

3rd International Symposium on Nuclear Symmetry Energy (NuSYM13)
NSCL/FRIB, East Lansing, Michigan, U.S.A.
22-26 July, 2013

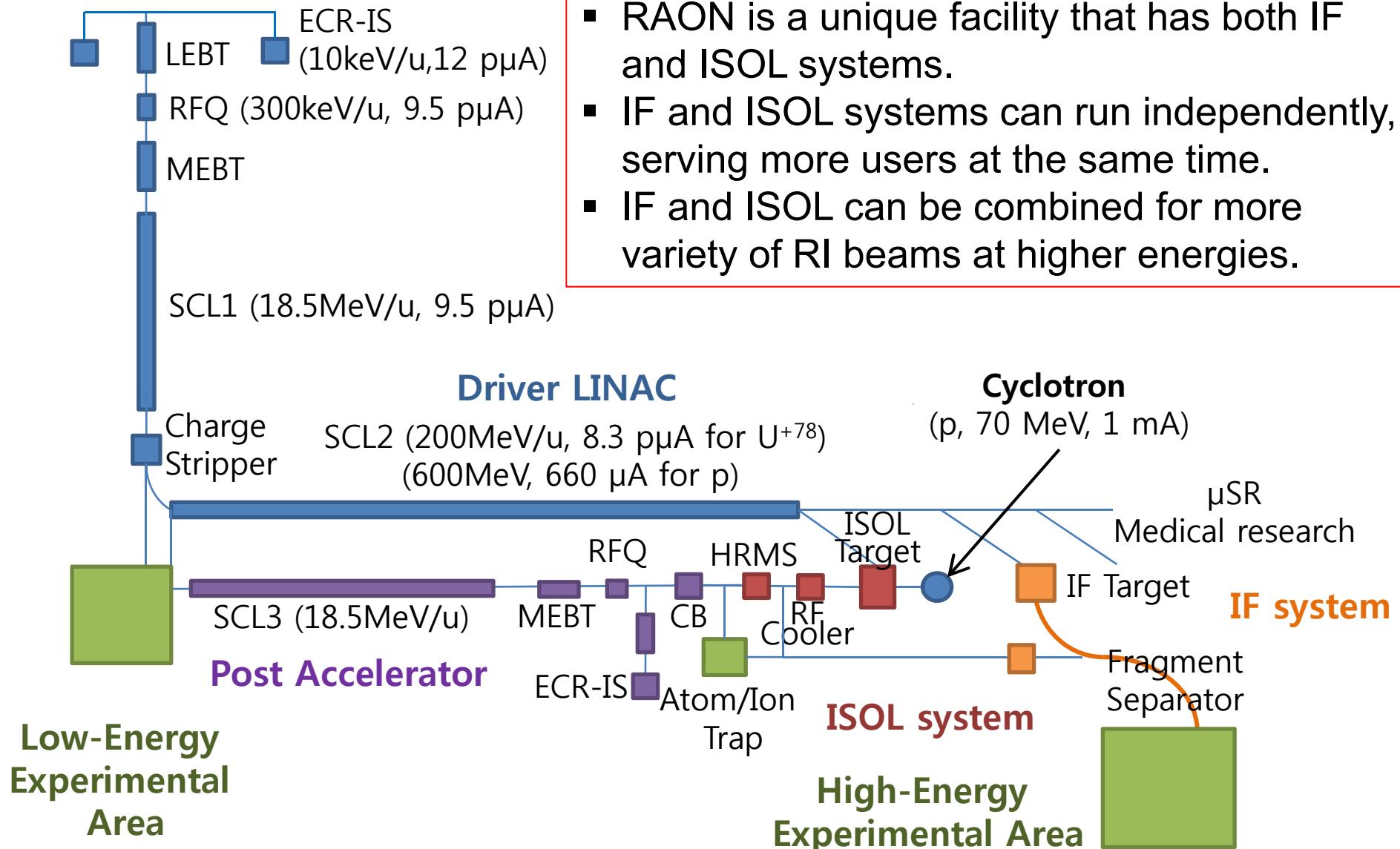
Plan for the Nuclear Symmetry Energy Measurements at RAON, LAMPS Facility

Byungsik Hong
(Korea University)

Outline

1. Introduction to RISP
 - Rare Isotope Science Project in Korea
 - RAON rare isotope beam facility
2. KOBRA
 - Broad acceptance recoil spectrometer at low-energy experimental area
3. LAMPS
 - Large-acceptance multipurpose spectrometer
 - A dedicated system for symmetry energy at high-energy experimental area
 - In addition, low-energy LAMPS will be also built and used at KOBRA for symmetry energy
4. Summary

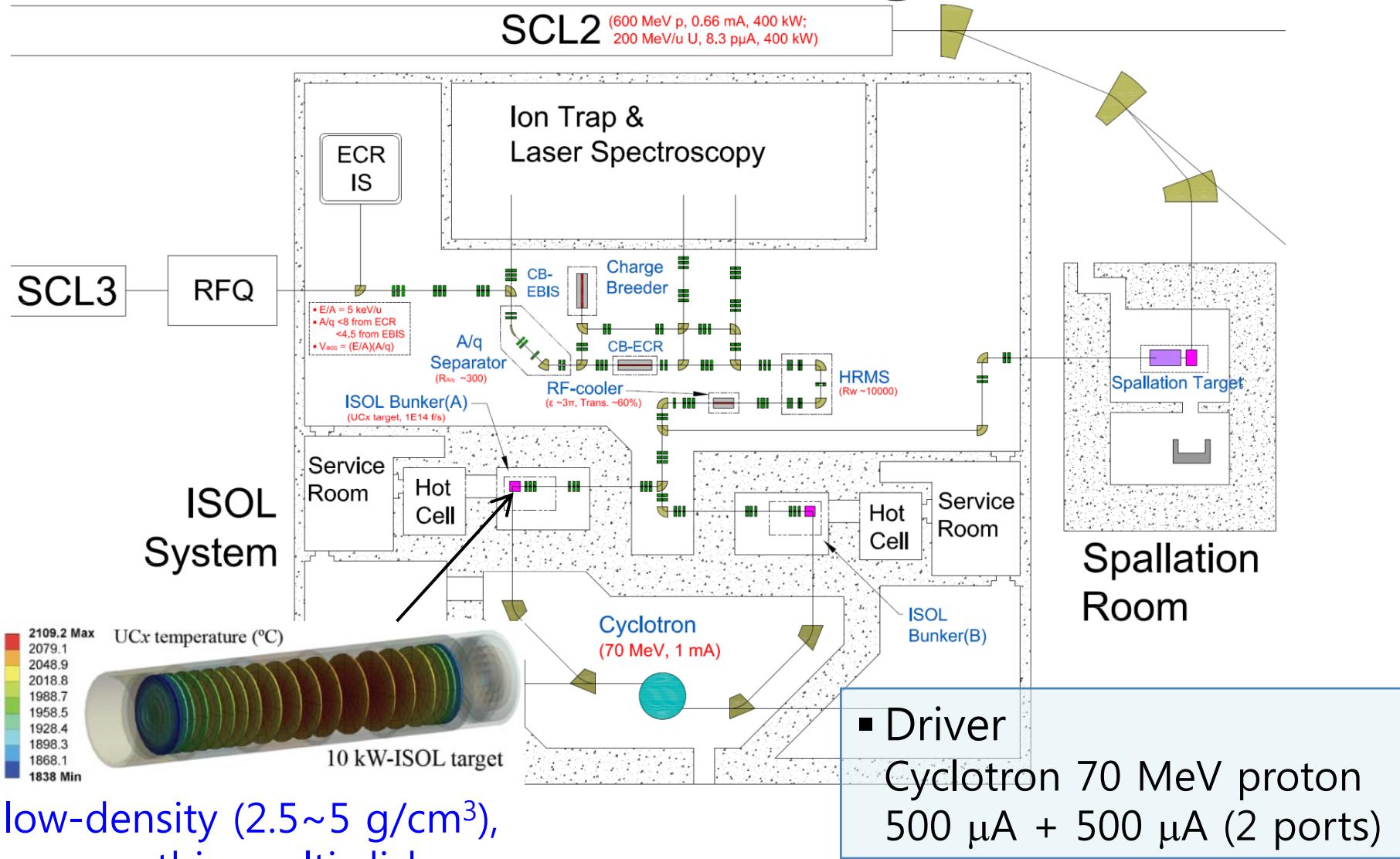
RAON



Beam Parameters of RAON

	Driver Linac				Post Acc.	Cyclotron
Particle	H ⁺	O ⁺⁸	Xe ⁺⁵⁴	U ⁺⁷⁹	RI beam	proton
Beam energy (MeV/u)	600	320	251	200	18.5	70
Beam current (pμA)	660	78	11	8.3	-	1000
Power on target (kW)	>400	400	400	400	-	70

ISOL Design



RI Yield Estimation

$$Y_{\text{ISOL}} = \Phi_P \sigma_f N_{\text{target}} \varepsilon_{\text{release}} \varepsilon_{\text{ionization}} \varepsilon_{\text{cooler}} \varepsilon_{\text{transport}} \varepsilon_{\text{charge-breeding}} \varepsilon_{\text{acceleration}}.$$

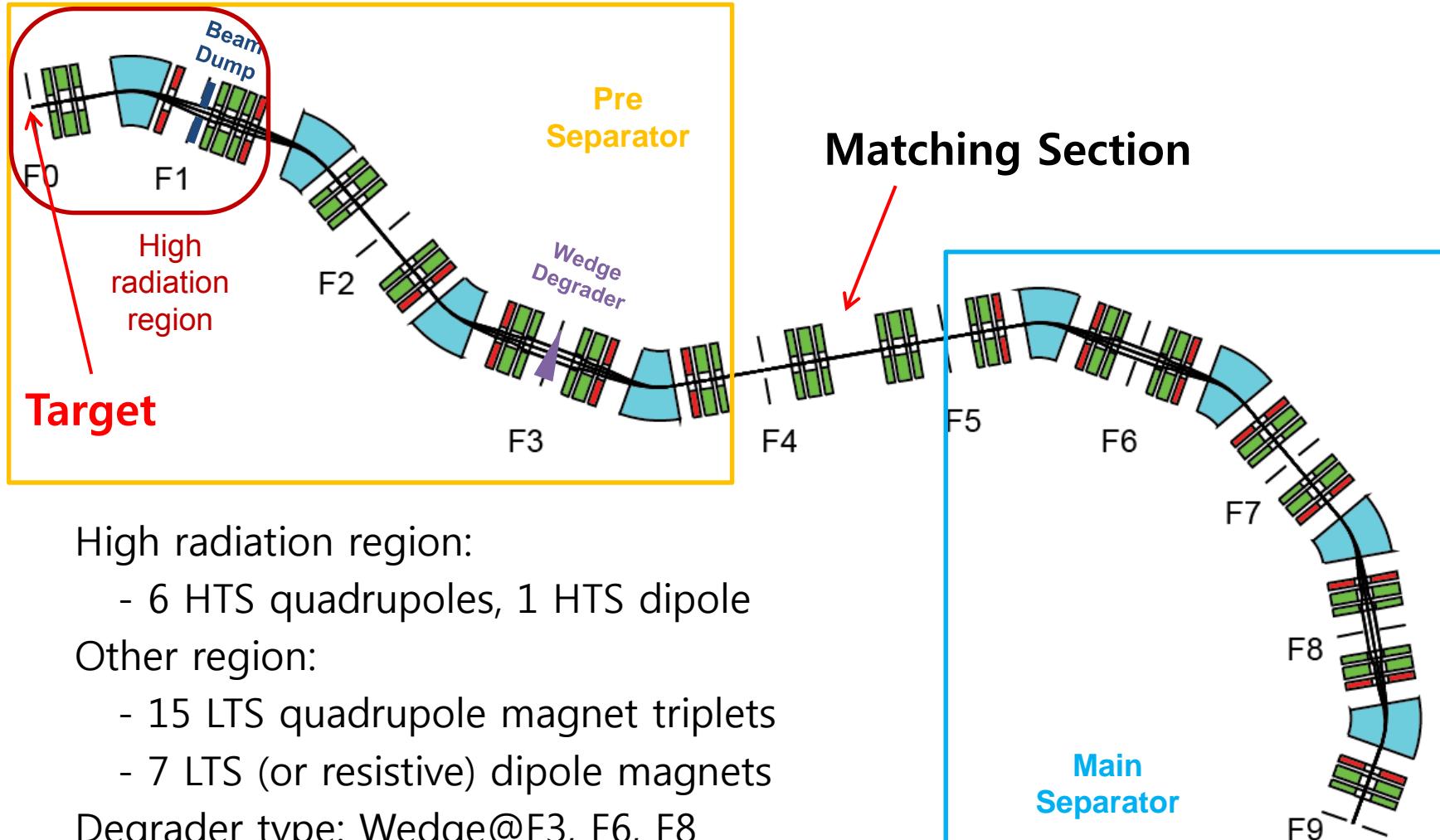
^{132}Sn intensities for 10 kW and 35 kW ISOL targets

- BERTINI-ORNL model for 2.5 g/cm³ target
- Assuming overall efficiency of ~0.5 %

	10 kW	35 kW
Deposited power (kW)	5.1	32.2
In-target fission rate (s ⁻¹)	1.6×10^{13}	7.3×10^{13}
In-target ^{132}Sn production rate (s ⁻¹)	2.3×10^9	9.7×10^9
^{132}Sn release rate (s ⁻¹)	2.2×10^9	8.2×10^9
Experimental hall (s ⁻¹)	1.1×10^7	4.1×10^7

B.-H. Kang at RISP/IBS

IF Design



High radiation region:

- 6 HTS quadrupoles, 1 HTS dipole

Other region:

- 15 LTS quadrupole magnet triplets
- 7 LTS (or resistive) dipole magnets

Degrader type: Wedge@F3, F6, F8

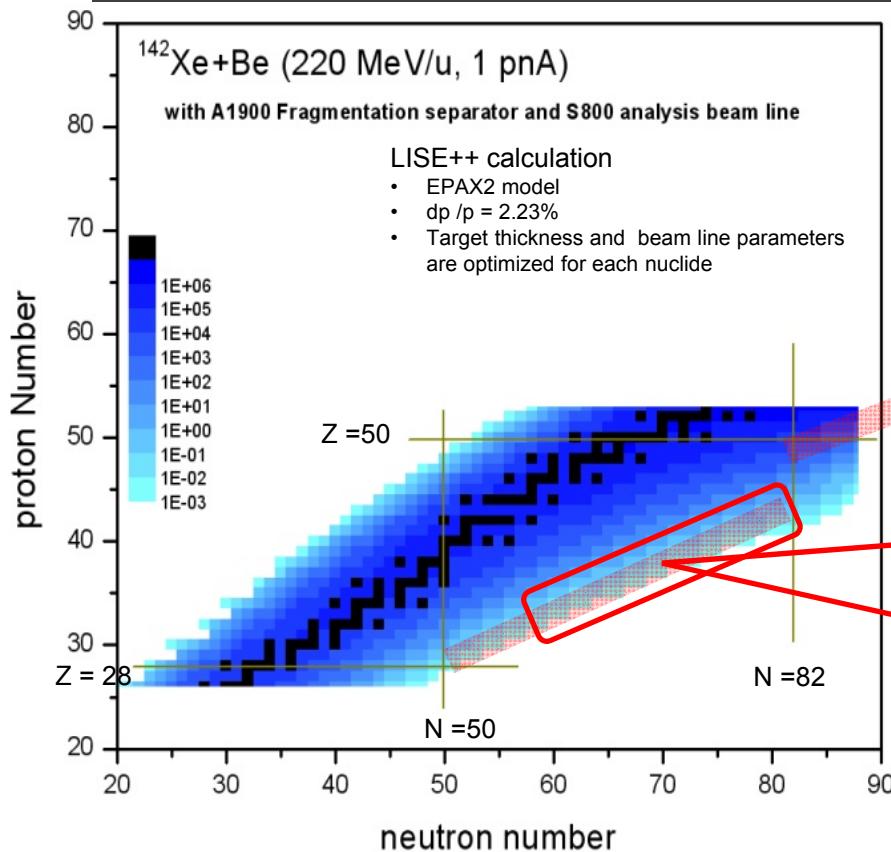
Production target at F0: Be, Graphite

Main Parameters of IF Separator

C. C. Yun, RISP/IBS

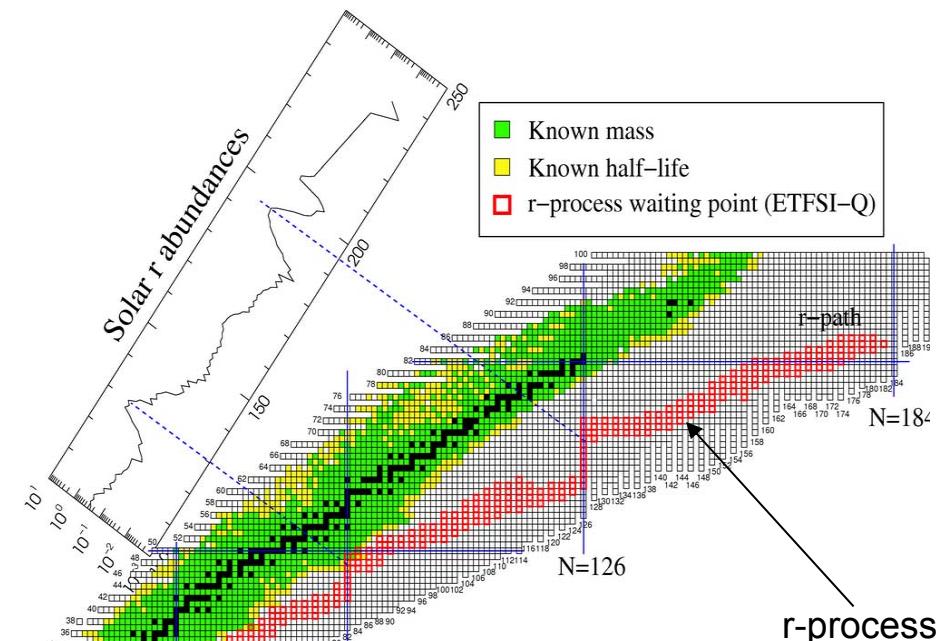
Parameter	Design Goal
Angular acceptance	90 mrad (H), 100 mrad (V)
Momentum acceptance	8%
Maximum rigidity	10 Tm
Momentum dispersion	2.7 cm/% @F6 & F8
Resolving power ($p/\Delta p$)	1,350 @F6 & F8

Production of More Exotic RIB



^{142}Xe (ISOL) \rightarrow post-accelerator \rightarrow Driver Linac
 \rightarrow IF target \rightarrow Fragment separator \rightarrow Experiment

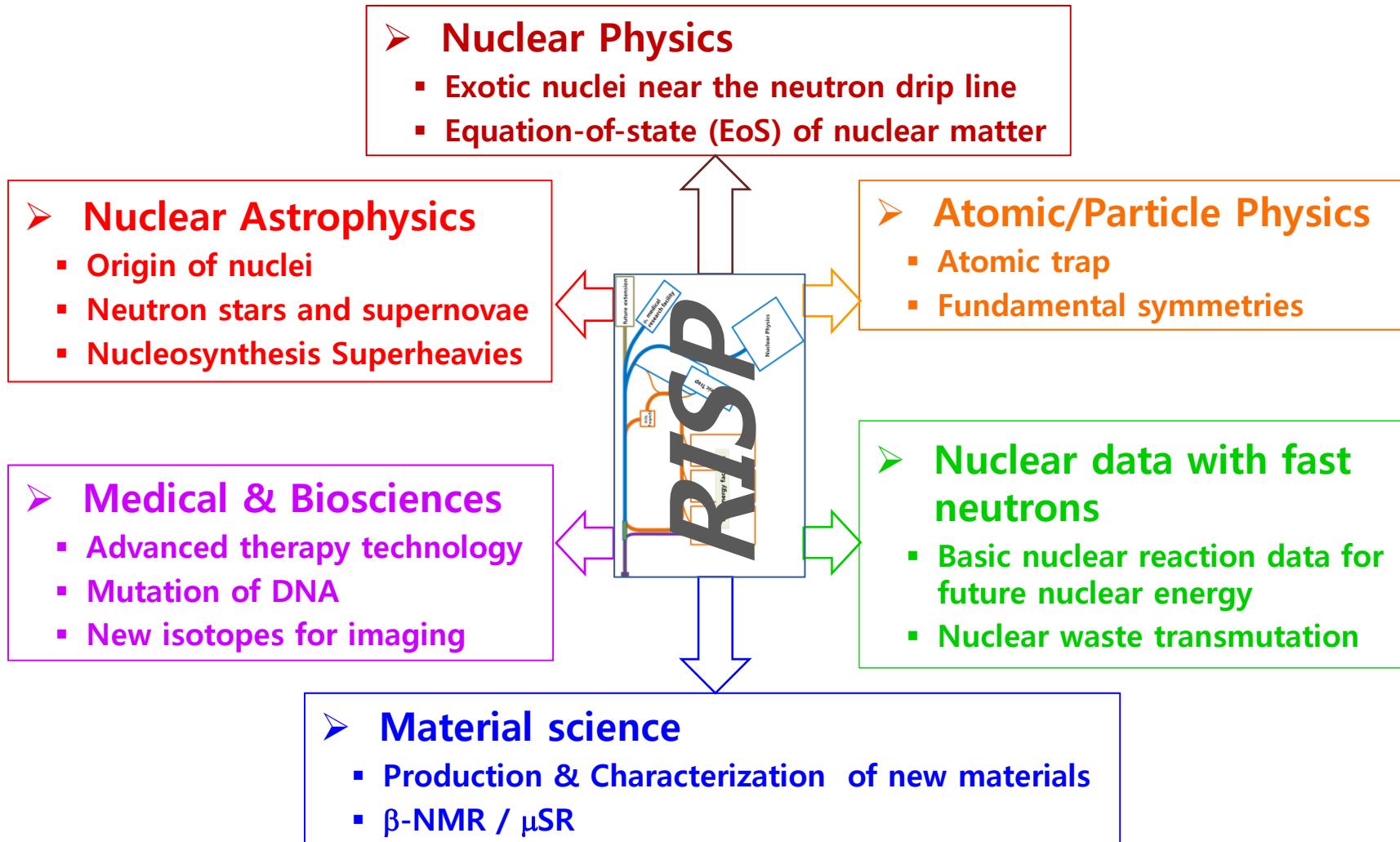
Note that $\sim 10^3$ times higher than ^{136}Xe (350 MeV/u, 10 pnA)+Be.



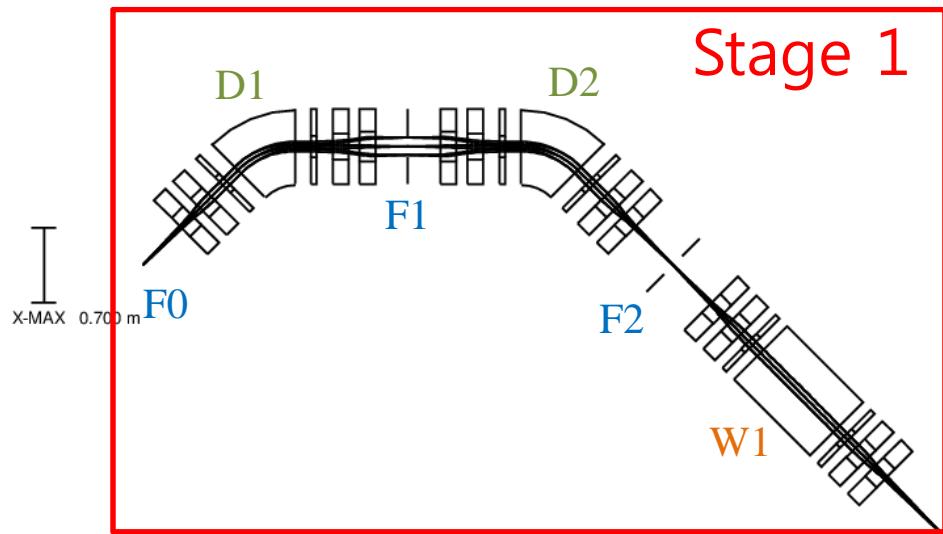
RAON can reach new n-rich isotope with rates of $10^{-3} \sim 10$ pps.

nuclide	Estimated Intensity (pps)
^{110}Y	1.8
^{110}Zr	1.8
^{114}Nb	1.1
^{116}Mo	3.8
^{118}Tc	1.4

Research Topics

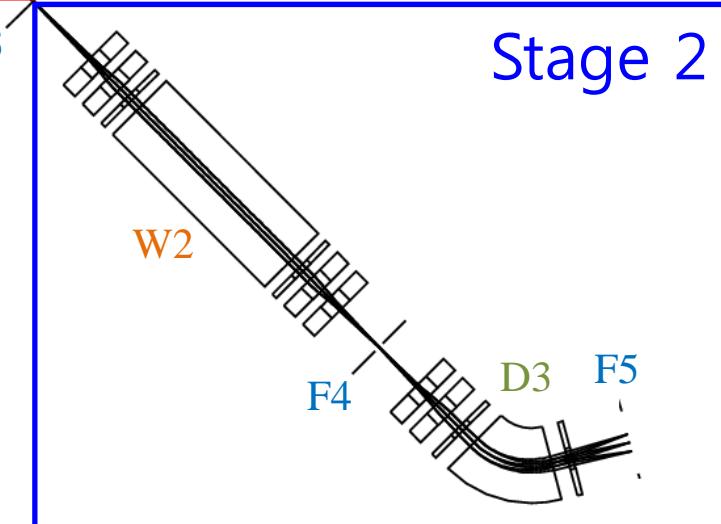


KOBRA

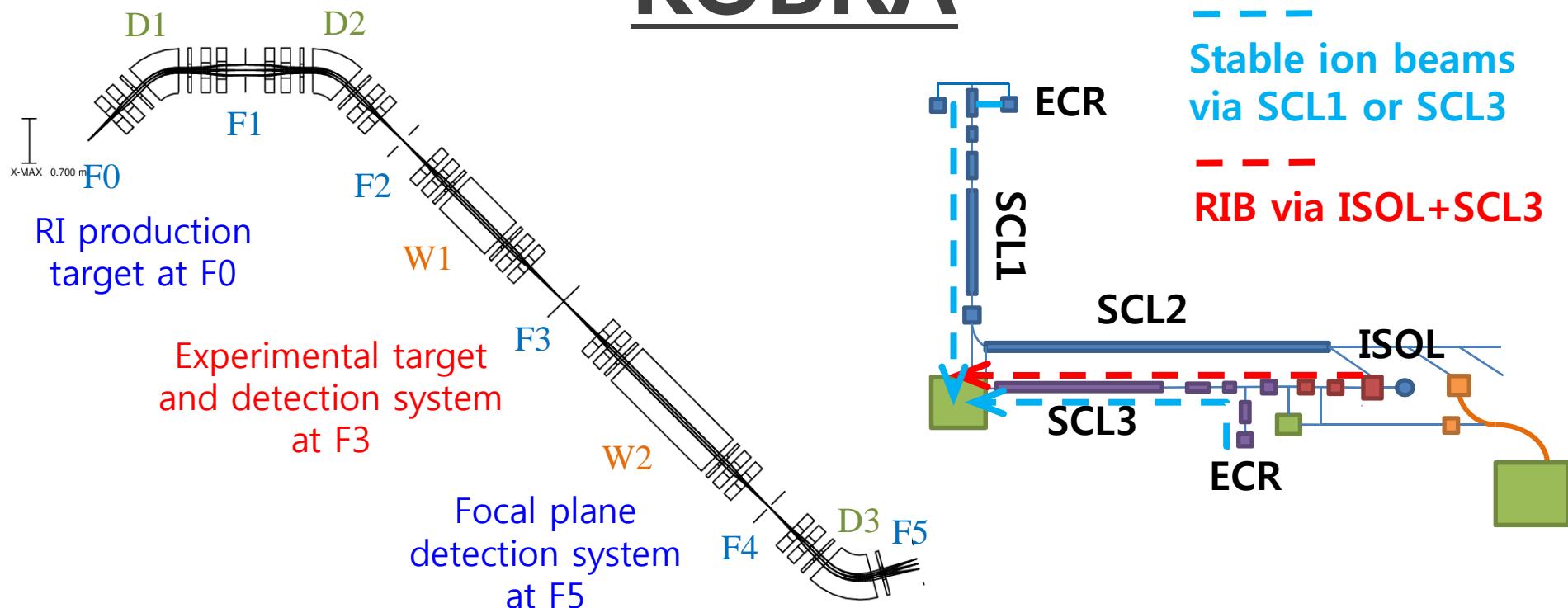


Korea Broad Acceptance
Recoil Spectrometer
and Apparatus
at low-energy
experimental area

- High-performance spectrometer with detection systems
- Main experimental facility for nuclear physics with low-energy beams up to 18.5 MeV/u



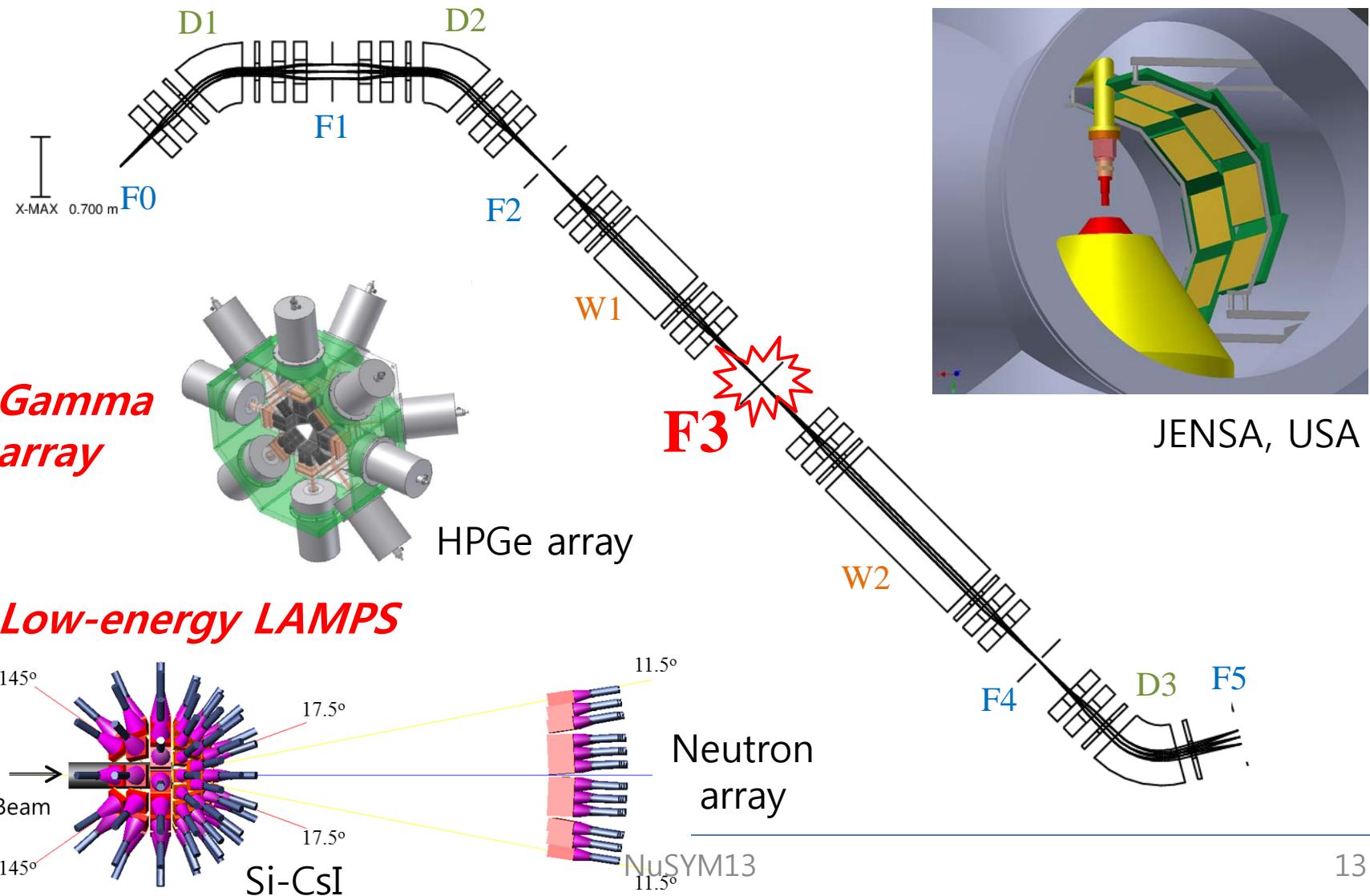
KOBRA



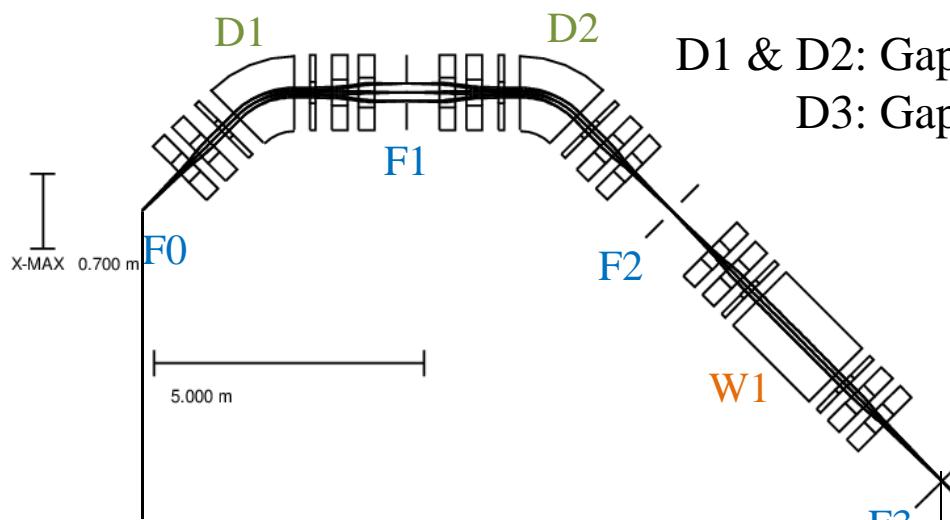
- Stage 1 (F0~F3): Production and separation of RIBs by in-flight method with high-intensity stable ion beams from ECRs
- Experimental target at F3 (available space of ~3 m): In-beam γ -ray spectroscopy, Symmetry energy & charged particle spectroscopy, Spin dependence, etc.
- Stage 2 (F3~F5): Big-bite spectrometer with Wien filter

Target and Detection Systems for KOBRA

Supersonic gas-jet target



KOBRA



D1 & D2: Gap 20 cm/Radius 1.5 m/Deflection angle 45°

D3: Gap 20 cm/Radius 1.5 m/Deflection angle 60°

W1: Length 2 m/E. gap 20 cm/ ± 300 kV

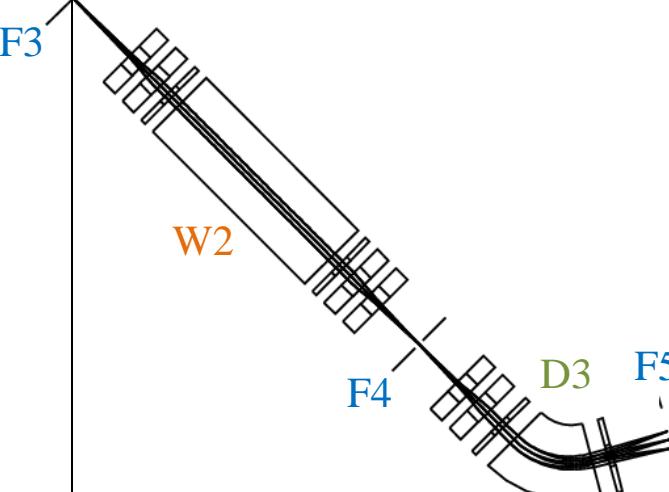
W2: Length 4 m/E. gap 20 cm/ ± 300 kV

F0, F2, F3 & F4: Achromatic focusing

F1: Dispersive focus ($D=2.03$ cm/%)

F5: Dispersive focus ($D=2.05$ cm/%)

Maximum magnetic rigidity	~ 3 T·m
Mass resolution ($M/\Delta M$)	< 200
Dispersion	~ 2 cm/%
Momentum acceptance @ stage1	14%
Angular acceptance @ stage2	40 mrad (H) 200 mrad (V)



Stage 1: In-flight Separator
F0~F3
{QQDQQ-QQDQQ-QQWQQ}

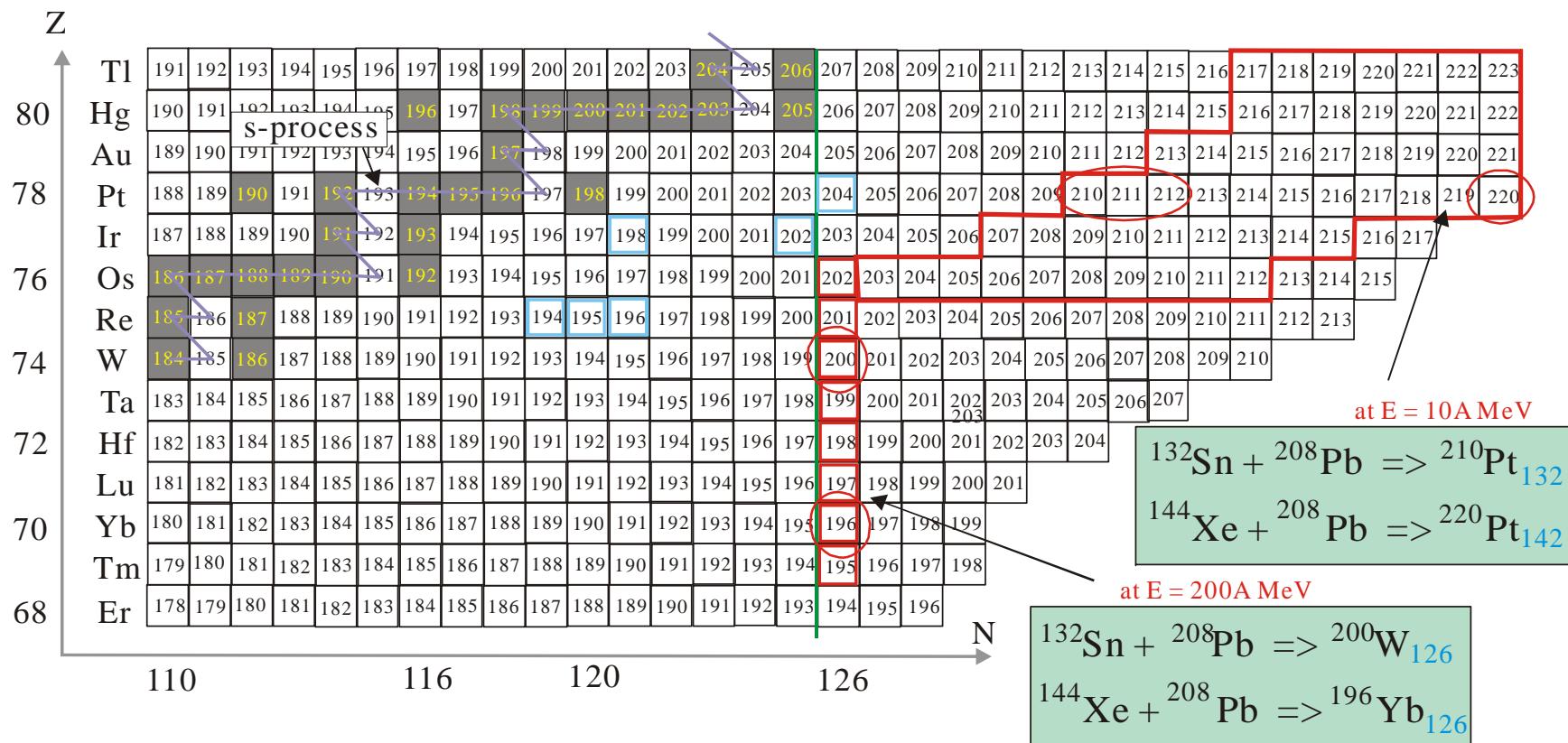
Stage 2: Large-Acceptance
Spectrometer F3~F5
{QQWQQ-QQD}

KOBRA: Physics Program

1. Nuclear structure
 - Comparison of the nuclear structures for the isobaric mirror nuclei at drip lines (charge symmetry and/or independence)
 - Resonant conditions of unbound nuclear states
 - Spin dependence of basic properties
2. Nuclear astrophysics
 - Capture reactions: (p,γ) , (α,γ) , (n,γ)
 - Transfer reactions: (d,p) , (α,p) , etc.
 - Resonant scattering: p and α resonant elastic scattering
3. Rare events
 - Super-heavy elements
 - Decay spectroscopy
4. Nuclear symmetry energy
 - Charged particle and neutron productions in central collisions
 - Electric dipole excitations

New Neutron-Rich Heavy Nuclei

“High Intensity Stable Beams in Europe”
NUPECC Report (July 2007)



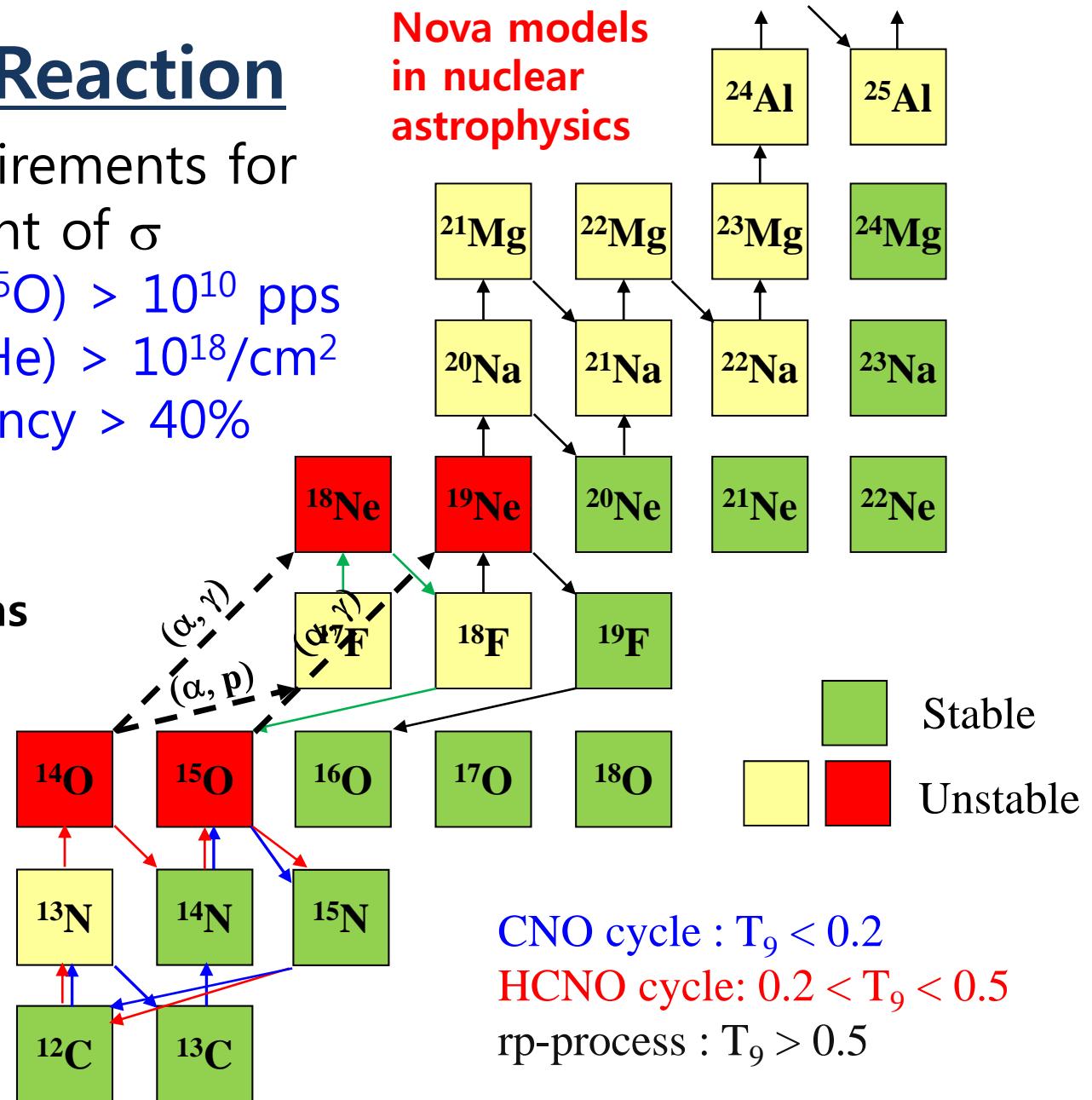
$^{15}\text{O}(\alpha,\gamma)^{19}\text{Ne}$ Reaction

- Experimental requirements for direct measurement of σ
 - Beam intensity (^{15}O) $> 10^{10}$ pps
 - Target density (^4He) $> 10^{18}/\text{cm}^2$
 - Recoil det. efficiency $> 40\%$

$\Rightarrow >1 \text{ Count/hr}$

→ Breakout paths to rp-process

- rp-process
- Hot-CNO II
- Hot-CNO I
- CNO cycle



$^{44}\text{Ti}(\alpha,\text{p})^{47}\text{V}$ Reaction

INTEGRAL

Gamma Ray Emission from Cassiopeia A

29 Sep 2006

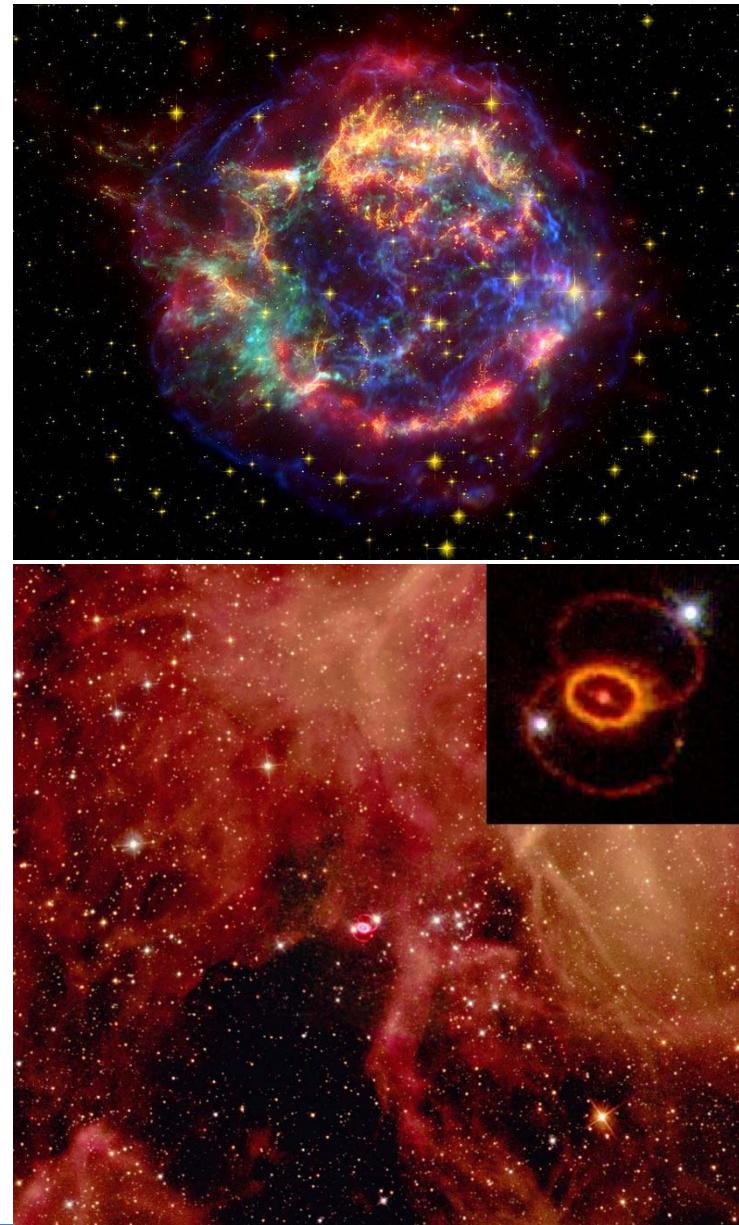
Supernovae and their remnants are the main galactic nucleosynthesis sites. Few radioactive isotopes are accessible to gamma-ray astronomy for probing these stellar explosions. Among them, ^{44}Ti is a key isotope for the investigation of the inner regions of supernovae and their young remnants.

INTEGRAL

INTEGRAL finds titanium in supernova remnant 1987A

17 Oct 2012

Astronomers using INTEGRAL have detected the first direct signature of titanium-44 in the remnant of the nearby supernova 1987A. The discovery reveals a large amount of this key isotope in the remnant, equivalent to 0.03 per cent the mass of the Sun. This value is close to upper bounds from theoretical predictions and exceeds the amount of titanium-44 observed in Cassiopeia A - the only other supernova remnant where this isotope has been found. The amount of titanium-44 found in SNR 1987A demonstrates that its radioactive decay has been powering the source for the past 22 years.



$^{44}\text{Ti}(\alpha, p)^{47}\text{V}$ Reaction

ORDER OF IMPORTANCE OF REACTIONS PRODUCING ^{44}Ti AT $\eta = 0$

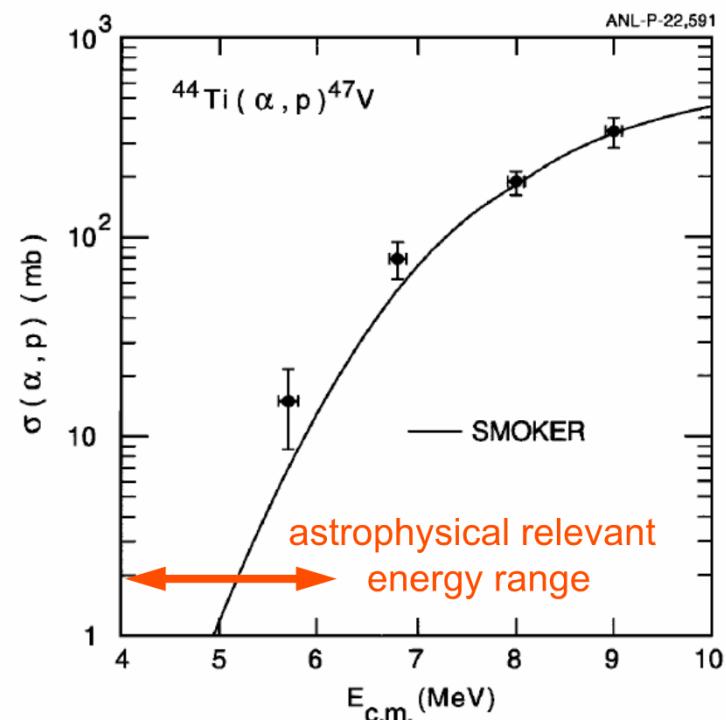
REACTION RATE MULTIPLIED BY 1/100	REACTION RATE MULTIPLIED BY 100		
	^{44}Ti Change (percent)	REACTION	^{44}Ti Change (percent)
$^{44}\text{Ti}(\alpha, p)^{47}\text{V}$	+173	$^{45}\text{V}(p, \gamma)^{46}\text{Cr}$	-98
$\alpha(2\alpha, \gamma)^{12}\text{C}$	-100	$\alpha(2\alpha, \gamma)^{12}\text{C}$	+67
$^{40}\text{Ca}(\alpha, \gamma)^{44}\text{Ti}$	-72	$^{44}\text{Ti}(\alpha, p)^{47}\text{V}$	-89
$^{45}\text{V}(p, \gamma)^{46}\text{Cr}$	+57	$^{44}\text{Ti}(\alpha, \gamma)^{48}\text{Cr}$	-61
$^{57}\text{Ni}(p, \gamma)^{58}\text{Cu}$	-47	$^{57}\text{Co}(p, n)^{57}\text{Ni}$	+25
$^{57}\text{Co}(p, n)^{57}\text{Ni}$	-33	$^{40}\text{Ca}(\alpha, \gamma)^{44}\text{Ti}$	+22
$^{13}\text{N}(p, \gamma)^{14}\text{O}$	-16	$^{57}\text{Ni}(n, \gamma)^{58}\text{Ni}$	+10
$^{58}\text{Cu}(p, \gamma)^{59}\text{Zn}$	-14	$^{54}\text{Fe}(\alpha, n)^{57}\text{Ni}$	+9.4
$^{36}\text{Ar}(\alpha, p)^{39}\text{K}$	-11	$^{36}\text{Ar}(\alpha, p)^{39}\text{K}$	+5.5
$^{12}\text{C}(\alpha, \gamma)^{16}\text{O}$	+3.5	$^{36}\text{Ar}(\alpha, \gamma)^{40}\text{Ca}$	+5.3

The et al., *Astrophys. J.* 504 (1998)

- Presently, TRIUMF, CERN, CNS CRIB are working on this reaction.
- At RISP, the direct measurement will be possible with **an active target in the IF mode of KOBRA**.

Measurement at FMA at Argonne Nat. Lab.

^{44}Ti intensity of $\sim 5 \times 10^5 \text{ s}^{-1}$



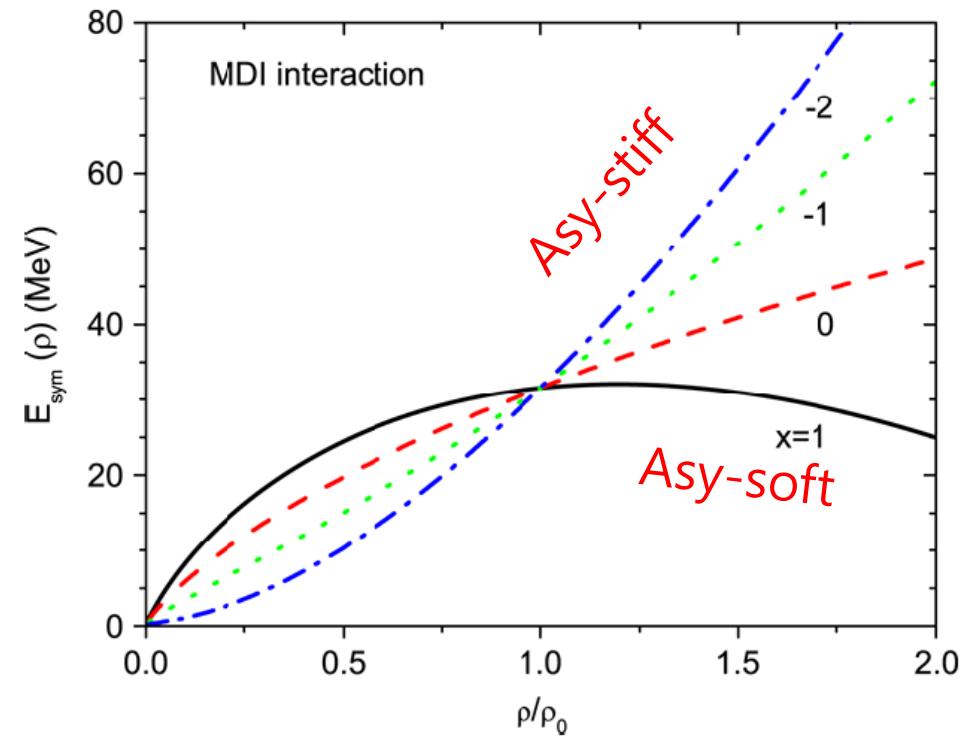
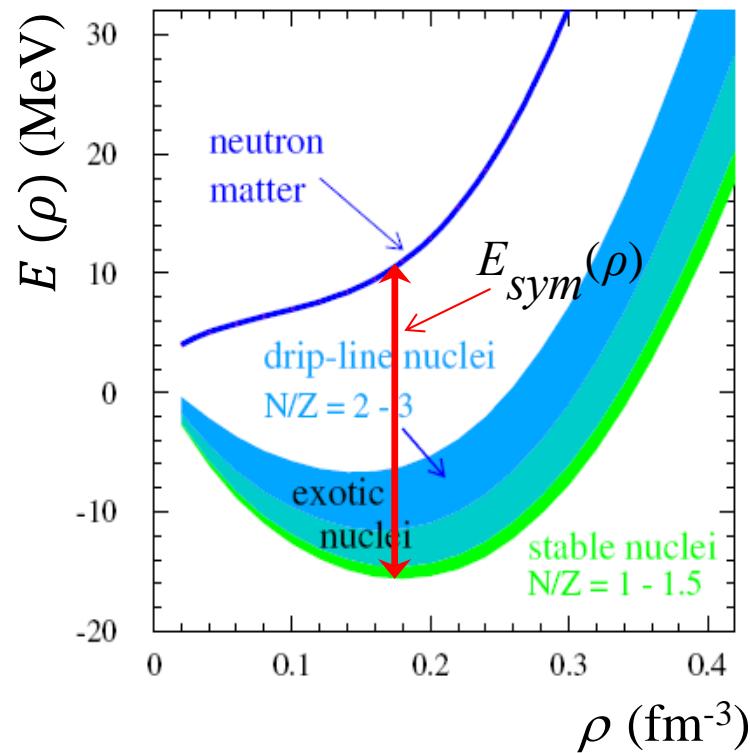
Sonzogni et al., *Phys. Rev. Lett.* 84 (2000)

EOS & Symmetry Energy

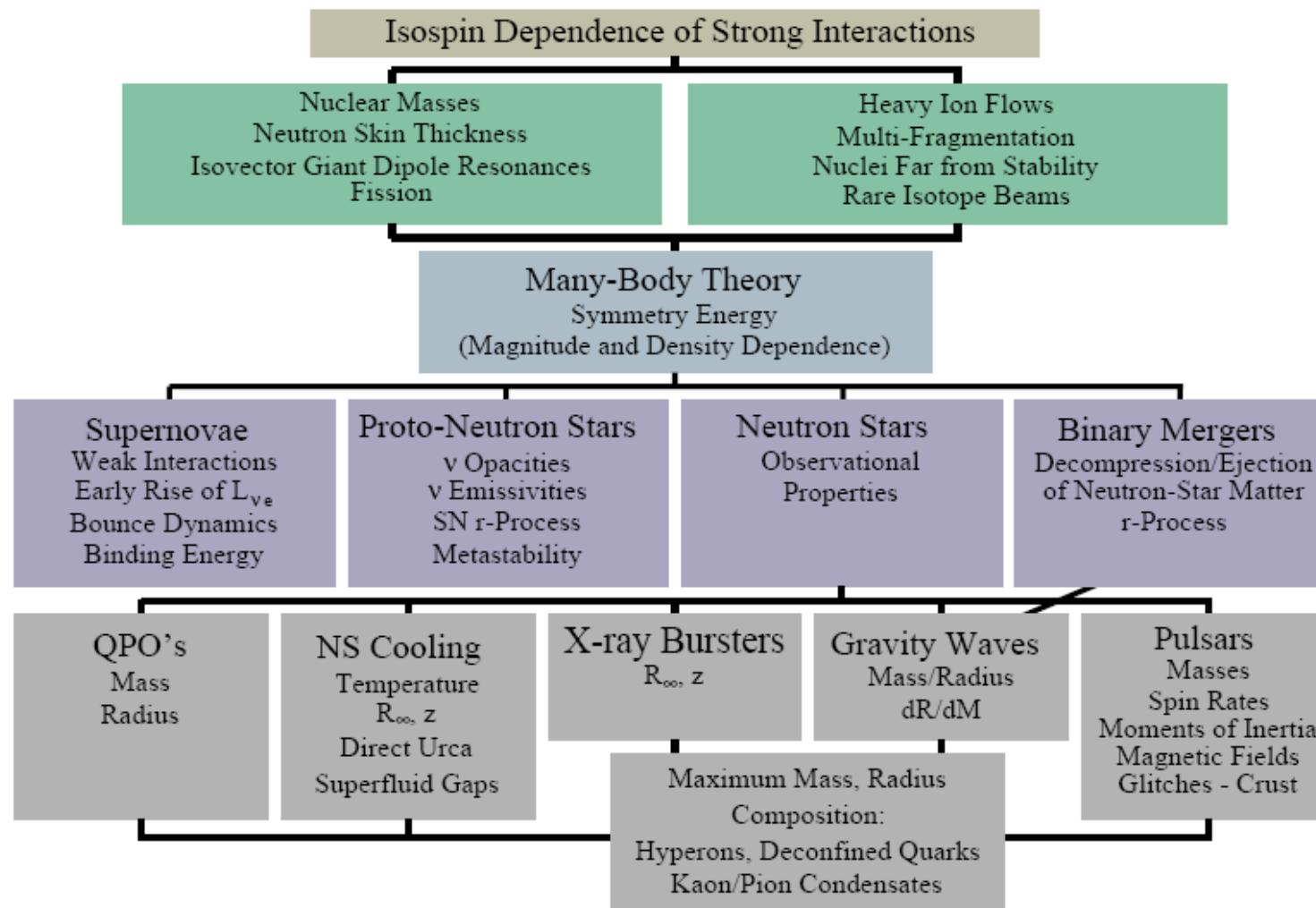
$$E(\rho, \delta) = E(\rho, \delta = 0) + E_{sym}(\rho)\delta^2 + \mathcal{O}(\delta^4) + \dots$$

where $\rho = \rho_n + \rho_p$ and $\delta = \frac{\rho_n - \rho_p}{\rho_n + \rho_p}$

L.W. Chen et al.,
PRL 94, 032701 (2005)



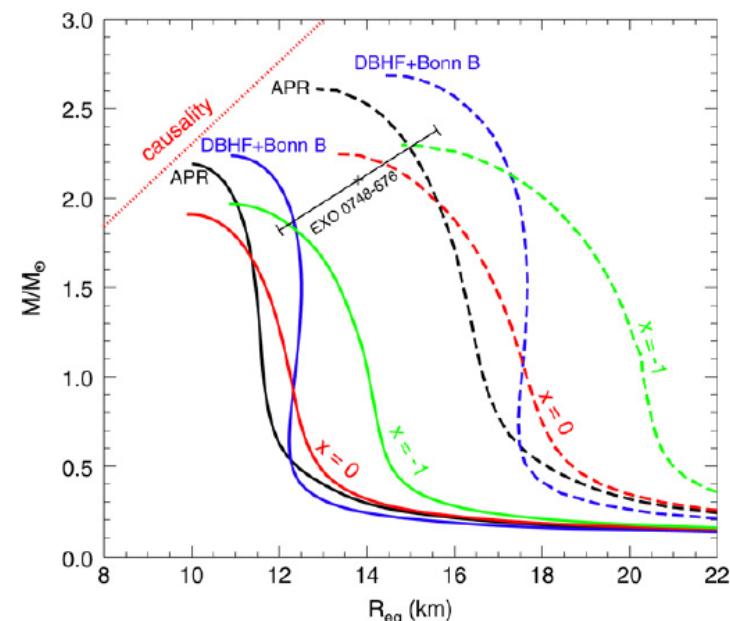
Symmetry Energy



- A.W. Steiner, M. Prakash, J.M. Lattimer and P.J. Ellis, Physics Report 411, 325 (2005)

Symmetry Energy & Neutron Stars

- Neutron star stability against gravitational collapse
- Determine stellar density profile and internal structure
- Observational consequences
 - Cooling rates of proto-neutron stars
 - Stellar masses, radii & moment of inertia from temperatures & luminosities of X-ray bursters
- M vs. R relationship
 - Uncertainty of softness of EOS and influence of E_{sym}
- Need to provide additional laboratory constraints at specific densities



P. Krastev, B.A. Li, and A. Worley,
Astrophys. J. 676, 1170 (2008)

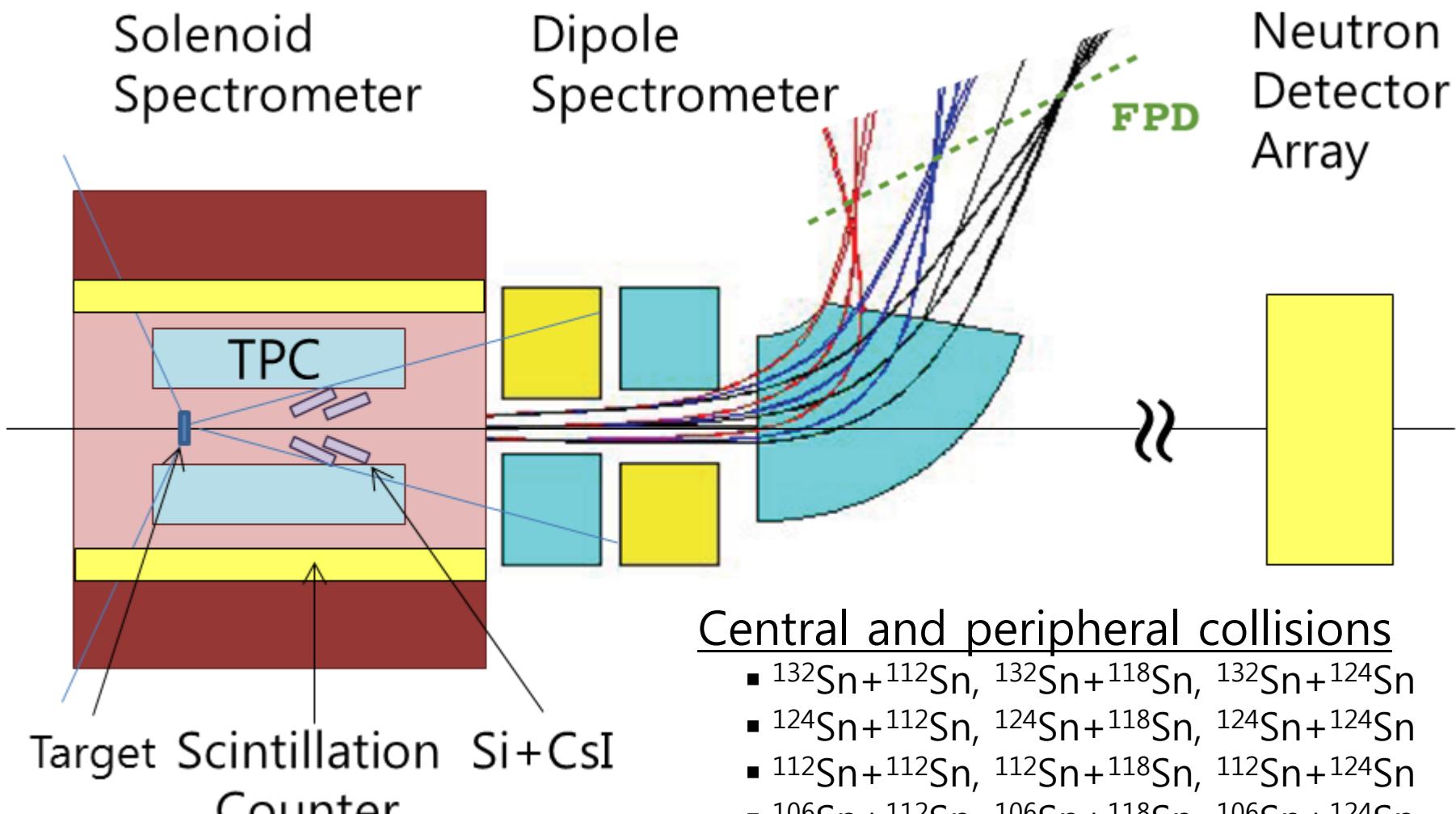
Experimental Observables

1. Particle ratios of mirror nuclei and pions
 - n/p, $^3\text{H}/^3\text{He}$, $^7\text{Li}/^7\text{Be}$, π^-/π^+ , etc.
2. Collective flow
 - Directed (or sideward) and elliptic flow parameters of n, p, and heavier fragments
3. Various isospin-dependent phenomena
 - Isoscaling in nuclear multifragmentation
 - Isospin transport/diffusion
4. Electric dipole resonances
 - Energy spectra of the excitation energy and/or gammas
 - PDR~the size of n-skin for unstable nuclei

Experimental Requirements

1. We need to accommodate
 - Large acceptance
 - Precise measurement of momentum (or energy) for variety of particle species, including $\pi^{+/-}$ and neutrons, with high efficiency
 - Gamma detection for electric dipole resonances
 - Keep flexibility for other physics topics
2. This leads to the design of **LAMPS**
 - Large-acceptance Multipurpose Spectrometer
 - Low-energy LAMPS at F3 of KOBRA
 - High-energy LAMPS

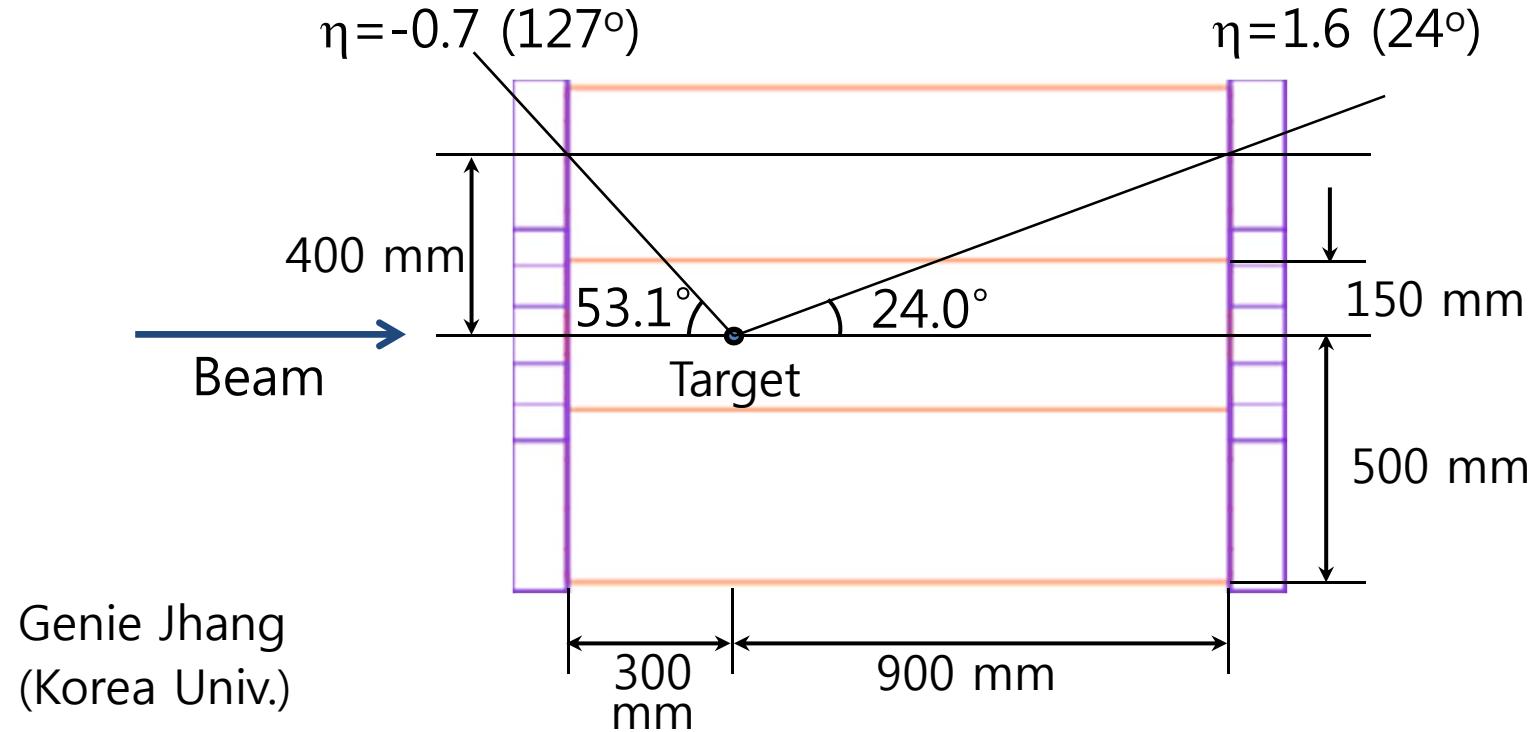
LAMPS



Central and peripheral collisions

- $^{132}\text{Sn} + ^{112}\text{Sn}$, $^{132}\text{Sn} + ^{118}\text{Sn}$, $^{132}\text{Sn} + ^{124}\text{Sn}$
- $^{124}\text{Sn} + ^{112}\text{Sn}$, $^{124}\text{Sn} + ^{118}\text{Sn}$, $^{124}\text{Sn} + ^{124}\text{Sn}$
- $^{112}\text{Sn} + ^{112}\text{Sn}$, $^{112}\text{Sn} + ^{118}\text{Sn}$, $^{112}\text{Sn} + ^{124}\text{Sn}$
- $^{106}\text{Sn} + ^{112}\text{Sn}$, $^{106}\text{Sn} + ^{118}\text{Sn}$, $^{106}\text{Sn} + ^{124}\text{Sn}$
etc.

Time Projection Chamber

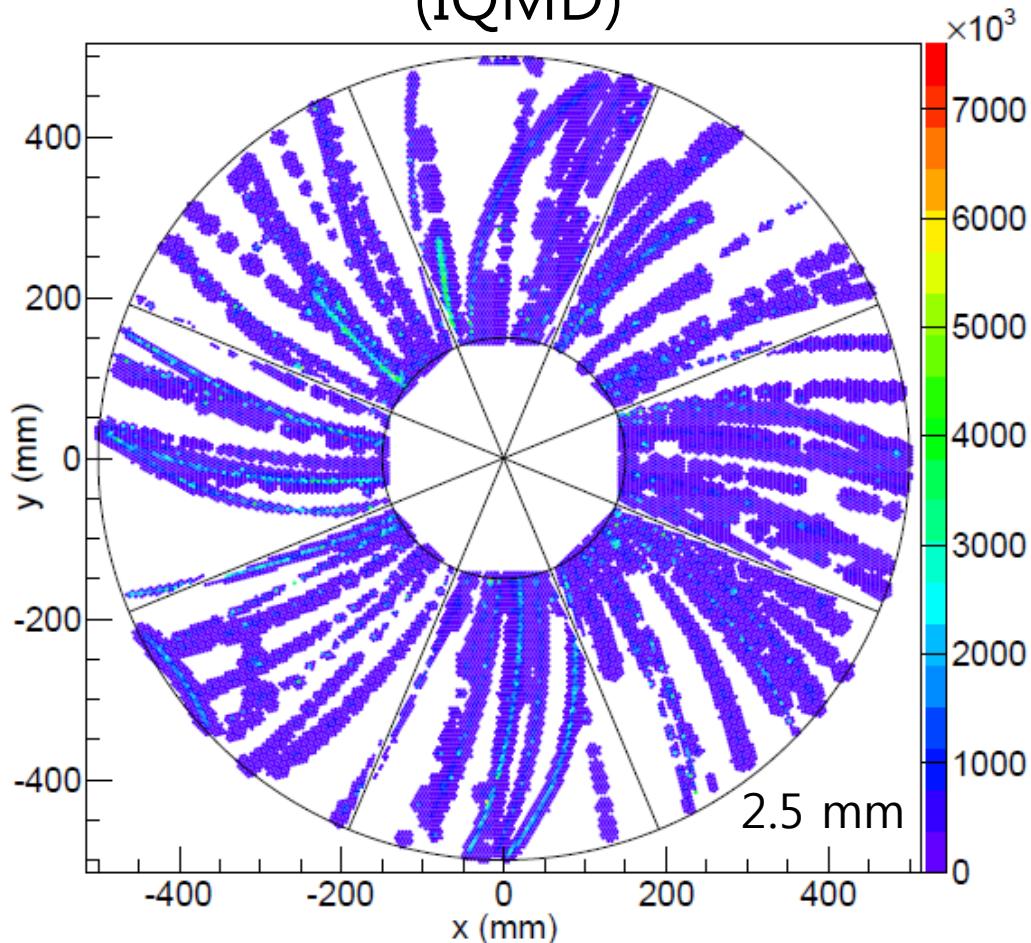


- Simulation with triple GEM readouts at both ends by Garfield++
 - Gas mixture: Ar 90%+CO₂ 10%, Voltage for each foil: 450 V
 - $\langle \text{Gain} \rangle \sim 1.4 \times 10^6$, $\langle \text{Drift velocity} \rangle \sim 50 \text{ mm}/\mu\text{s}$
 - $\langle \text{Dispersion} \rangle$ after 60 cm (maximum drift distance) < 3 mm

Time Projection Chamber

Central Au+Au at 250 AMeV
(IQMD)

Genie Jhang (Korea Univ.)

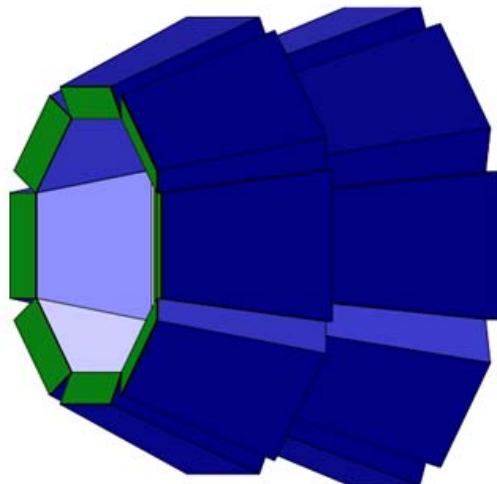


Color scale: the number of electrons in each pad

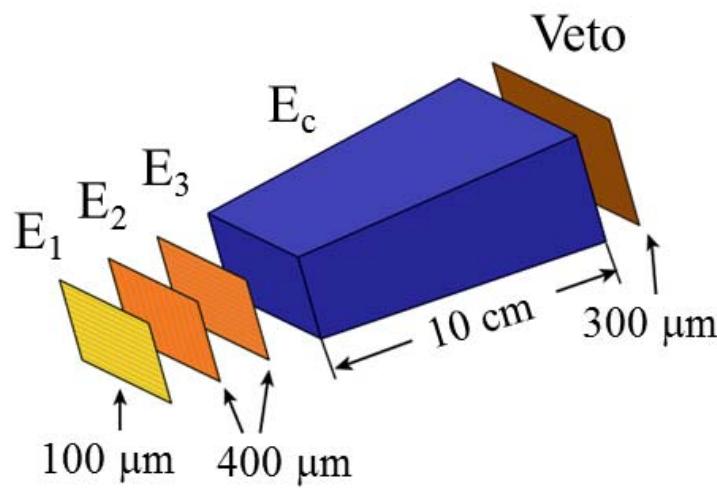
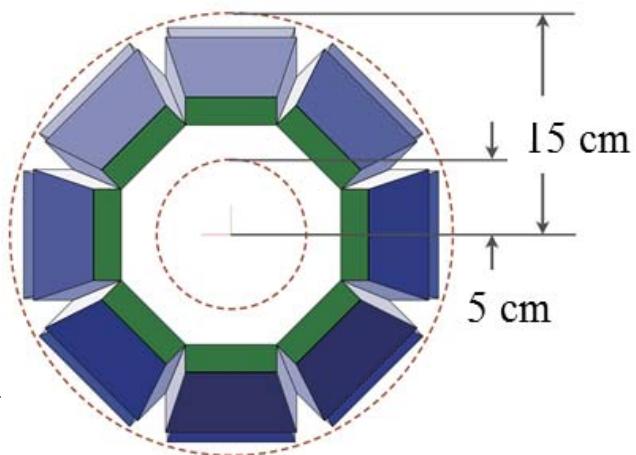
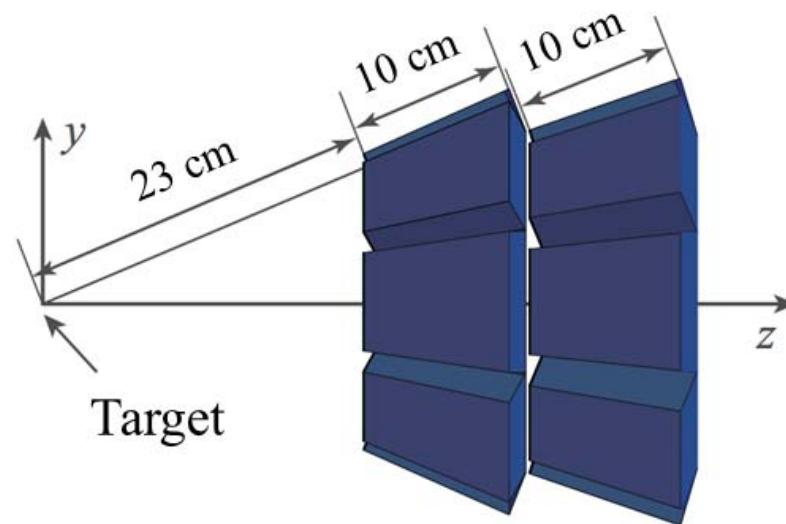
- Pad
 - Shape: hexagonal
 - Total number 90,000 for 2.5 mm
20,000 for 5 mm
- Signal processing
 - GET: General Electronics for TPC

Si-CsI

$1.6 < \eta < 2.1$
 $(14^\circ < \theta_{\text{Lab}} < 24^\circ)$

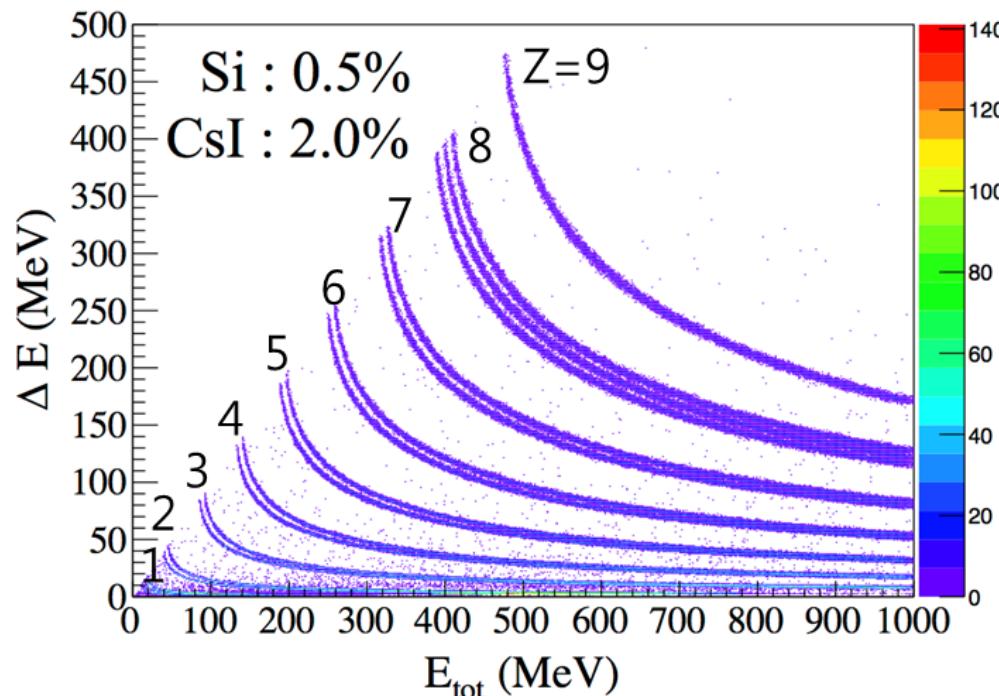


Suhyun & Songkyo Lee
(Korea Univ.)



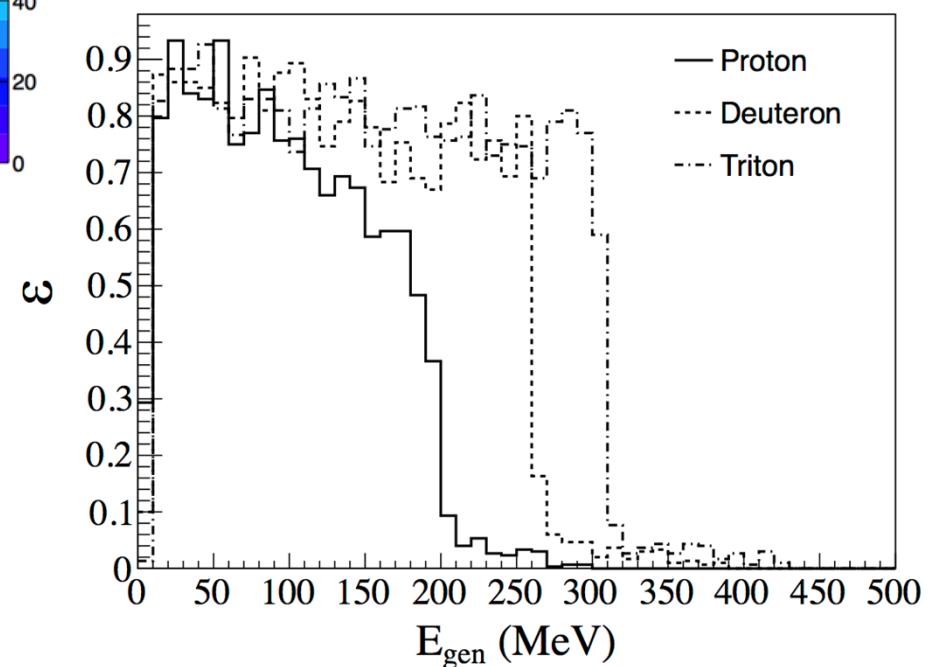
Si-CsI

Suhyun & Songkyo Lee
(Korea Univ.)



$$\Delta E = E_1 + E_2 + E_3$$

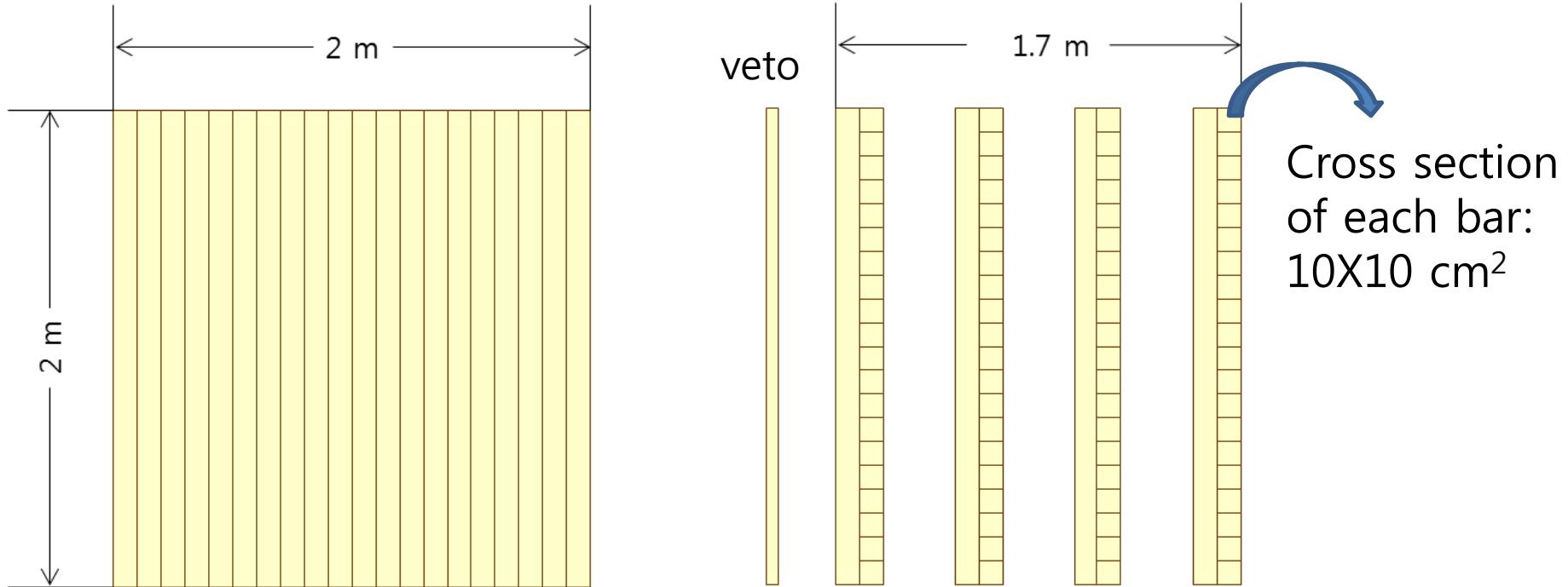
$$E_{\text{tot}} = \Delta E + E_c$$



- $\Delta E = E_1 + E_2 + E_3$ is preferred at high energies
- $\Delta E = E_1$ is preferred at low energies

Neutron Detector Array

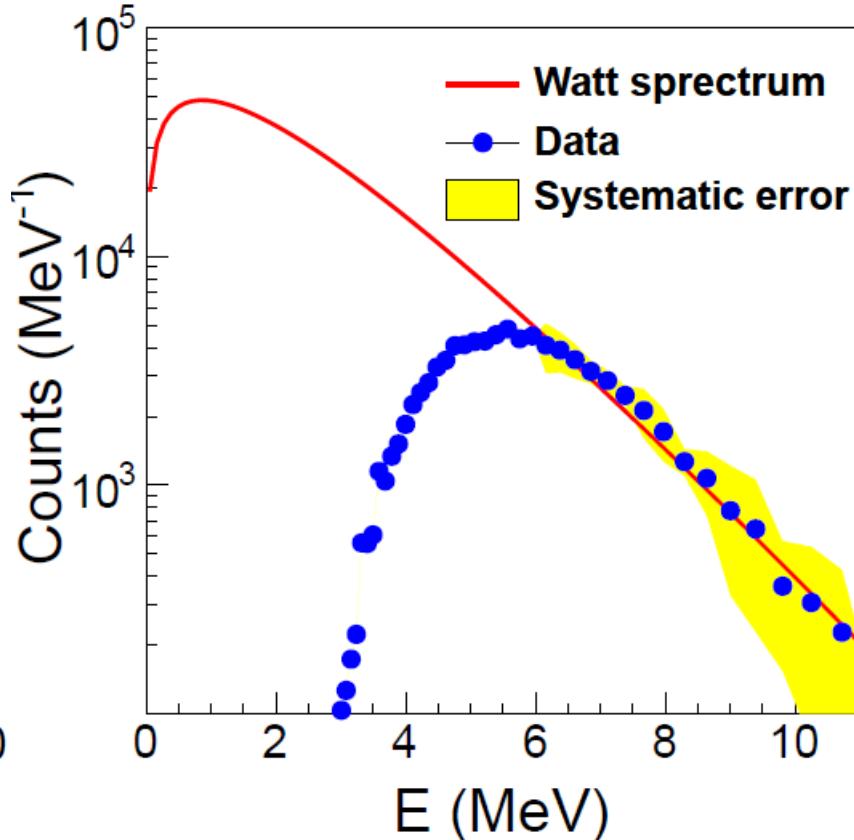
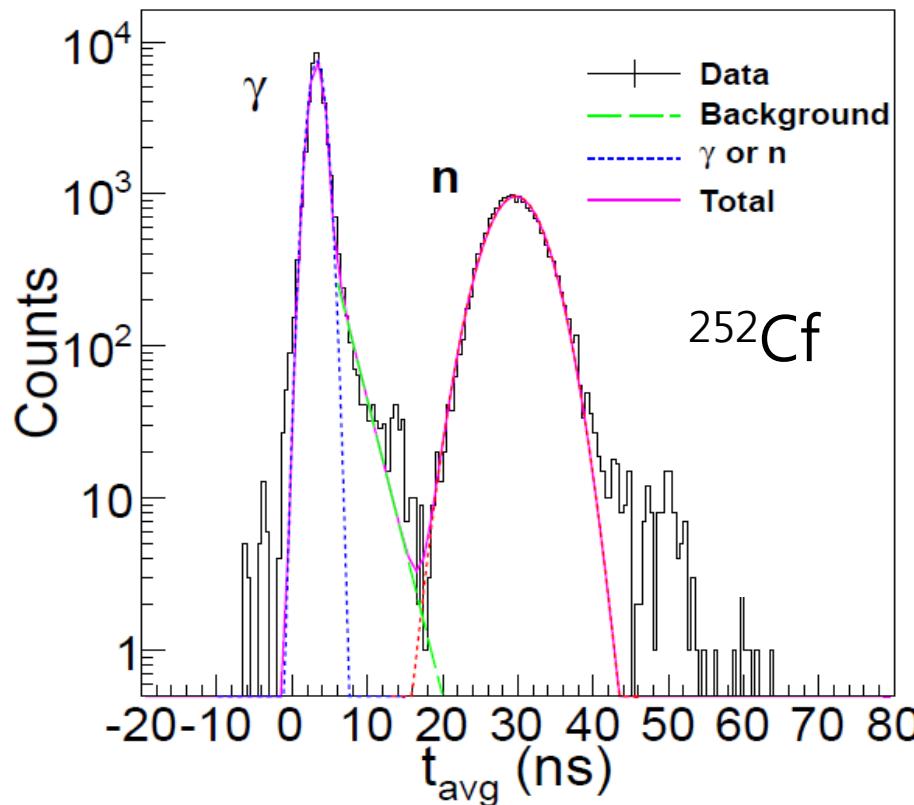
Kisoo Lee & Eunah Joo (Korea Univ.)



- Construction of the prototype and test with radiation sources
 - Dimension: $0.1 \times 0.1 \times 1.0 \text{ m}^3$
 - Sources: ^{60}Co and ^{252}Cf
 - Time resolution: 488 ps, Position resolution: $\sim 8 \text{ cm}$ for CFD

Neutron Detector Array

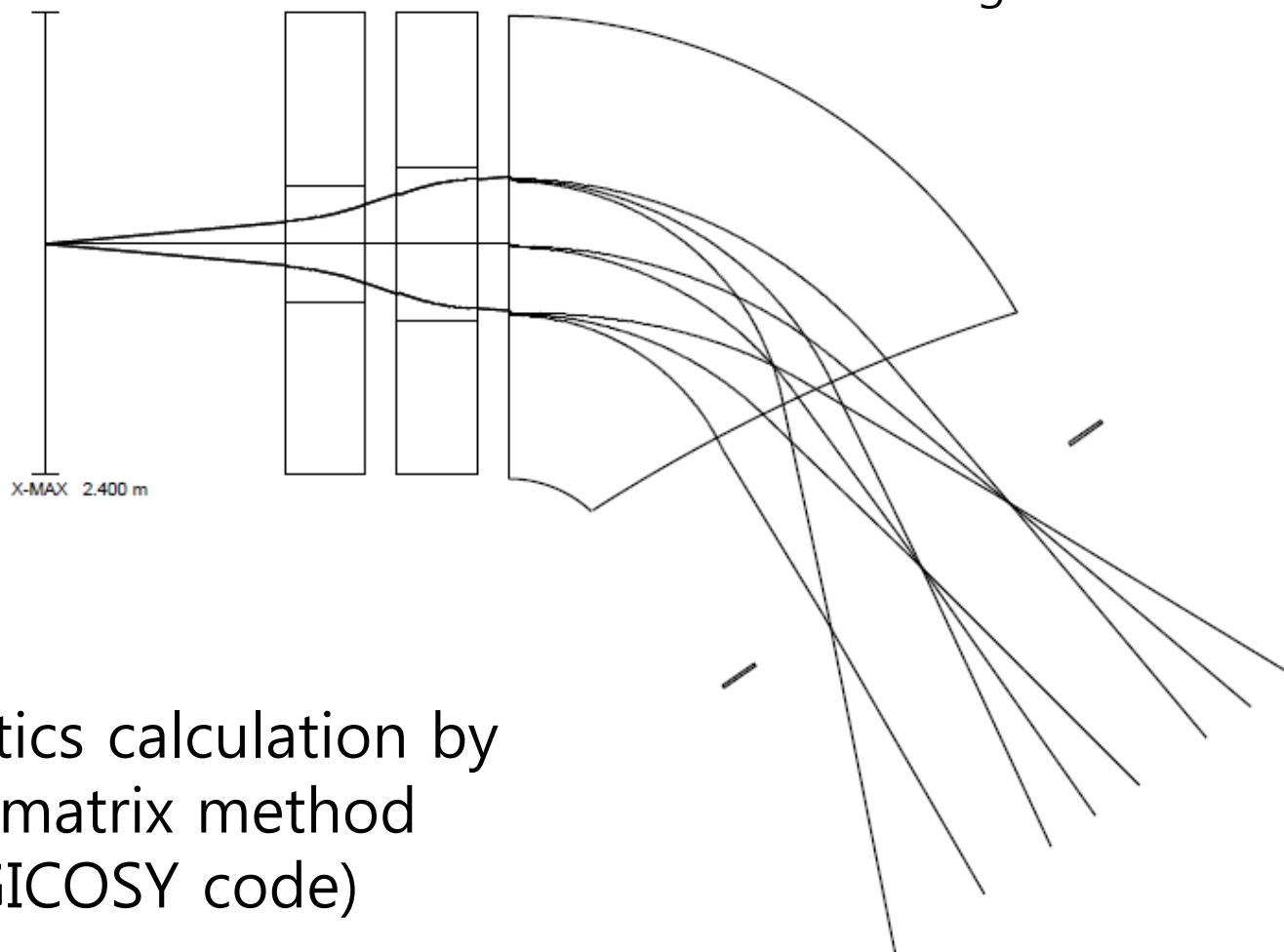
Kisoo Lee & Benard Mulilo (Korea Univ.)



- Watt spectrum: $\frac{dN}{dE} \propto e^{-aE} \sinh \sqrt{bE}$
with $a=0.88 \text{ MeV}^{-1}$ and $b=2.0 \text{ MeV}^{-1}$
Ref) B. Watt, Physical Review 87, 1037 (1952)

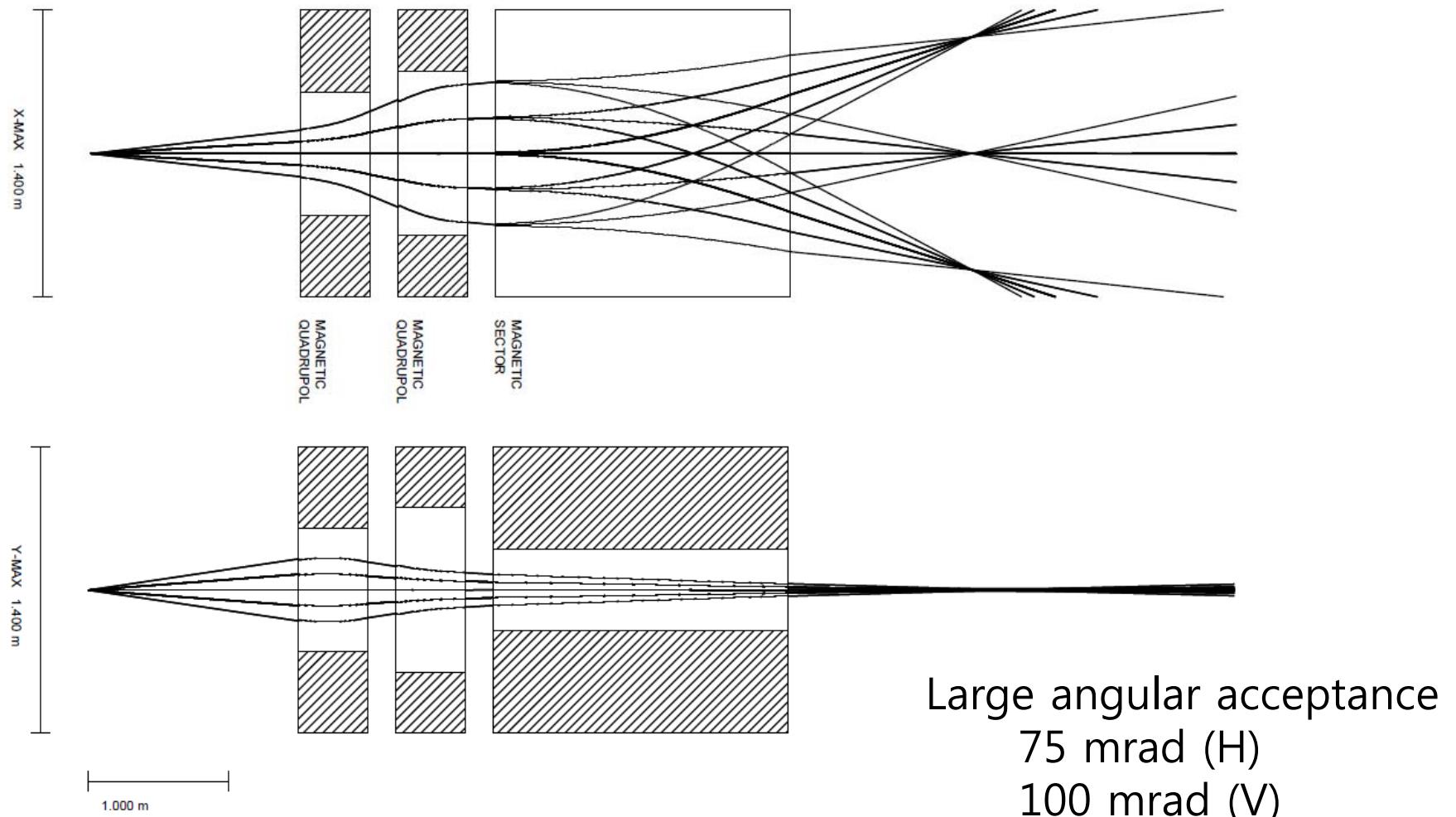
Dipole Spectrometer

Songkyo Lee (Korea Univ.)
Chong Cheoul Yun (RISP/IBS)

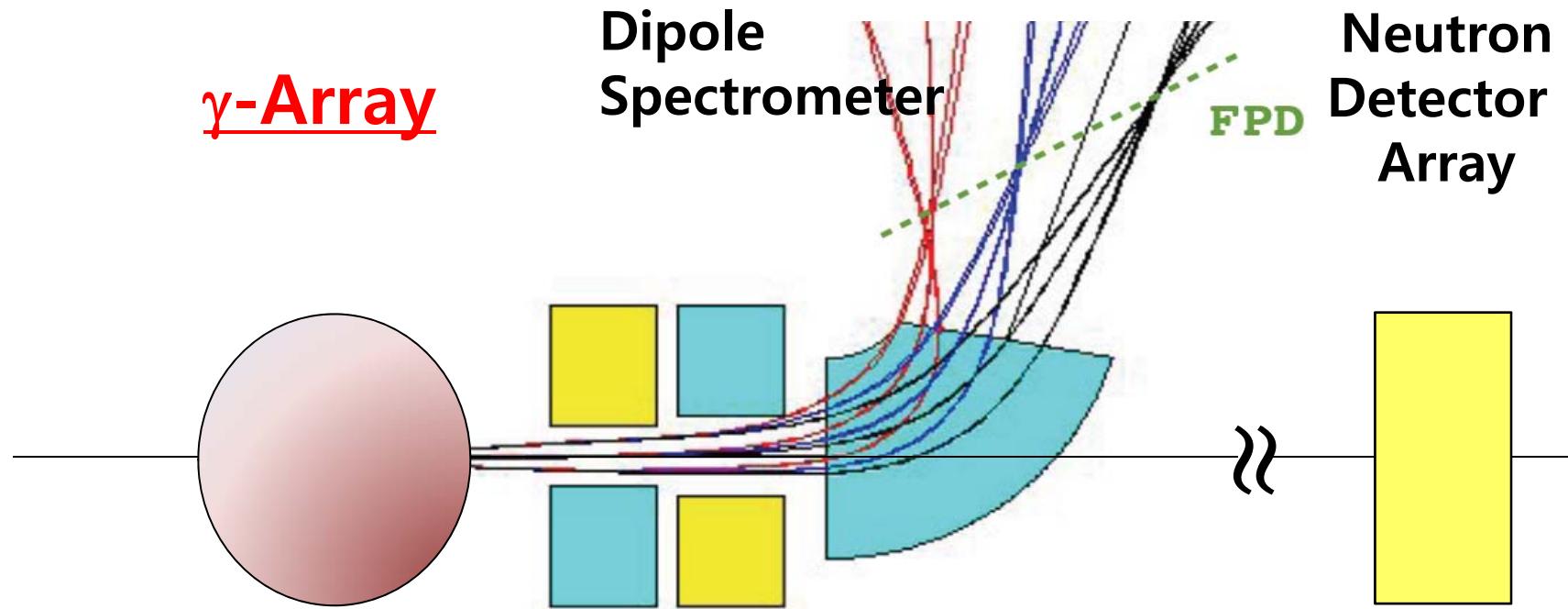


Dipole Spectrometer

Songkyo Lee (Korea Univ.) & Chong Cheoul Yun (RISP/IBS)



Coulomb Breakup & Transfer Reactions



- Photoabsorption measurements
 - PDR/GDR measurements via $^{124,130,132}\text{Sn} + ^{208}\text{Pb}$, $^{68,70,72}\text{Ni} + ^{208}\text{Pb}$, $^{50,54,60}\text{Ca} + ^{208}\text{Pb}$, etc.
 - 1n and 2n removal cross sections for unstable nuclei
 - Measuring E^* from beam fragment, n's, and γ 's
- For example, 2n transfer reaction is important for the structure

Summary

1. RAON
 - First large-scale facility for nuclear physics in Korea
2. KOBRA
 - Broad acceptance recoil spectrometer at low-energy experimental area
 - To cover nuclear structure, nuclear astrophysics, super-heavy elements, and nuclear symmetry energy
3. LAMPS
 - Large-acceptance multipurpose spectrometer at high-energy experimental area
(Low-energy version of LAMPS at KOBRA)
 - Primary purpose is to measure the nuclear symmetry energy at sub- and supra-saturation densities
 - Useful also to study various photoabsorption processes and transfer reactions