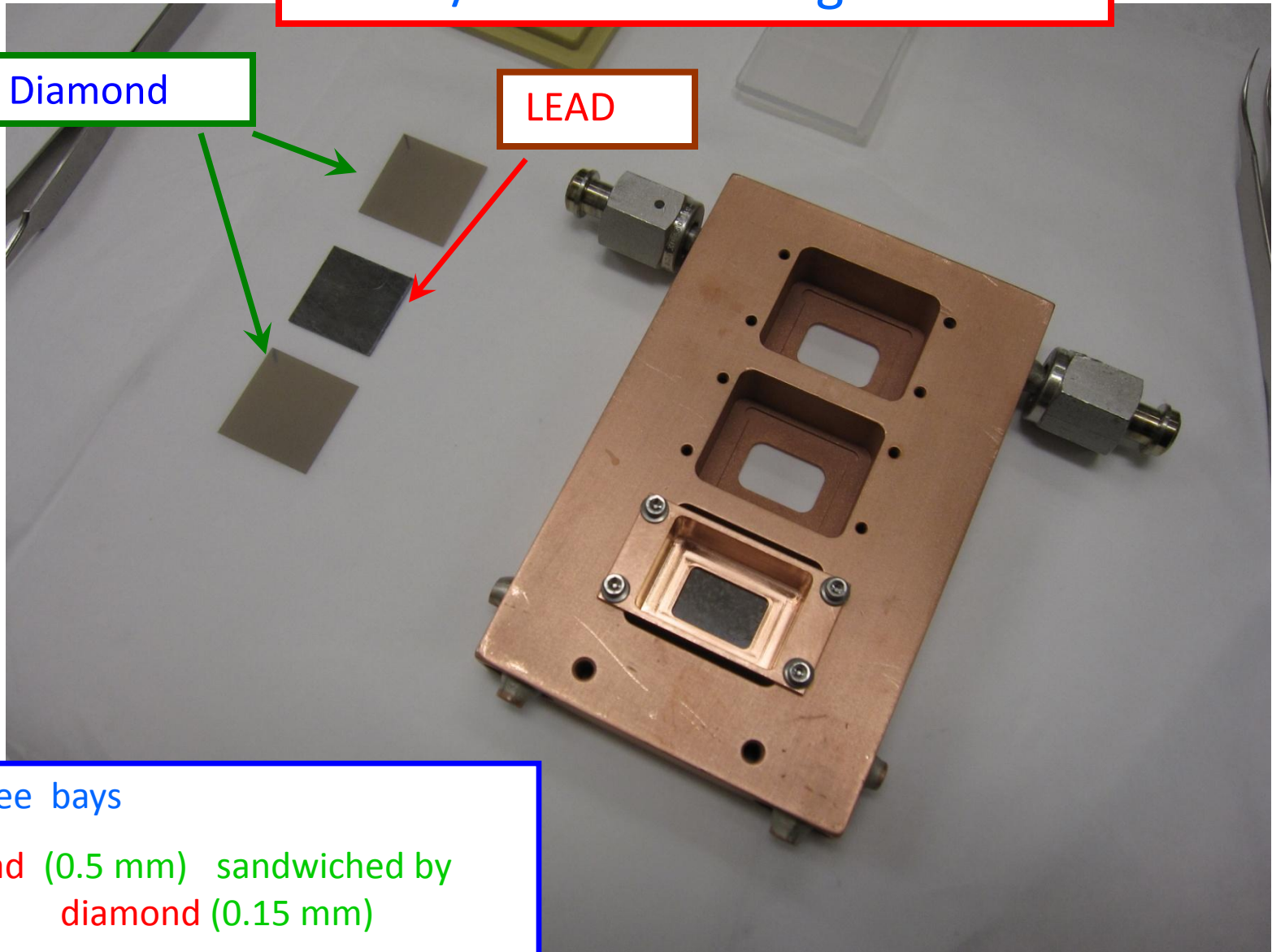
Pure, Thin ²⁰⁸Pb Target



Lead / Diamond Target

Diamond

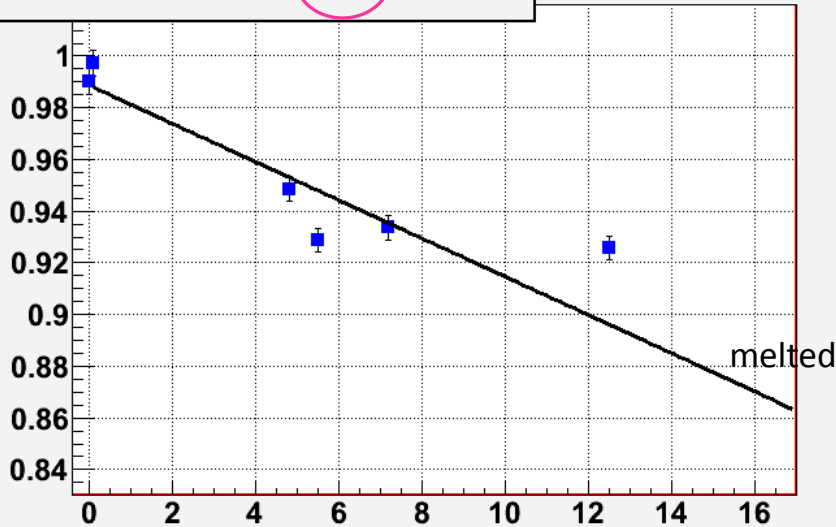
LEAD



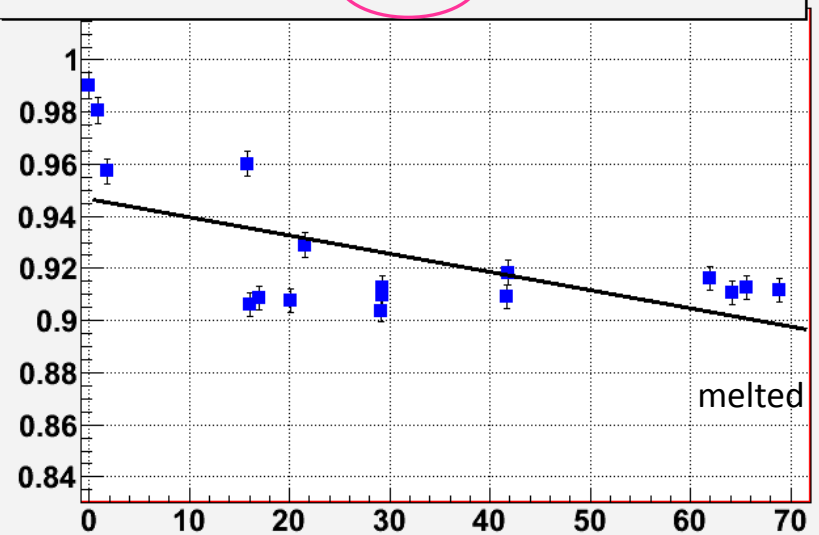
- Three bays
- Lead (0.5 mm) sandwiched by diamond (0.15 mm)
- Liquid He cooling (30 Watts)

Performance of Lead / Diamond Targets

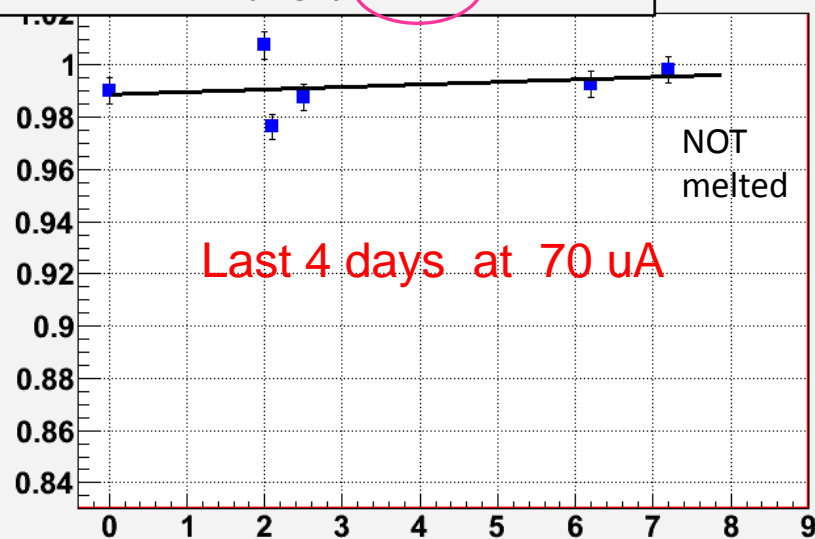
Rate vs Time (days) **THIN** Diamond



Rate vs Time (days) **MEDIUM**-thickness Diamond



Rate vs Time (days) **THICK** Diamond



Targets with thin diamond backing (4.5 % background) degraded fastest.

Thick diamond (8%) ran well and did not melt at 70 uA.

→ Solution: Run with 10 targets.

PREX-I Result

Systematic Errors

Error Source	Absolute (ppm)	Relative (%)
Polarization (1)	0.0083	1.3
Beam Asymmetries (2)	0.0072	1.1
Detector Linearity	0.0076	1.2
BCM Linearity	0.0010	0.2
Rescattering	0.0001	0
Transverse Polarization	0.0012	0.2
Q ² (1)	0.0028	0.4
Target Thickness	0.0005	0.1
¹² C Asymmetry (2)	0.0025	0.4
Inelastic States	0	0
TOTAL	0.0140	2.1

(1) Normalization Correction applied

(2) Nonzero correction (the rest assumed zero)

Physics Asymmetry

$$A = 0.656 \text{ ppm} \\ \pm 0.060 \text{ (stat)} \pm 0.014 \text{ (syst)}$$

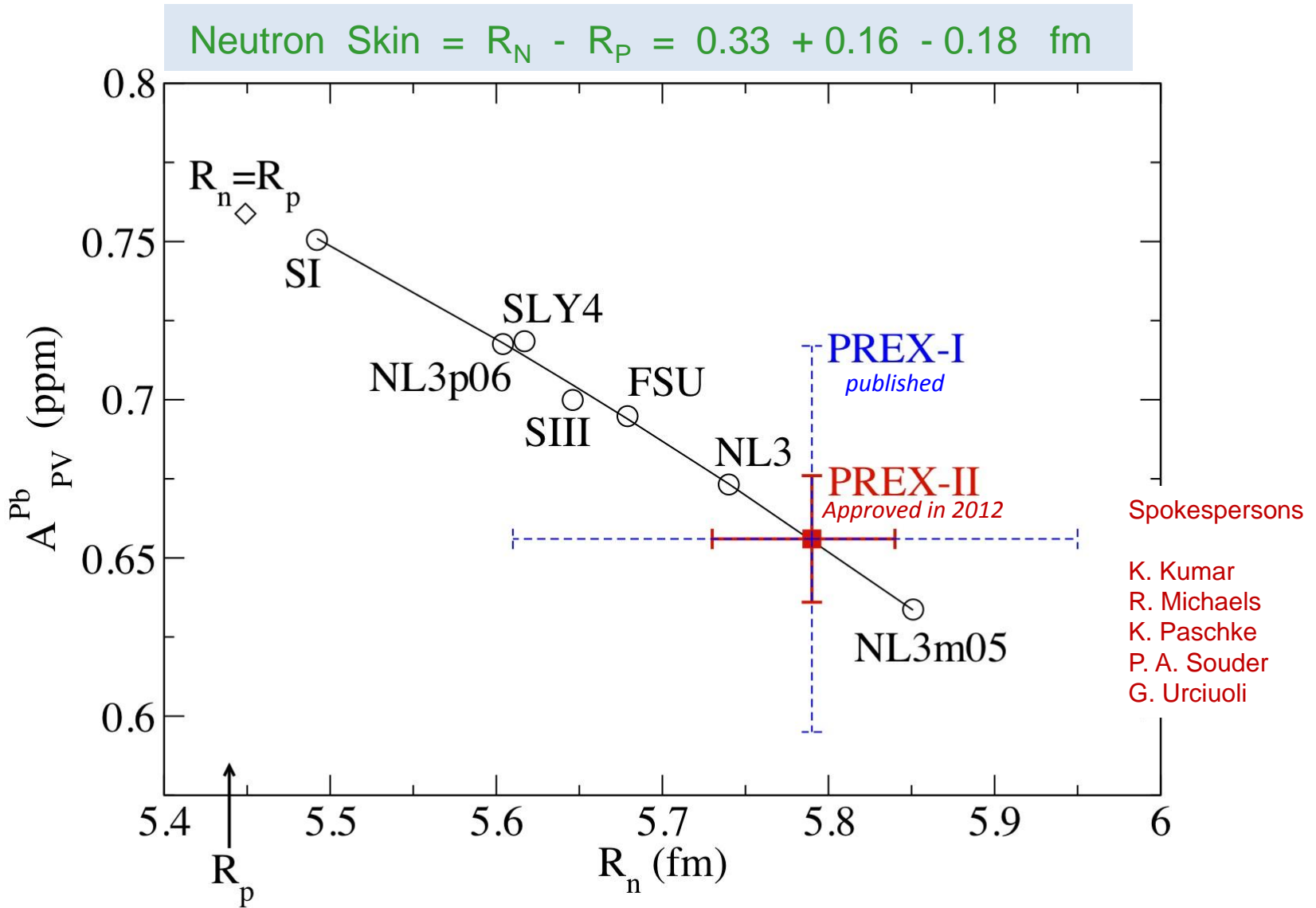
→ Statistics limited (9%)

→ Systematic error goal achieved! (2%)

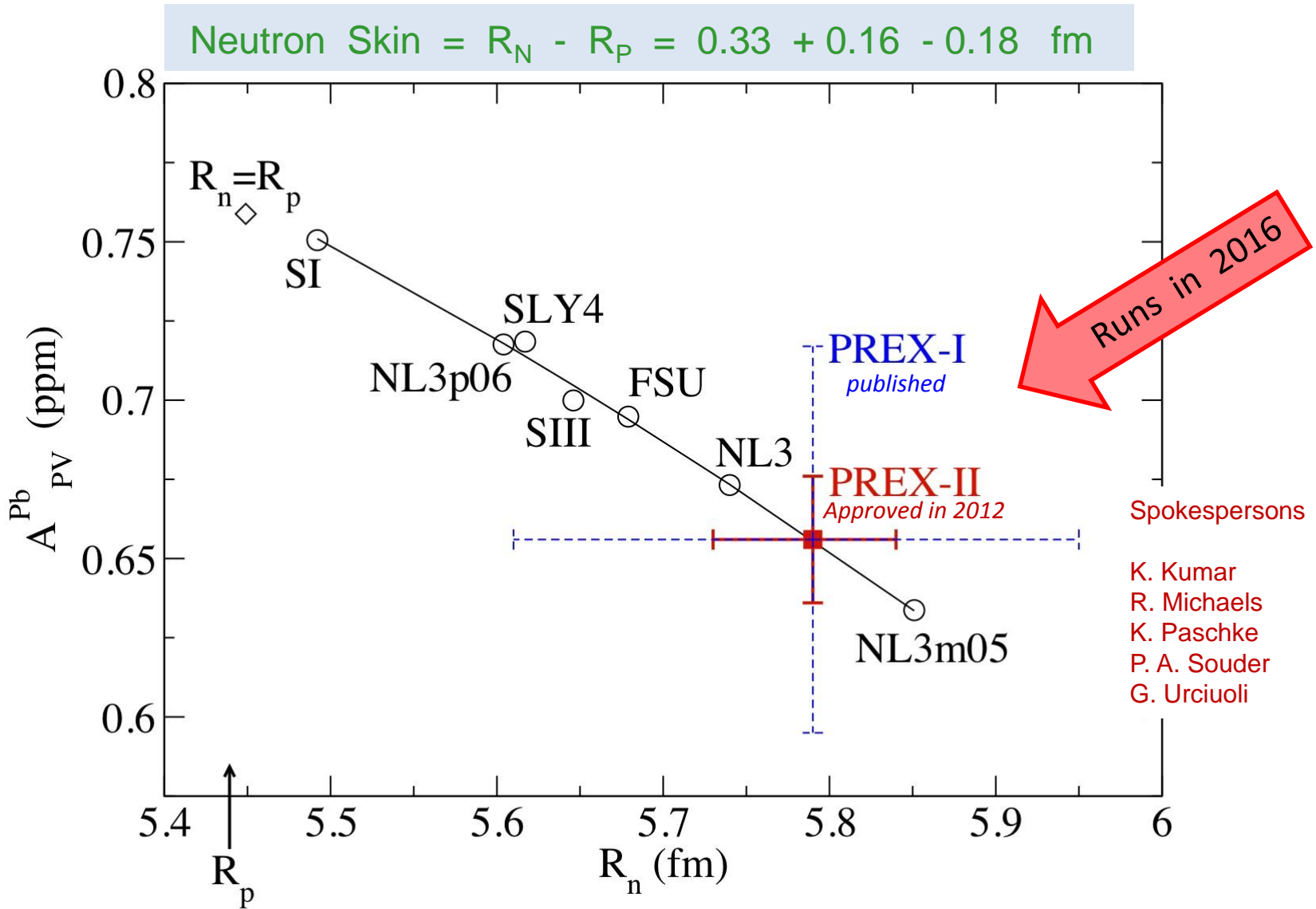
PRL 108, 112502 (2012)

PRC 85, 032501(R) (2012)

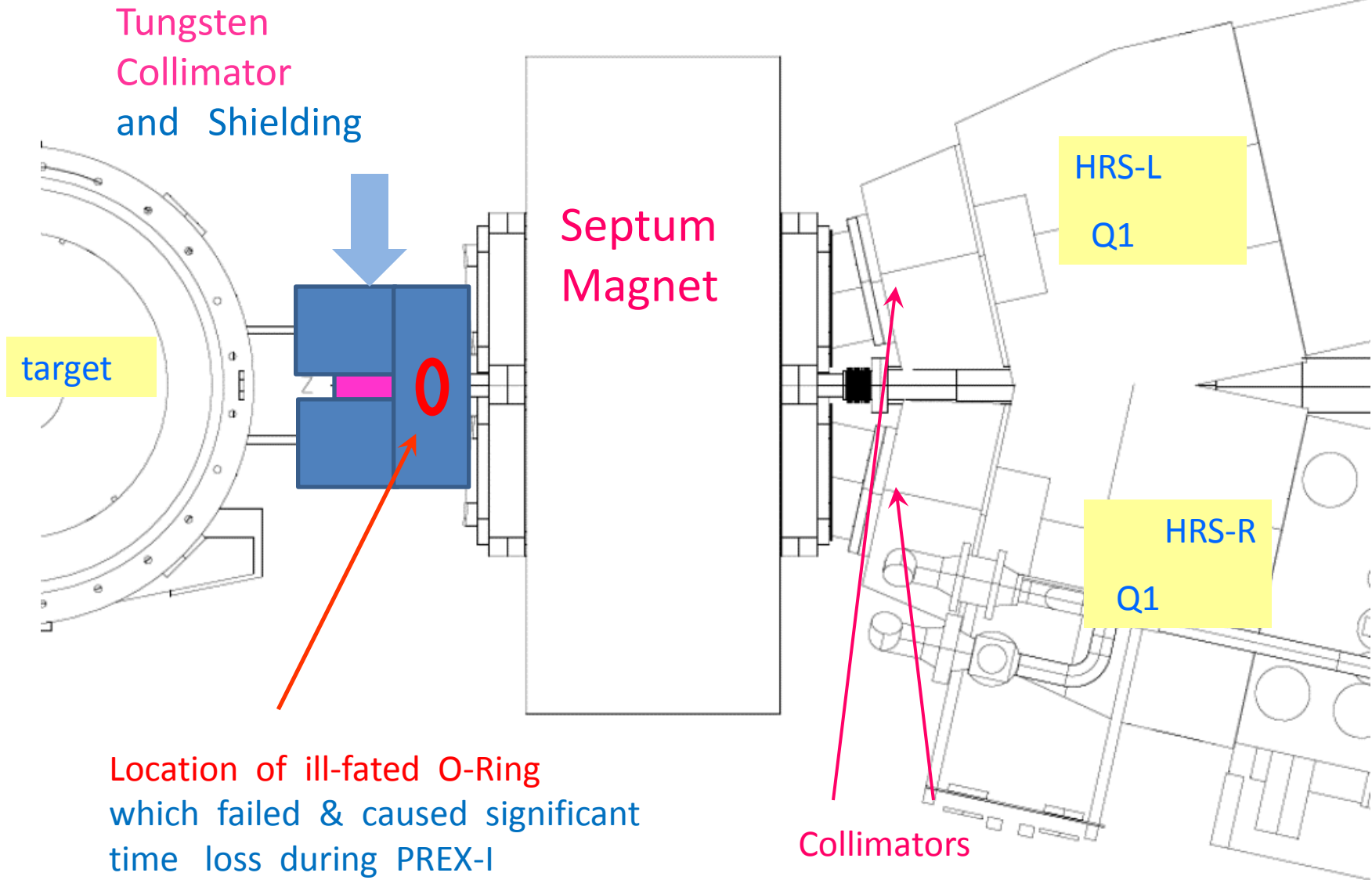
PREX: Asymmetry leads to R_N Establishing a neutron skin at ~95 % CL



PREX: Asymmetry leads to R_N Establishing a neutron skin at ~95 % CL



Improvements for PREX-II



Location of ill-fated O-Ring which failed & caused significant time loss during PREX-I

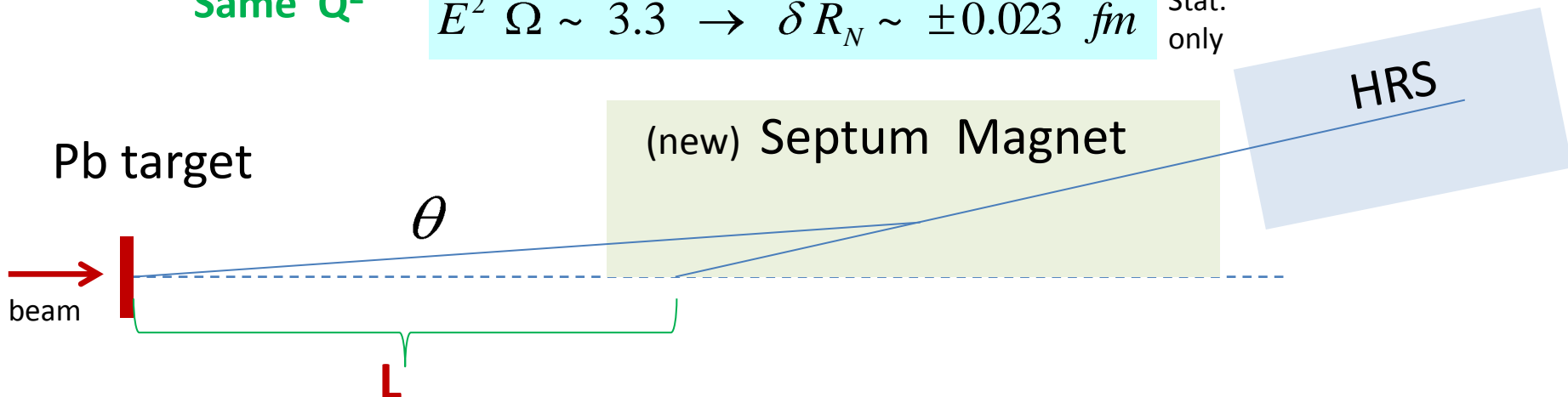
→ PREX-II to use all-metal seals

Chuck asks : “ What about a **PREX-III** ? ”

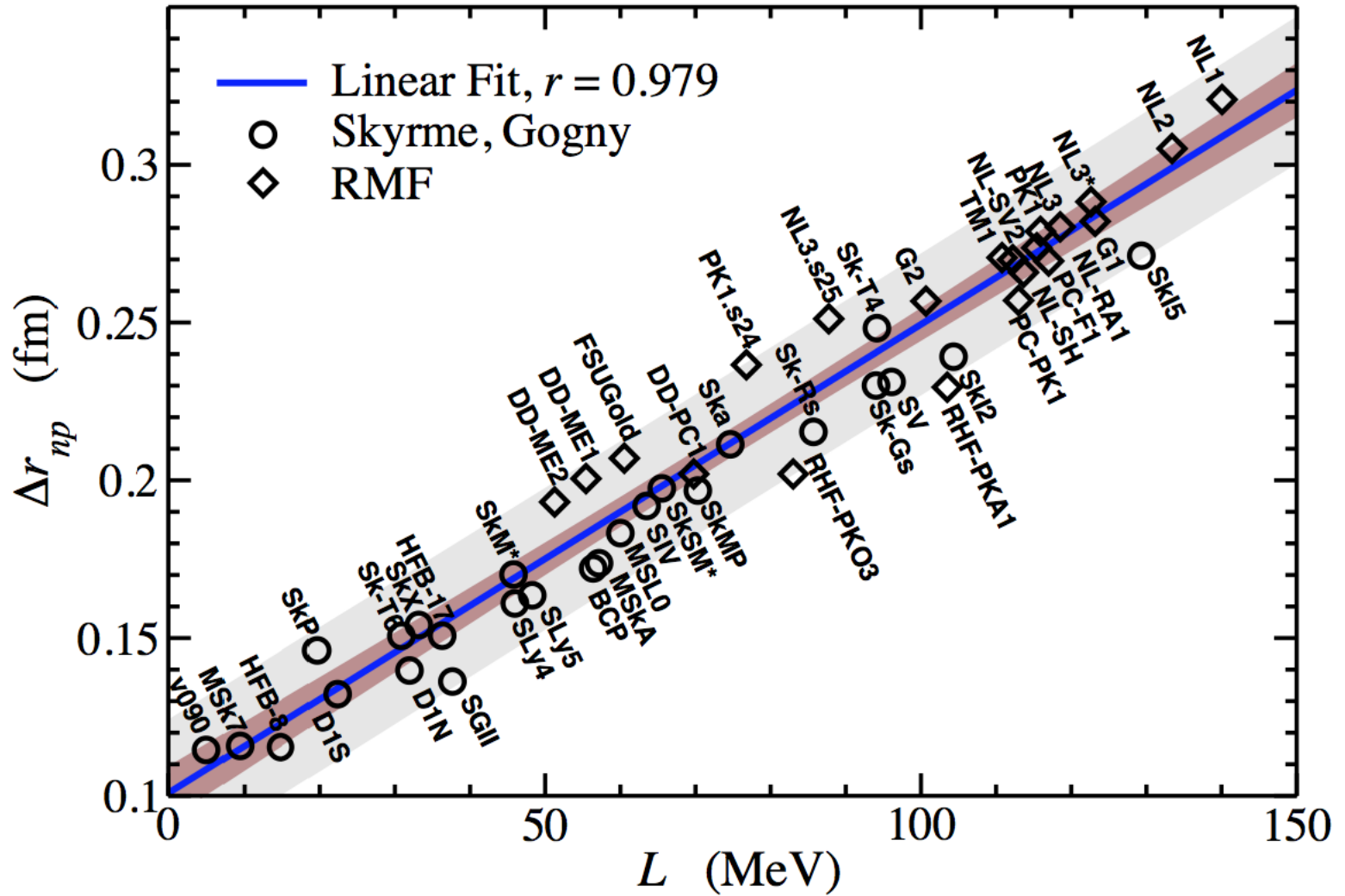
2.2 GeV - 1-pass JLab beam energy (2 month run)

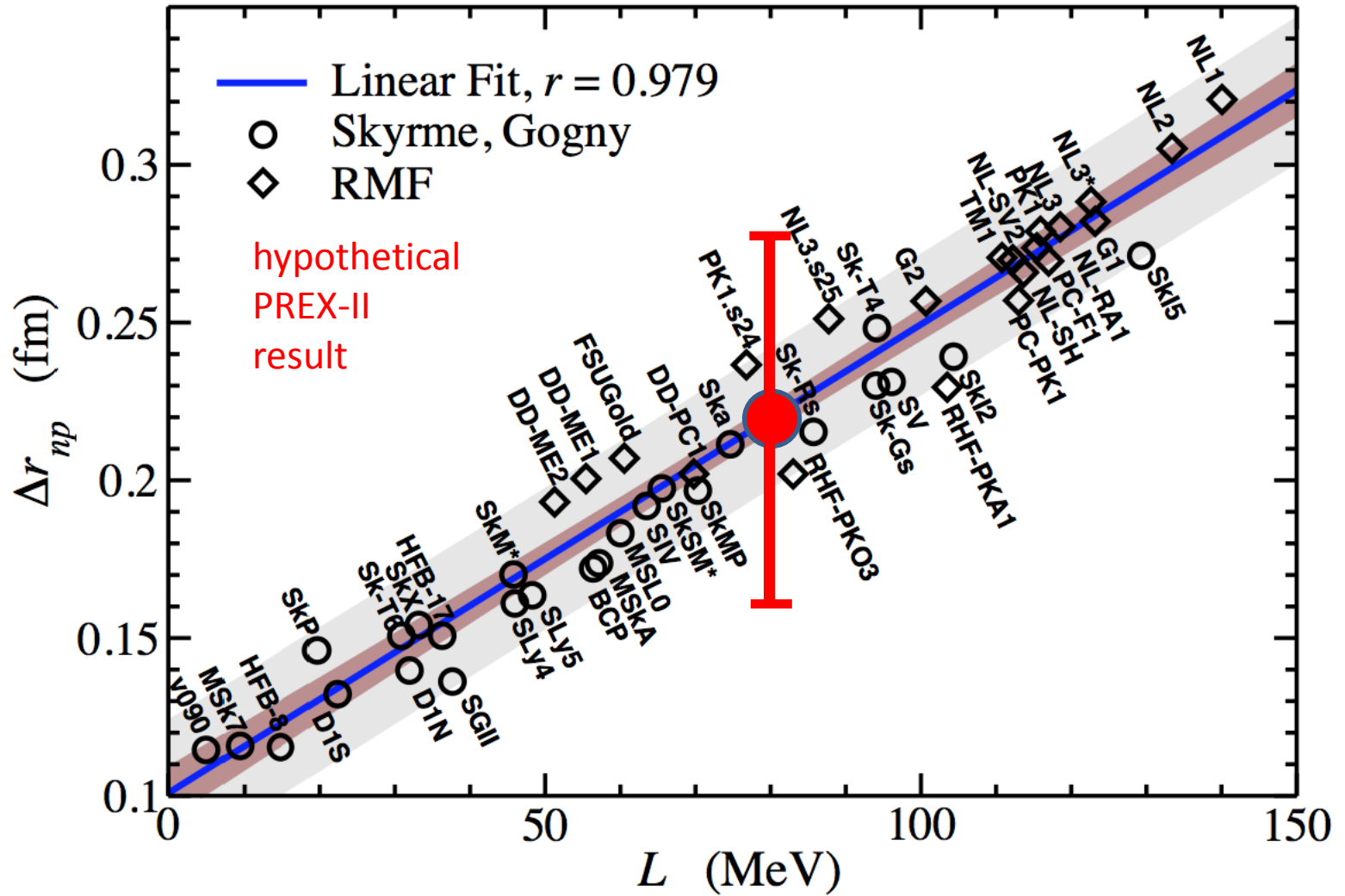
Same Q^2

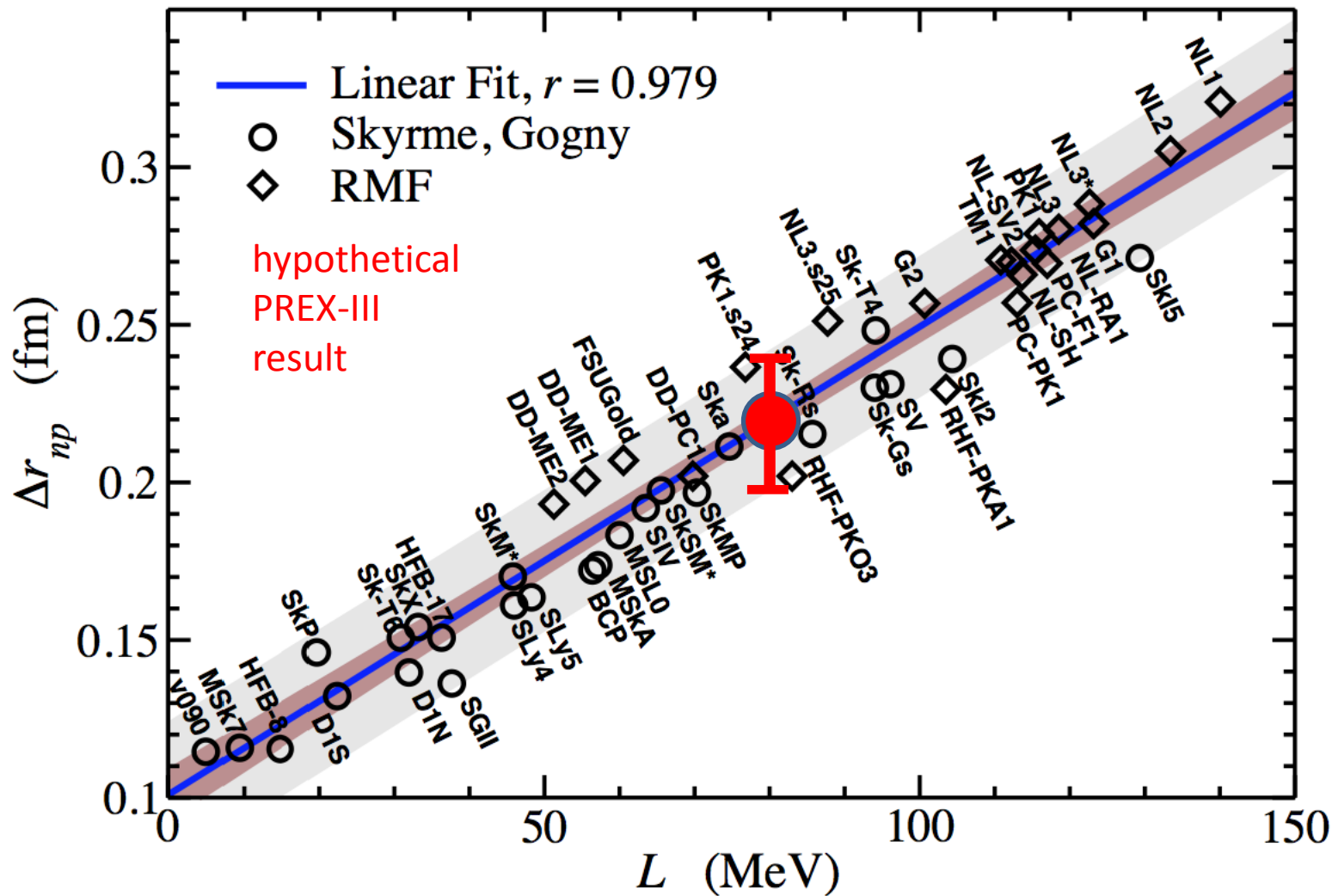
$$E^2 \Omega \sim 3.3 \rightarrow \delta R_N \sim \pm 0.023 \text{ fm} \quad \text{Stat. only}$$



quantity	PREX-II or CREX	new PREX-III	comment
Energy, angle	1.05 GeV, 5°	2.2 GeV, 2.4°	same Q^2
L	1.36 m (CREX)	2 m	Re-engineer beamline
Septum, Acceptance Ω	1 m x 11.3 kG 2.9 msr (CREX)	1.7 m x 19 kG $\Omega \sim 2 \text{ msr}$	new magnet, more windings
Q^2 Error	0.4 %	1.7 %	Problem: need $\ll 1\%$
Polarimetry	1 %	$< 1\%$	Easier at higher E



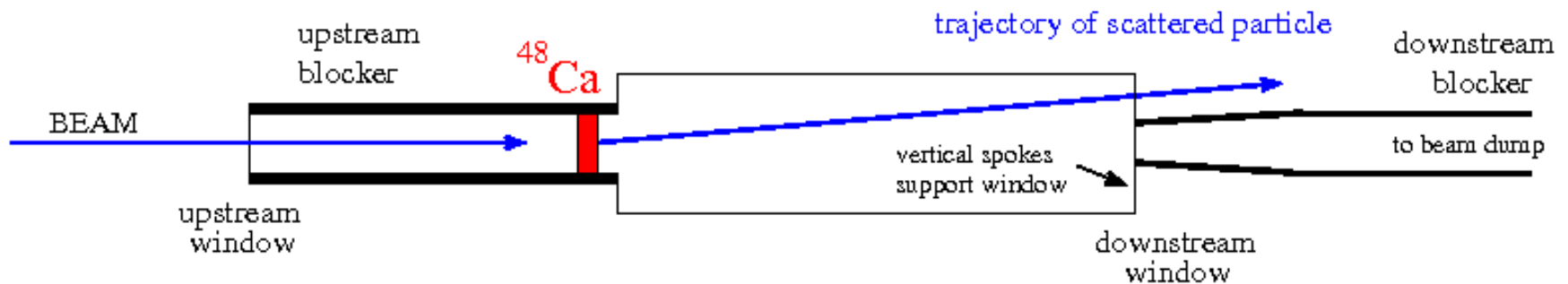




CREX Target



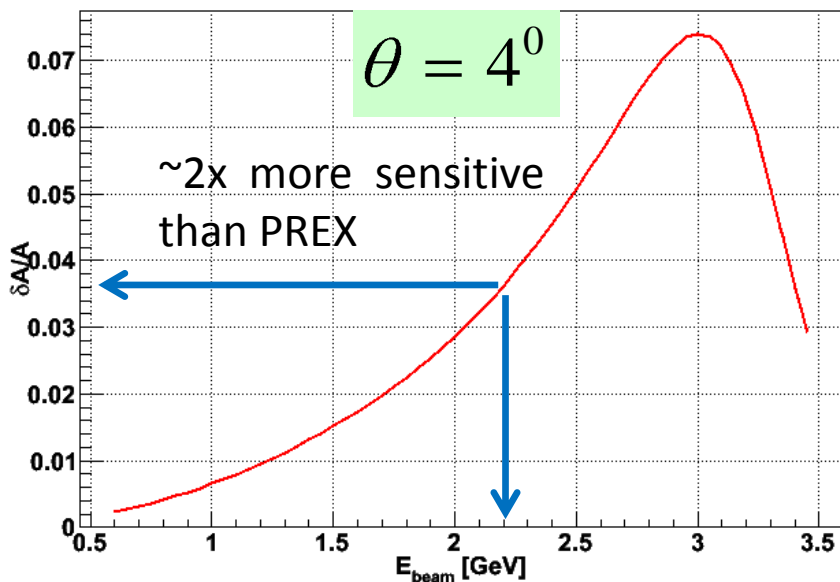
- Isotopically pure ^{48}Ca
- Vacuum seal to trap atoms if beam destroys target
- Higher thermal conductivity and melting point than lead :
should take 100 μA .
- Similar to target used in E08014 (at 40 μA)



CREX: Optimization of Kinematics

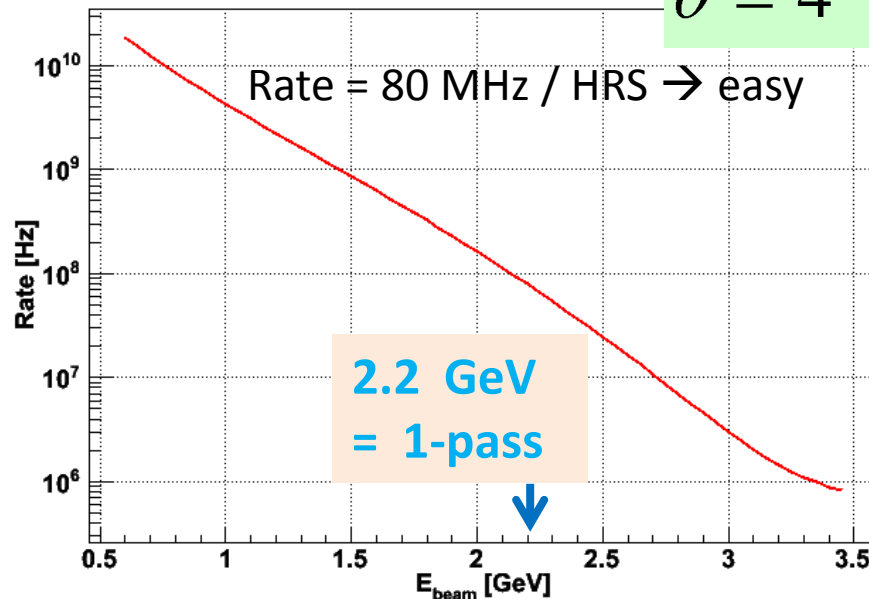
Sensitivity of Asymmetry (A) to change in R_N

dA/A for ^{48}Ca 1% change in R vs. Energy, $\theta = 4^\circ$

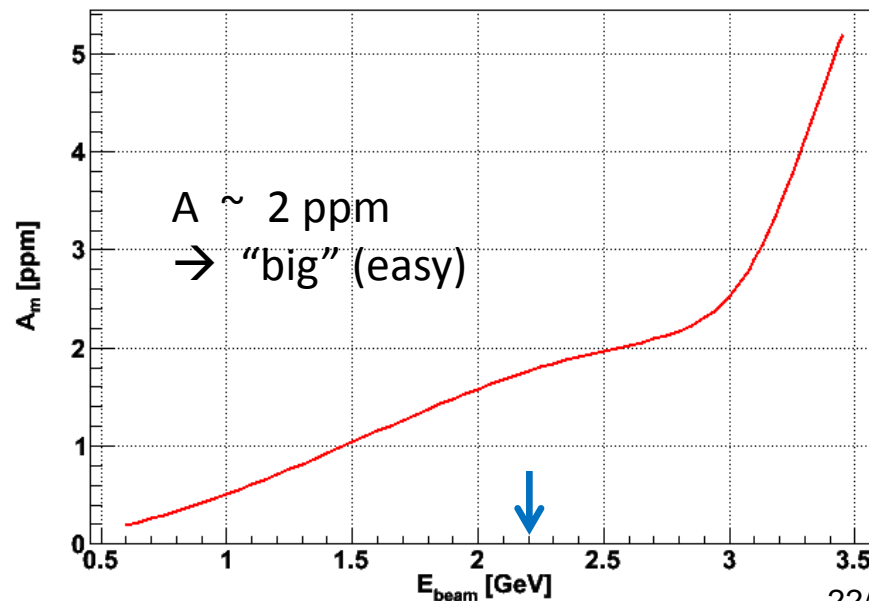


Rate vs. Energy, ^{48}Ca , $\theta = 4^\circ$, $100\mu\text{A}$, 1 HRS

$\theta = 4^\circ$



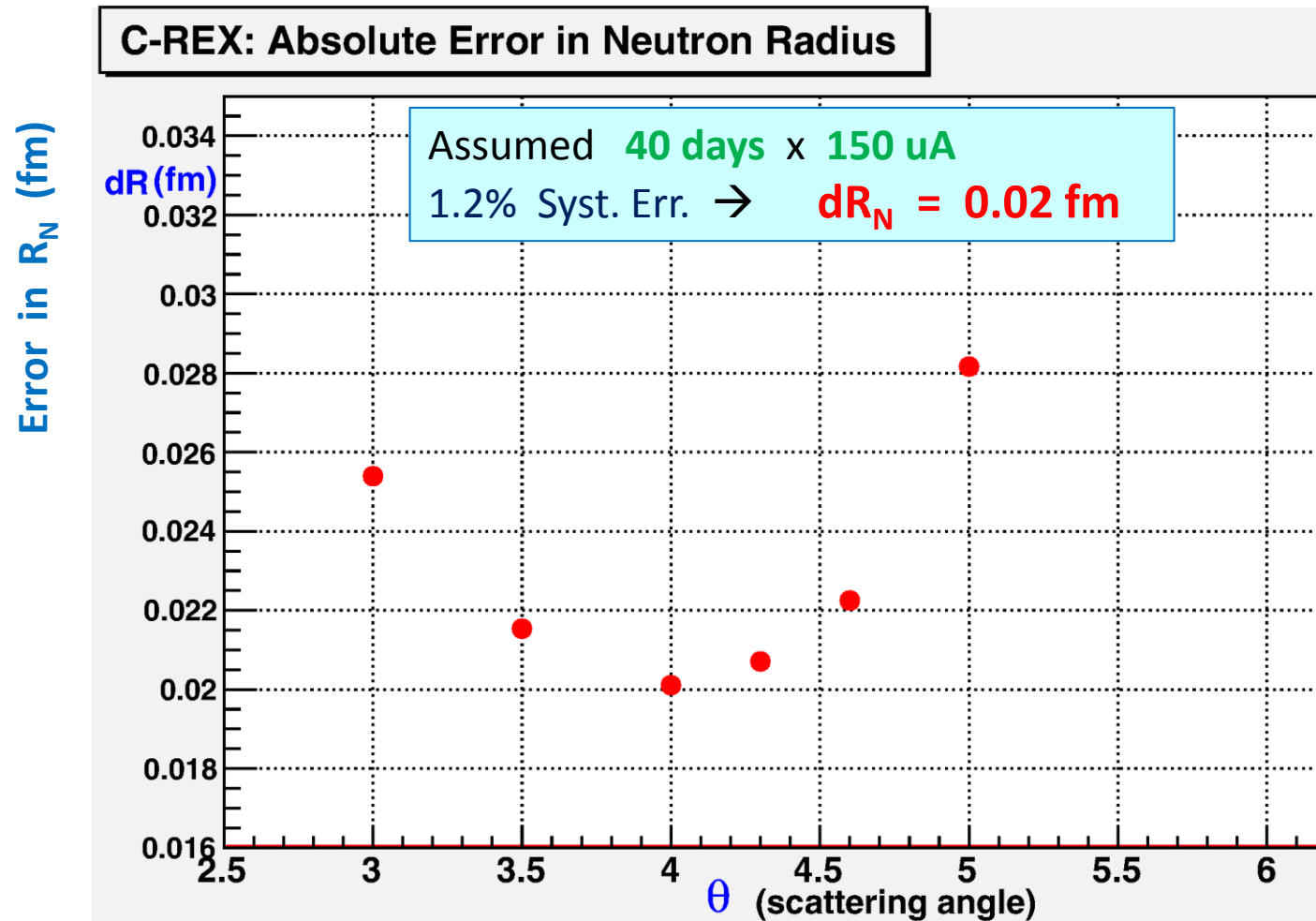
Measured asymmetry vs. Energy, ^{48}Ca , $\theta = 4^\circ$



CREX -- choice of kinematics

$$E = 2.2 \text{ GeV}, \theta = 4^\circ$$

Optimization of FOM equivalent to minimum error in R_N

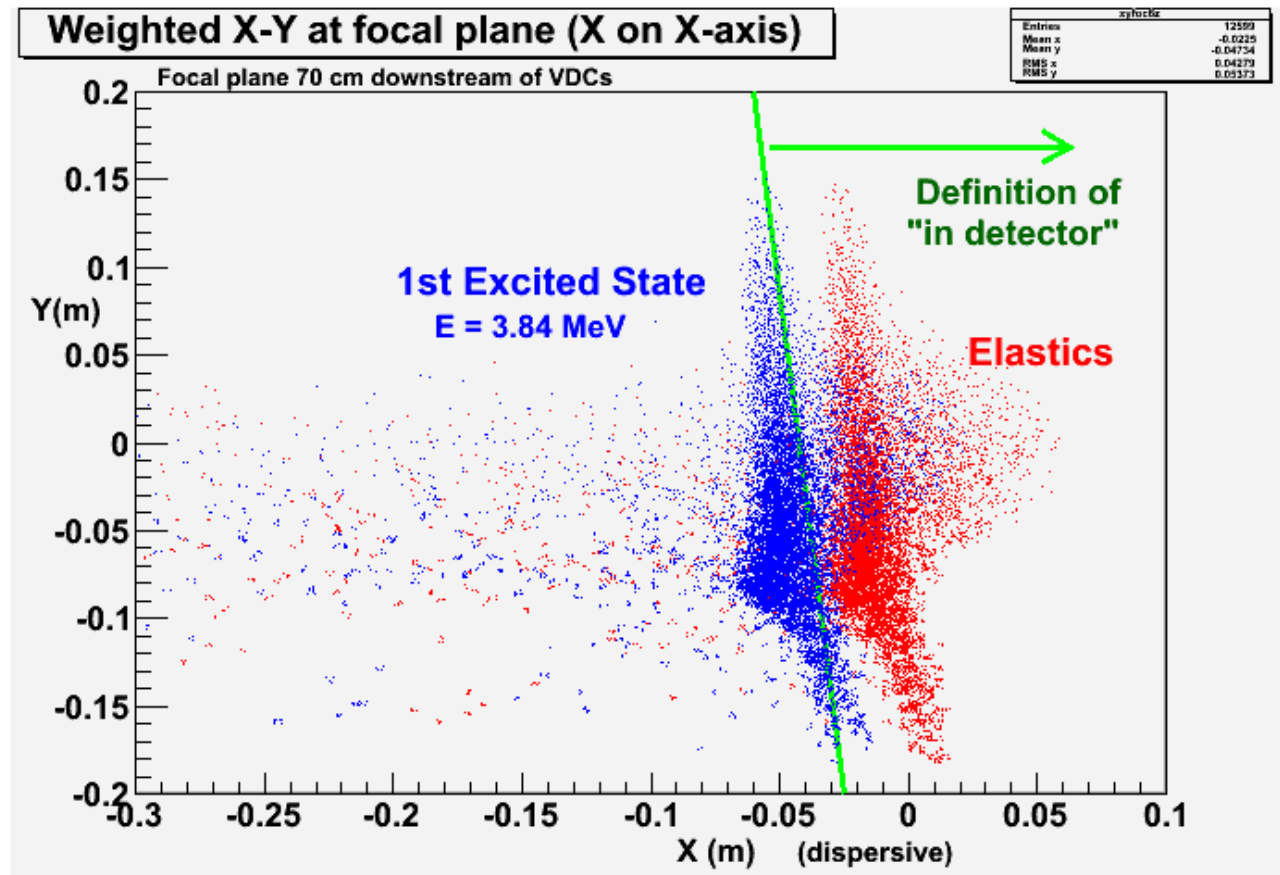


Inelastic Contributions -- background Correction $dA = f * A_{\text{state}}$

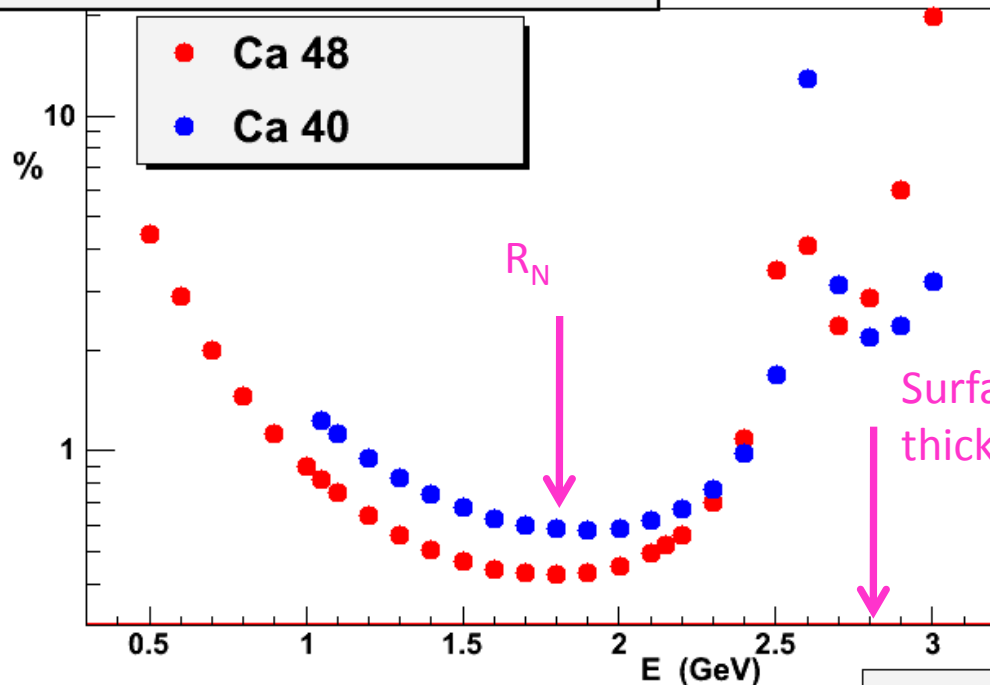
For ^{48}Ca : 1st state $f = 0.19\%$, 2nd state $f = 0.18\%$, next 8 states $f_{\text{tot}} = 0.03\%$

Need estimates for A_{state} and further optimization of HRS tune

Simulated Events in HRS



Percent Error in R_N vs Energy (Calcium Isotopes)



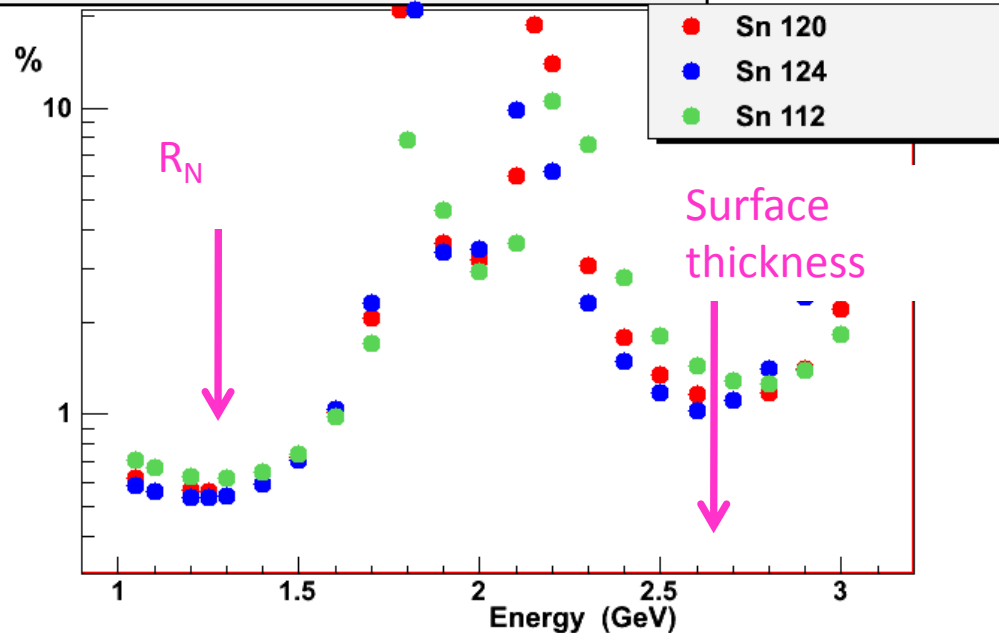
After PREX / CREX

Other Nuclei ?

and Shape Dependence ?

each point 30 days

Percent Error in R_N vs Energy (Tin Isotopes)

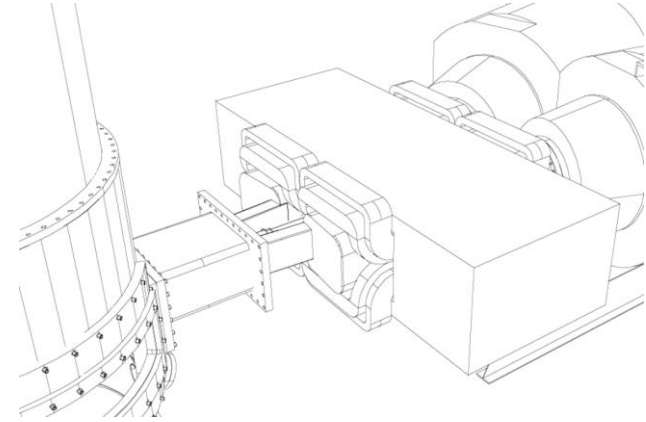


Parity Violating Electron Scattering
Measurements of Neutron Densities

Shufang Ban, C.J. Horowitz, R. Michaels

J. Phys. G39 014104 2012

Possible Future PREX - type Program ?



Each point 30 days, except PREX-III 60 days stat. error only

Nucleus	E (GeV)	dR_N / R_N	comment
^{208}Pb	1	1 % (0.06 fm)	PREX-II (approved by JLab PAC, A rating)
^{48}Ca	2.2 (1-pass)	0.02 fm	CREX natural 12 GeV exp't (approved, A-)
^{208}Pb	2.2 (1-pass)	0.02 fm	PREX-III discussed today
^{48}Ca	2.6	2 %	surface thickness
^{40}Ca	2.2 (1-pass)	0.6 %	basic check of theory
tin isotope	1.8	0.6 %	apply to heavy ion ?
tin isotope	2.6	1.6 %	surface thickness

Not yet proposed

PREX / CREX : Summary

- Fundamental Nuclear Physics with many applications
- PREX-I achieved a 9% stat. error in Asymmetry (original goal : 3 %)
- Systematic Error Goals Achieved !!
- PREX-II approved to reach original goals (runs in 2016)
- CREX approved (unscheduled)