

Density Dependence of Symmetry Term from High Energy Experiments

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- Symmetry energy
- Observables and sensitivity (INDRA, FOPI)
- FOPI/LAND data (1993/2011)
- ASY-EOS Experiment (2011)
- Challenges and future directions

Symmetry Energy (nucleus)

$$\frac{E_B}{A} = \underbrace{-a_V}_{\text{Volume}} + \underbrace{a_S \frac{1}{A^{1/3}}}_{\text{Surface}} + \underbrace{a_C \frac{Z^2}{A^{4/3}}}_{\text{Coulomb}} + \underbrace{a_A \left(\frac{N-Z}{A} \right)^2}_{\text{Symmetry}} \pm \underbrace{a_P \frac{1}{A^{3/2}}}_{\text{Pairing}}$$

Binding Energy/Nucleon
(semiempirical
Bethe-Weizsäcker)
acc. better than 1%

$$a_V = 15.8$$

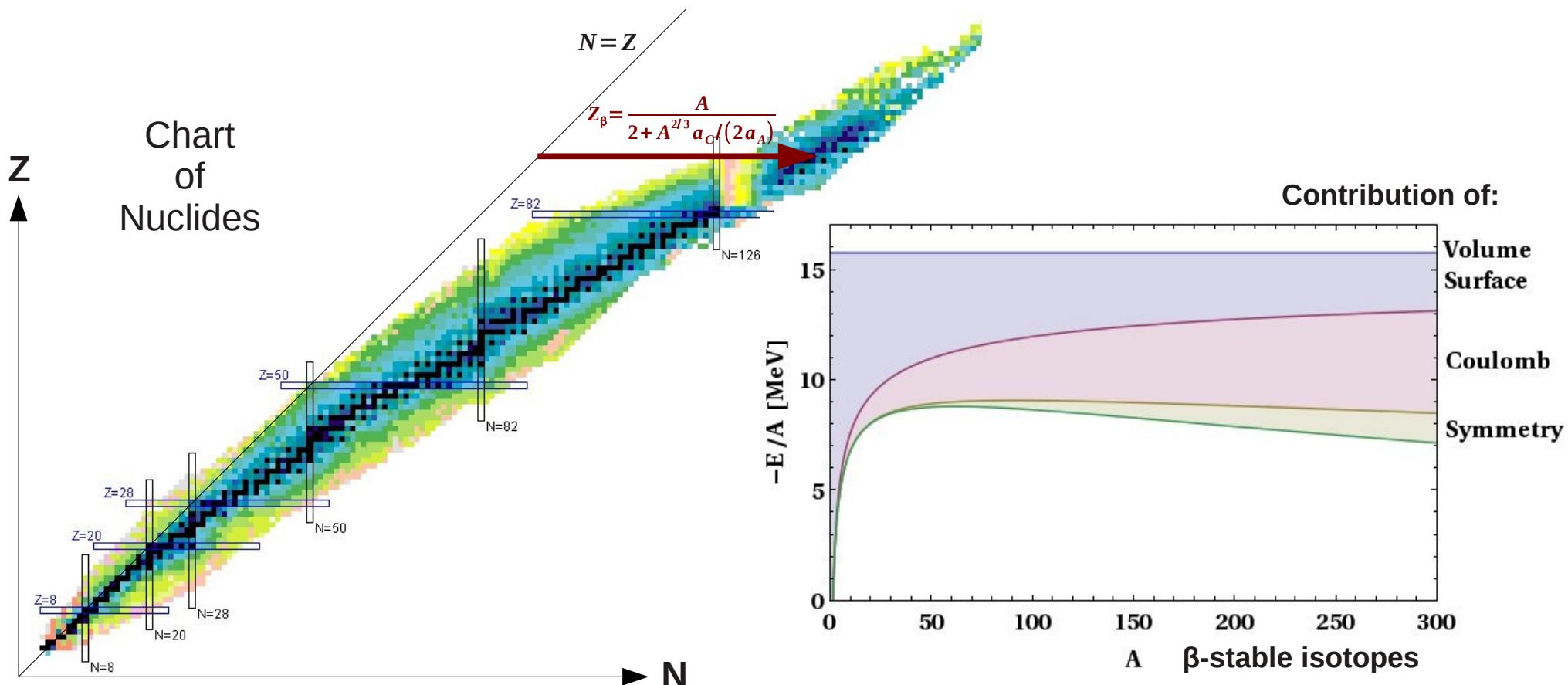
$$a_S = 18.0$$

$$a_C = 0.72$$

$$a_A = 23.5$$

$$a_P = 11.5 \mid 0$$

[MeV]



Symmetry Energy (nuclear matter)

$$\frac{E_B}{A} = \underbrace{-a_V}_{\text{Volume}} + \underbrace{a_S \frac{1}{A^{1/3}}}_{\text{Surface}} + \underbrace{a_C \frac{Z^2}{A^{4/3}}}_{\text{Coulomb}} + \underbrace{a_A \left(\frac{N-Z}{A} \right)^2}_{\text{Symmetry}} \pm \underbrace{a_P \frac{1}{A^{3/2}}}_{\text{Pairing}}$$

Energy per nucleon in nuclear matter (EoS):

$$E(\rho, \delta) = E(\rho, 0) + E_{sym}(\rho) \delta^2 + o(\delta^4)$$

dominant symmetric matter ($N=Z$) term:

$$E(\rho, 0) \approx -a_V + \frac{K}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots$$

symmetry term:

$$E_{sym}(\rho) \approx a_A^V + \frac{L}{3} \left(\frac{\rho - \rho_o}{\rho_o} \right) + \frac{K_{sym}}{18} \left(\frac{\rho - \rho_o}{\rho_o} \right)^2 + \dots$$

$\rho_n, \rho_p \rightarrow$ neutron, proton densities

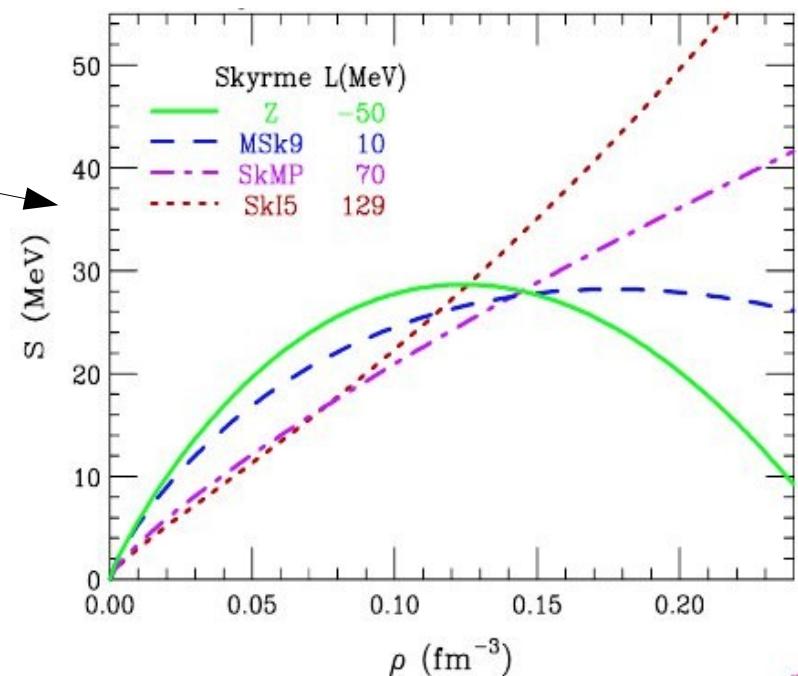
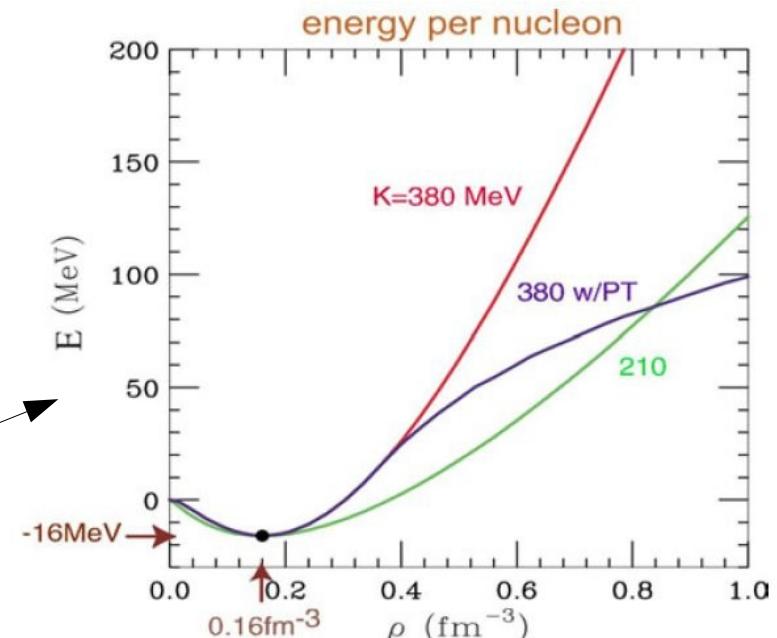
$\rho = \rho_n + \rho_p \rightarrow$ nucleon density

$\delta = \frac{\rho_n - \rho_p}{\rho} \rightarrow$ isospin asymmetry

$L = 3\rho_o \left. \frac{\partial E_{sym}}{\partial \rho} \right|_{\rho=\rho_o} \rightarrow \sim$ symmetry pressure

$K = 9\rho_o^2 \left. \frac{\partial^2 E}{\partial \rho^2} \right|_{\rho=\rho_o} \rightarrow$ compressibility

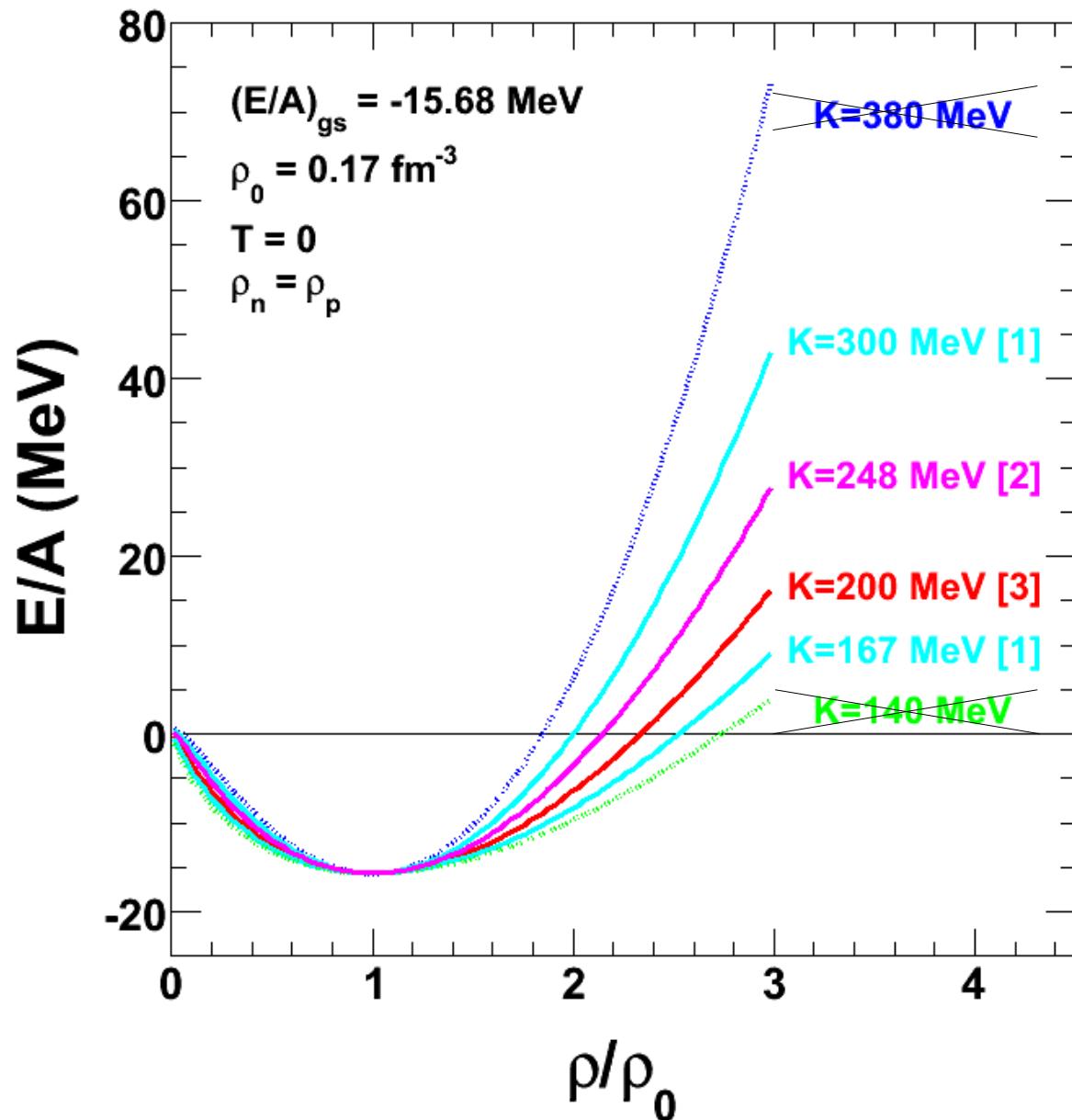
$a_V \approx 15.8 \text{ MeV}$ $a_A^V \approx 32.5^{(*)} \text{ MeV}$ $\delta^2(^{208}\text{Pb}) \approx 0.04$ $\delta^2(^{271}\text{Pb}) \approx 0.16$
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P. Danielewicz, arXiv:nucl-th/0512009

(*) P. Danielewicz, J. Lee, NPA 818(2009)36

Symmetric matter ($\delta=0$)



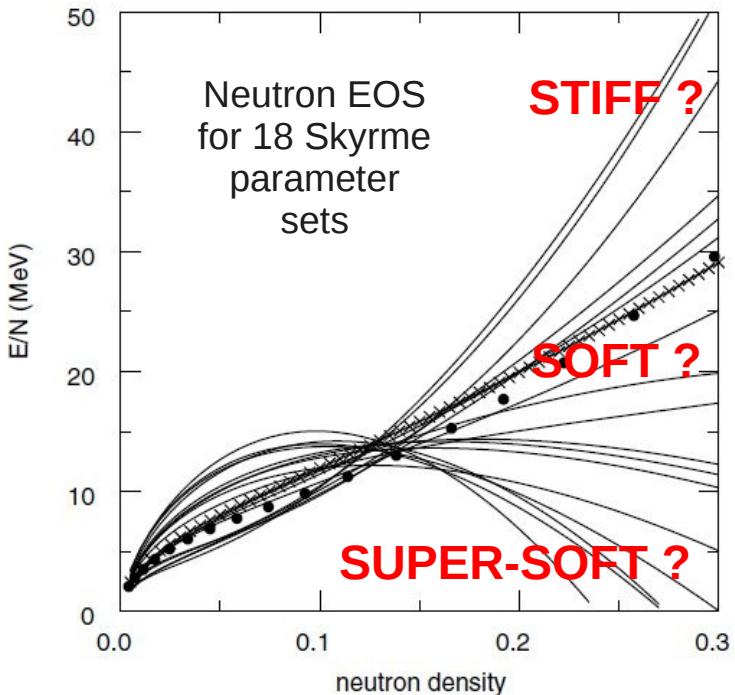
[1] Flow:
P. Danielewicz et al., Science 298 (02) 1592
BME: $K = 167\text{-}300 \text{ MeV}$

[2] ISGMR:
J. Piekarewicz, PRC 69 (04) 041301
RMF: $K=248 \text{ MeV}$
G. Colò et al., PRC 70 (04) 024307
Skyrme HF: $K=230 \text{ MeV}$

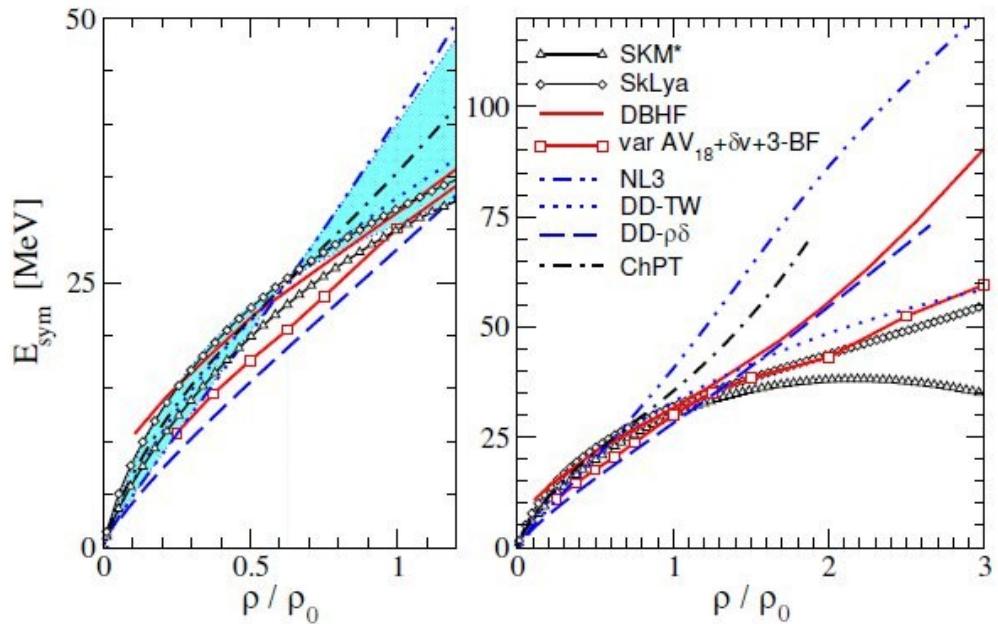
[3] Subthreshold K^+ :
C. Sturm et al., PRL 86 (01) 39
Ch. Hartnack et al., PRL 96 (06) 012302
Flow:
W. Reisdorf (FOPI) arXiv:1307.4210 [nucl-ex]
IQMD: $K=200 \text{ MeV}$

Model and data dependent
values of K for isoscalar part
Still $\sim 20\text{-}30\%$ uncertainty

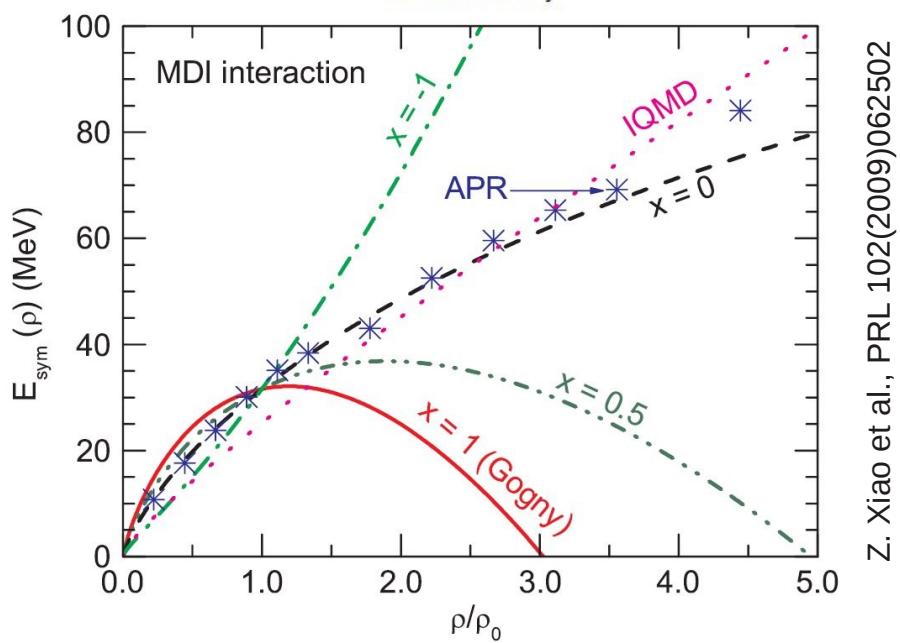
Symmetry term. Why so uncertain?



B. Alex Brown, PRL 85(2000)5299



C. Fuchs and H.H. Wolter, EPJA 30(2006)5

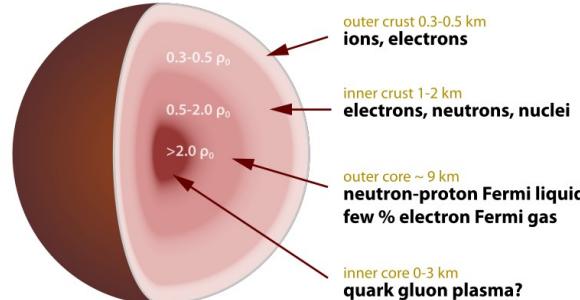


Symmetry energy uncertain at high density and modified by clustering at low density

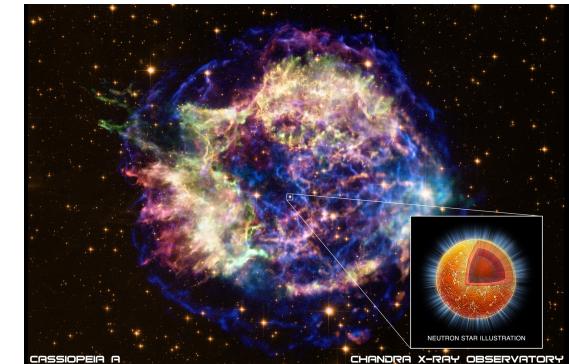
Phenomenological forces constrained around saturation and for nearly isospin-symmetric matter. Poor knowledge of effective forces in neutron-rich matter. Uncertainties in the nature of the three-neutron force. Uncertain extrapolations above the saturation density.

Why so important?

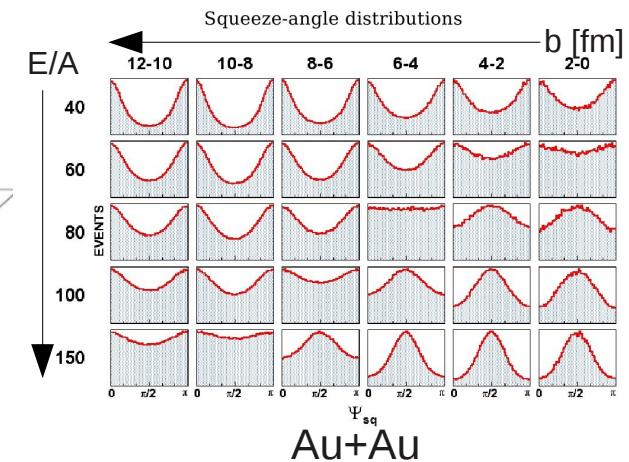
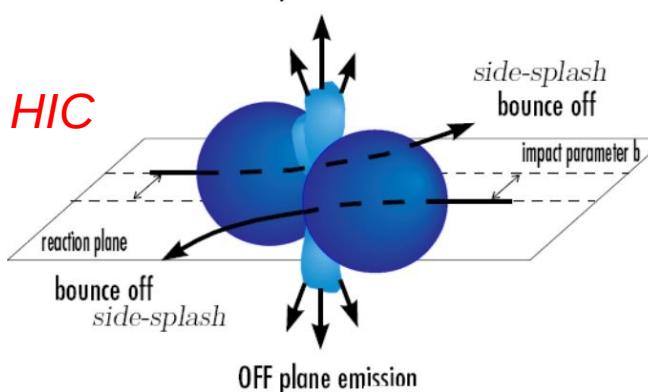
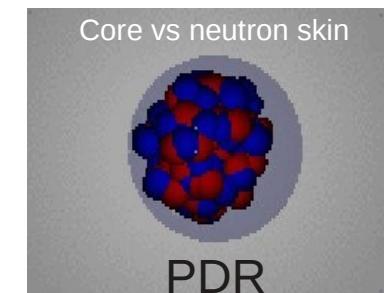
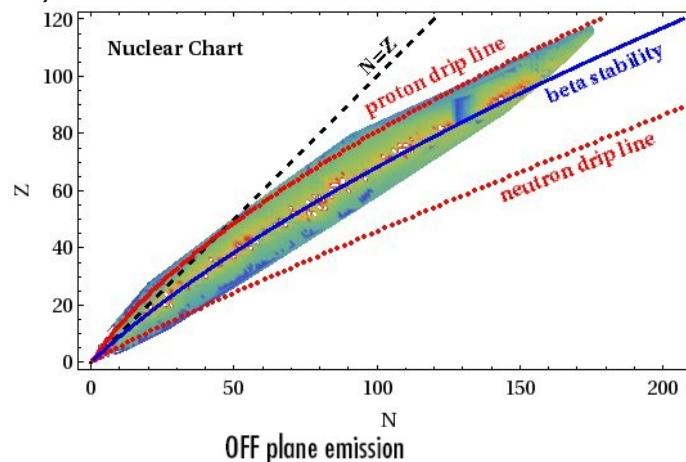
- Neutron star structure, composition, size, mass and cooling rate
- Supernova explosions
- Structure of the nucleus (masses, drip lines)
- Neutron skin thickness
- IvGDR
- Pygmy resonances
- Differences between IAS
- *Flow patterns in HIC*
- *n/p, t/³He, π/π⁺, K⁺/K⁰ ratios in HIC*



Robert Schulze (Wikipedia)

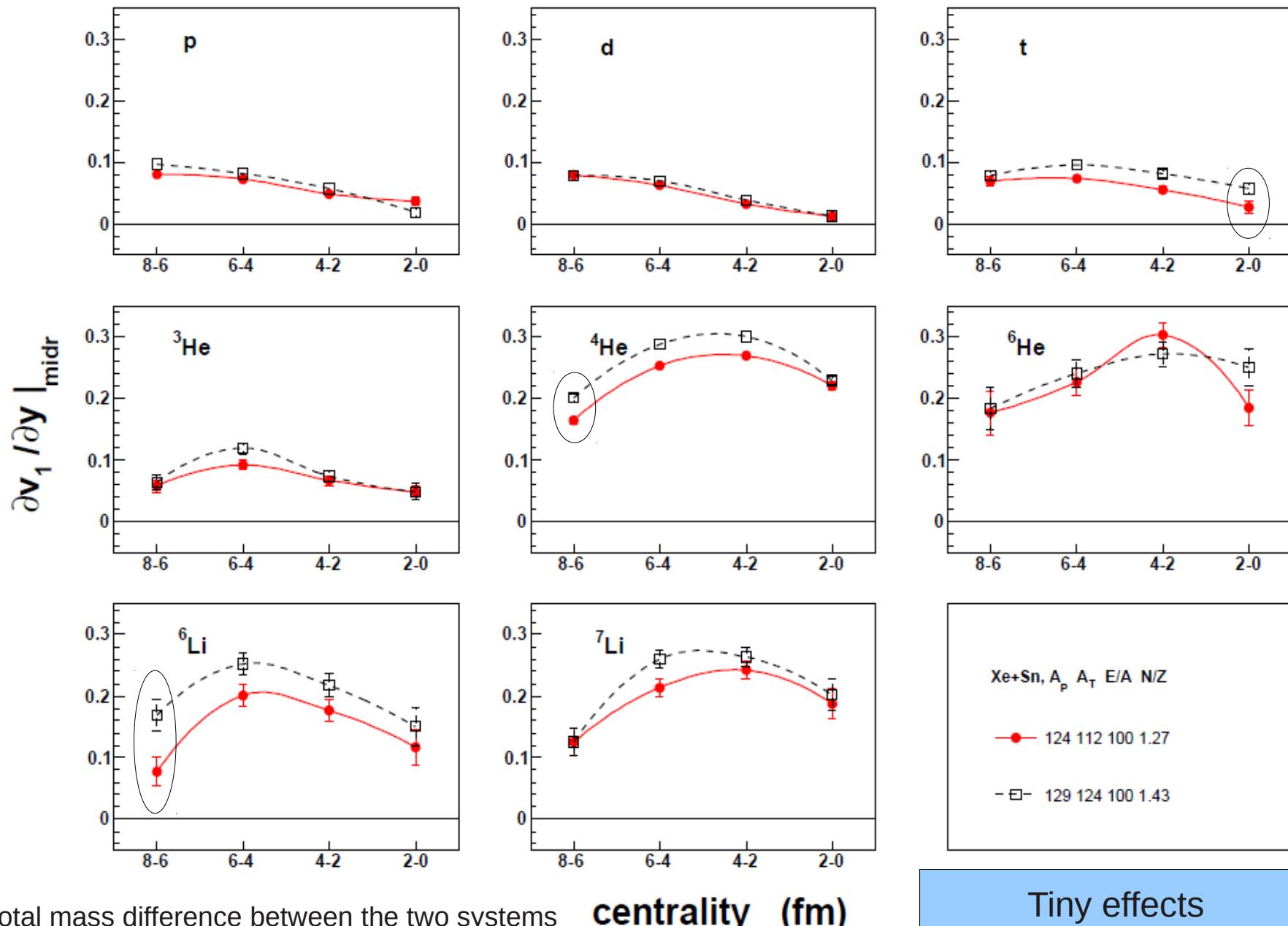


CASSIOPEIA A
CHANDRA X-RAY OBSERVATORY
Cassiopeia A Supernova Remnant
(http://chandra.harvard.edu/photo/printgallery/2004/)



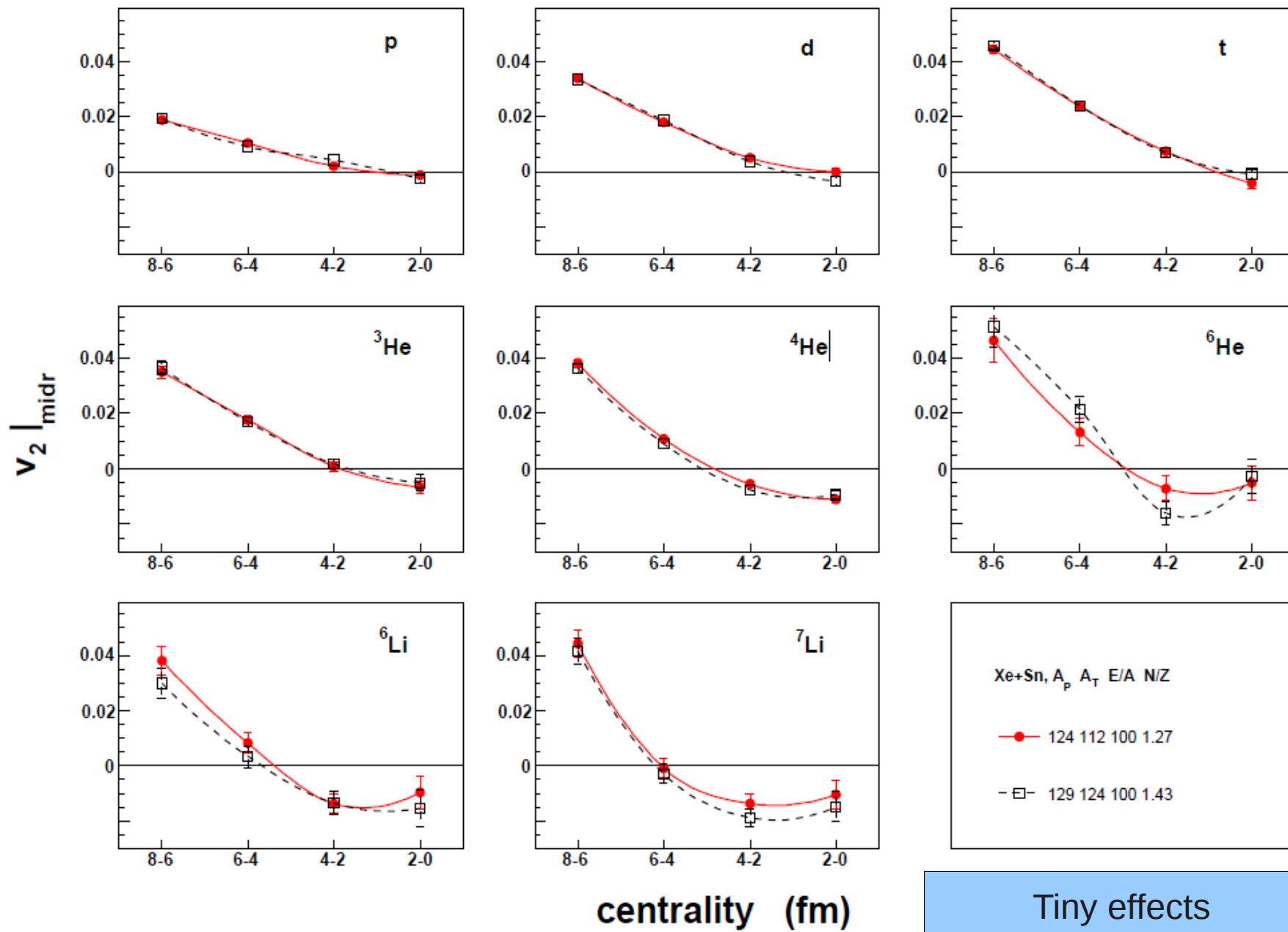
INDRA@GSI: $^{124,129}\text{Xe} + ^{112,124}\text{Sn}$ @ 100 AMeV

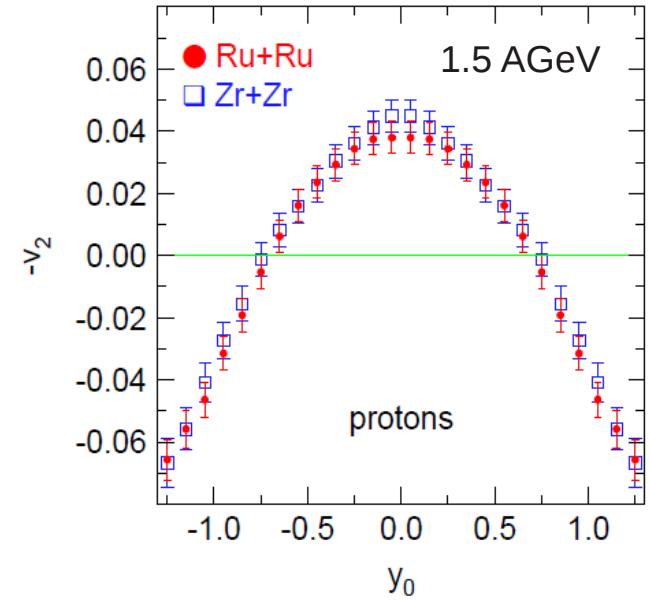
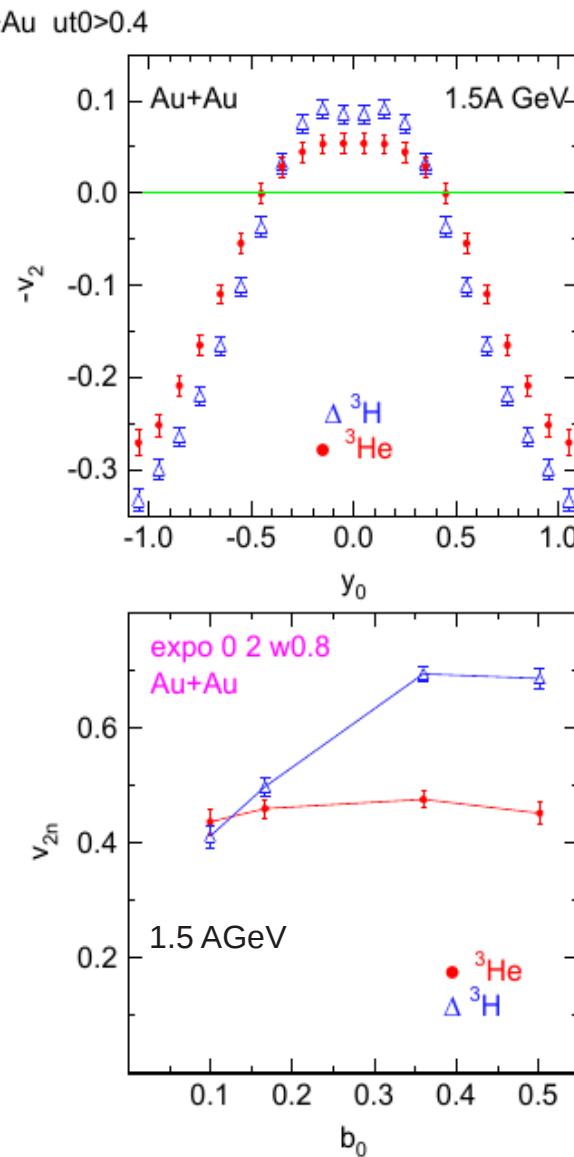
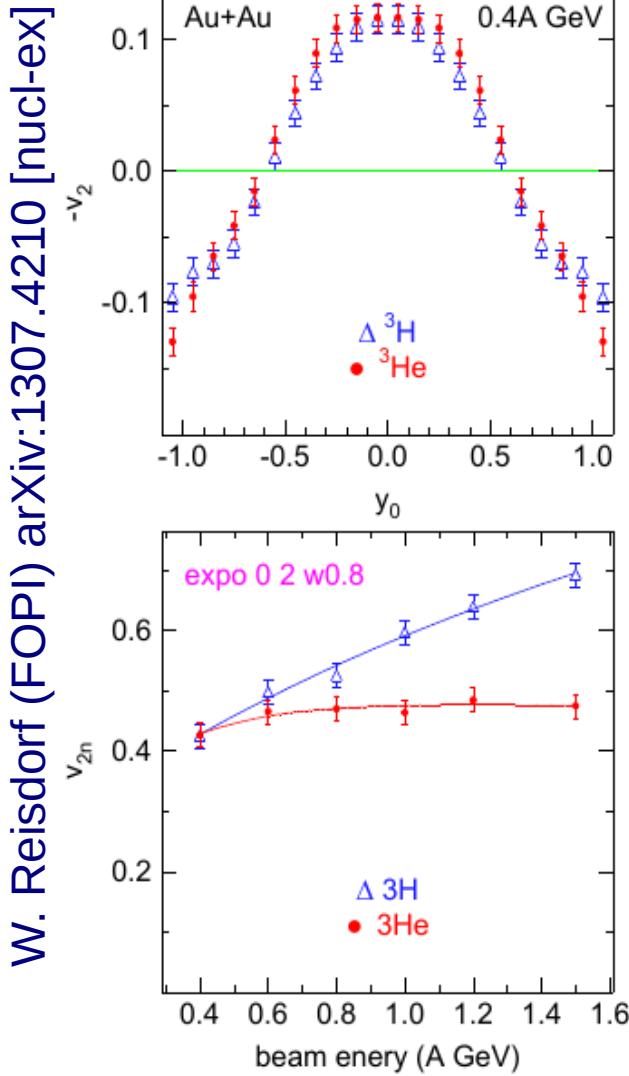
slope of v_1 at midrapidity



INDRA@GSI: $^{124,129}\text{Xe} + ^{112,124}\text{Sn}$ @ 100 AMeV

v2 at midrapidity



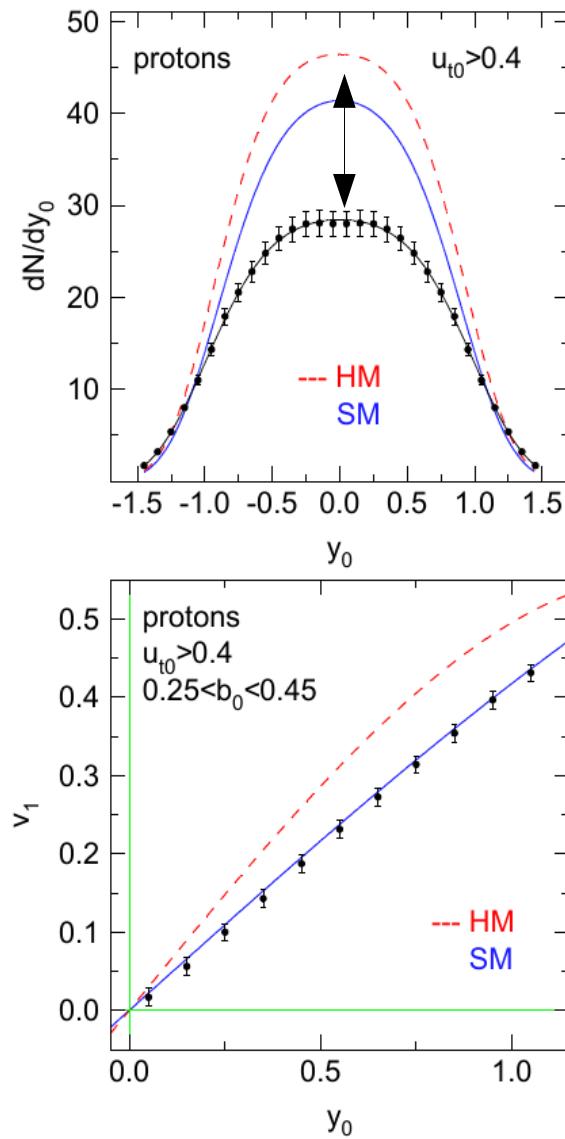


Increasing sensitivity of cluster isospin probes with
 ➤ beam energy
 ➤ impact parameter
 ➤ mass of the system
 ➤ mass of the probe

Reliable cluster production algorithms needed

v_{2n} parametrizes the whole v_2 rapidity distribution

FOPI + IQMD



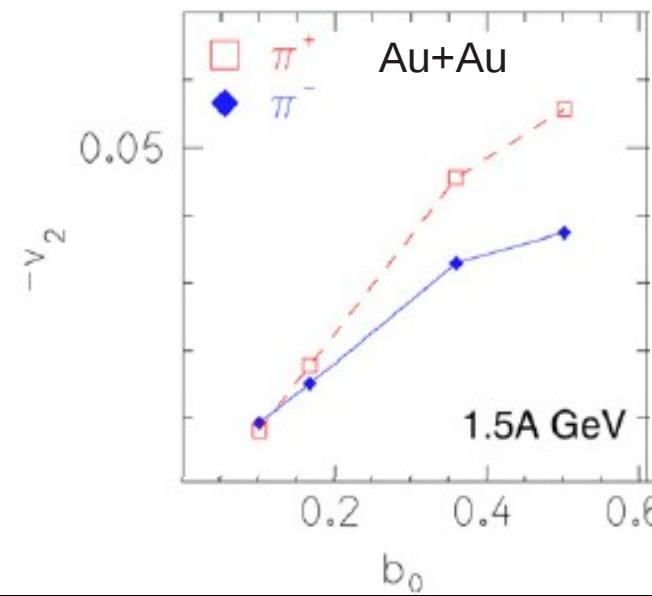
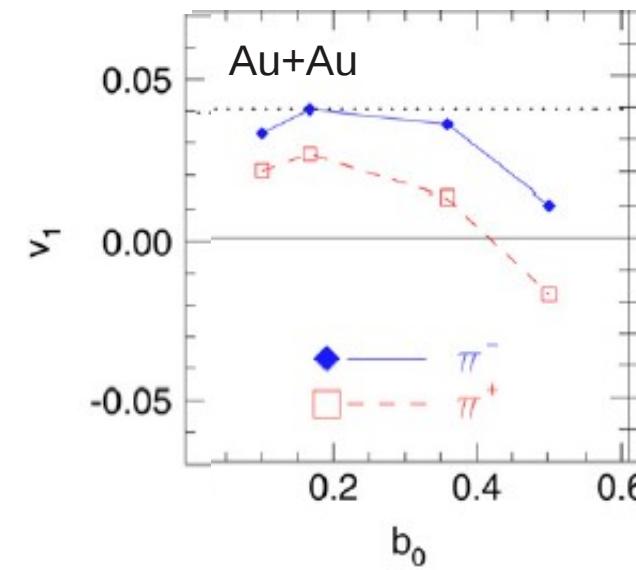
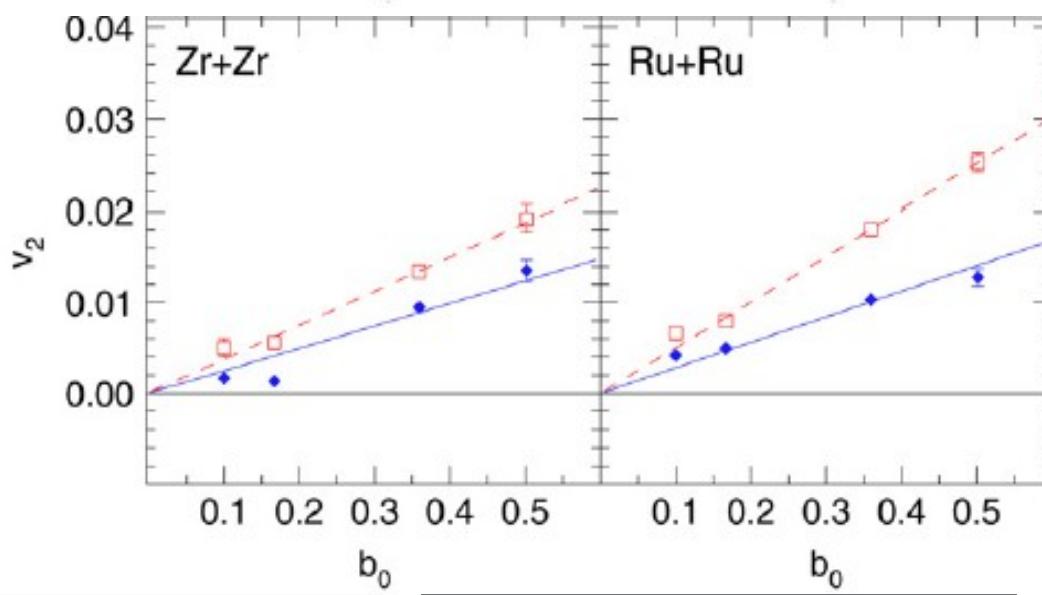
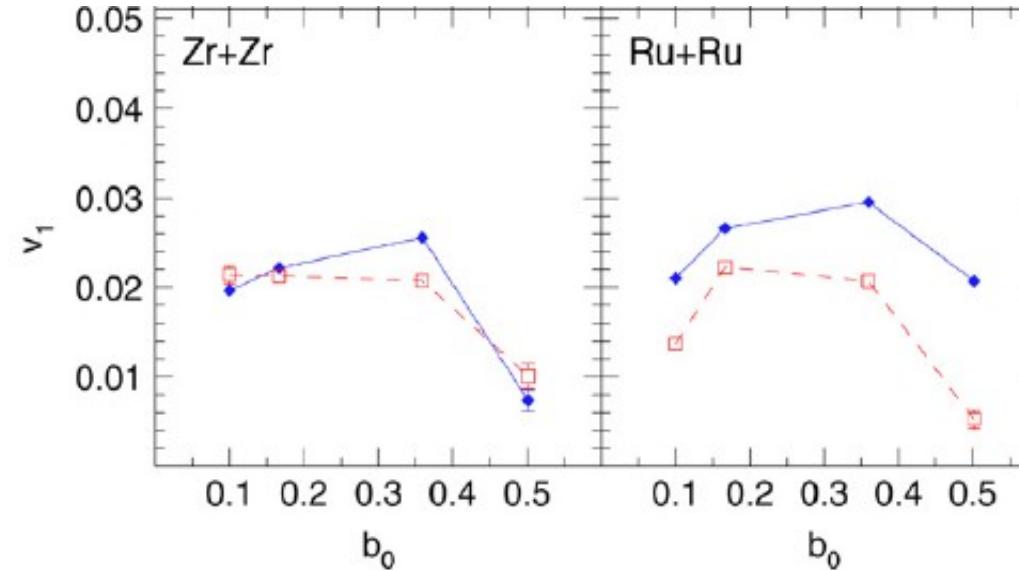
Au+Au @ 1 AGeV,
Scaled rapidity distribution

Reliable cluster production algorithms
needed

Au+Au @ 1 AGeV,
 v_1 vs scaled rapidity

HM, SM – HARD and SOFT isoscalar EOS
parametrizations with momentum dependence

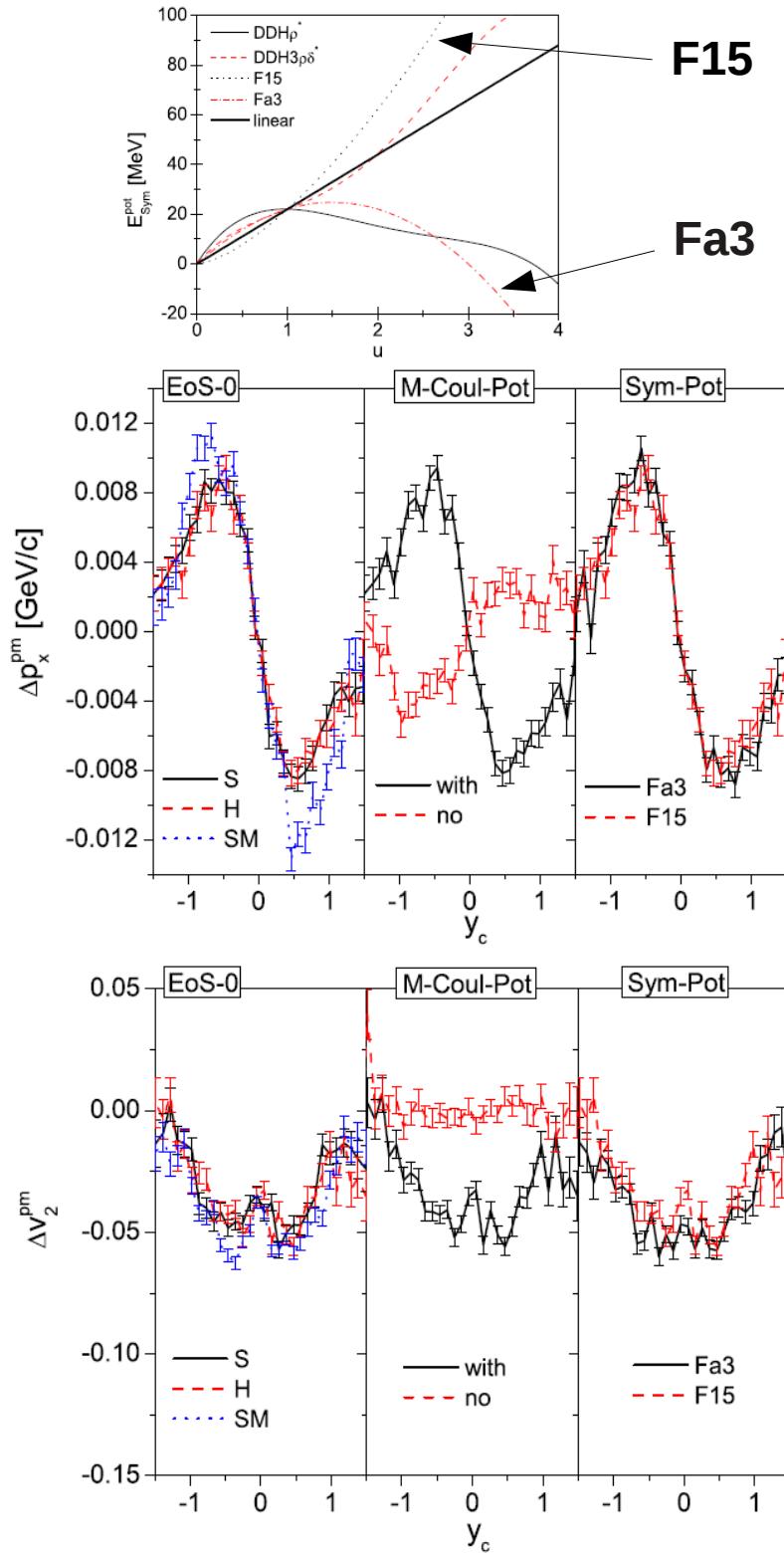
Pion flows, FOPI, 1.5 AGeV



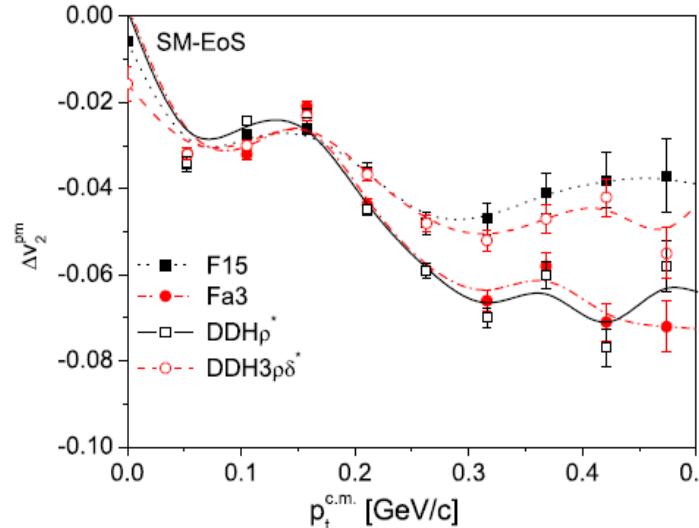
	N/Z	δ^2
⁹⁶ Zr	1.40	0.028
⁹⁶ Ru	1.18	0.007
¹⁹⁷ Au	1.49	0.039

Coulomb + shadowing + Symmetry energy effects

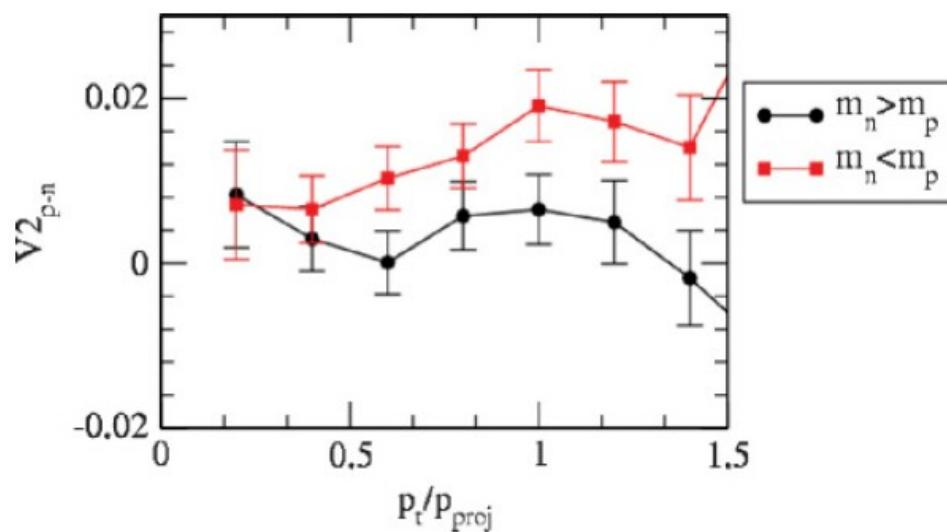
Q-F Li et al. JPG 32(2006) 151 (UrQMD): pion v_1 and v_2 depend weakly on E_{sym} but are sensitive to the momentum dependence of the meanfield. But, at high pT , v_2 should discriminate between E_{sym}



Q-F Li et al. JPG 32(2006) 151 (UrQMD)
208Pb + 208 Pb reaction at $E_b = 0.8 \text{ A GeV}$ and $b = 7-9 \text{ fm}$

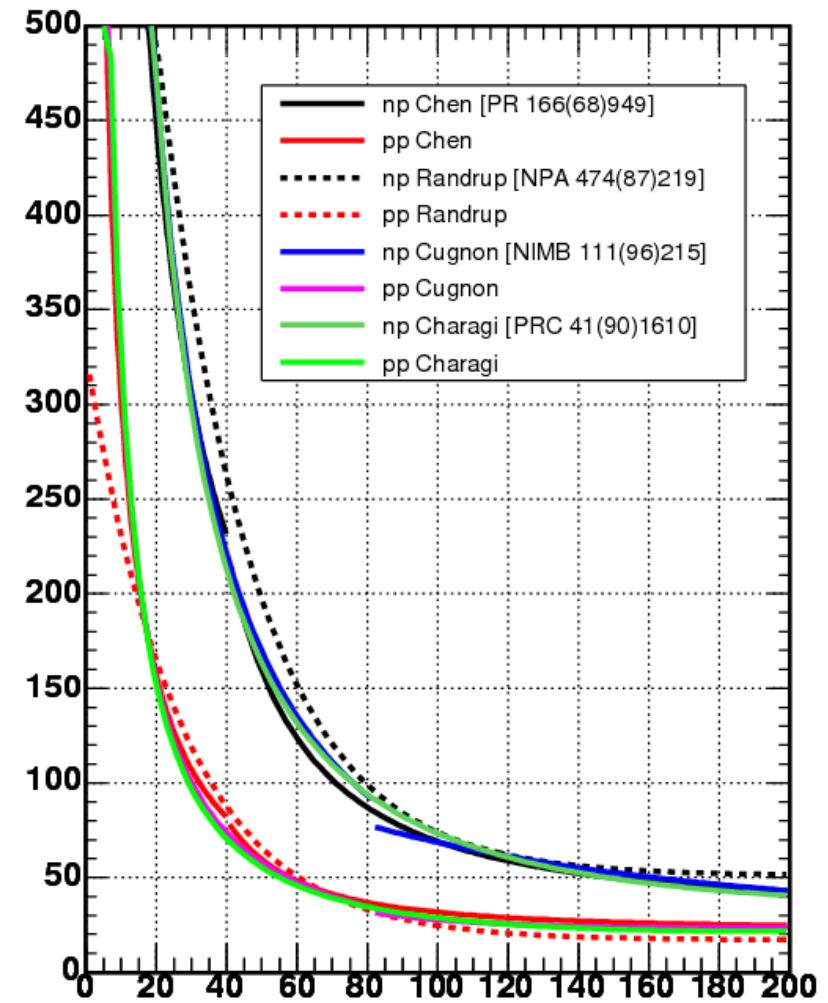
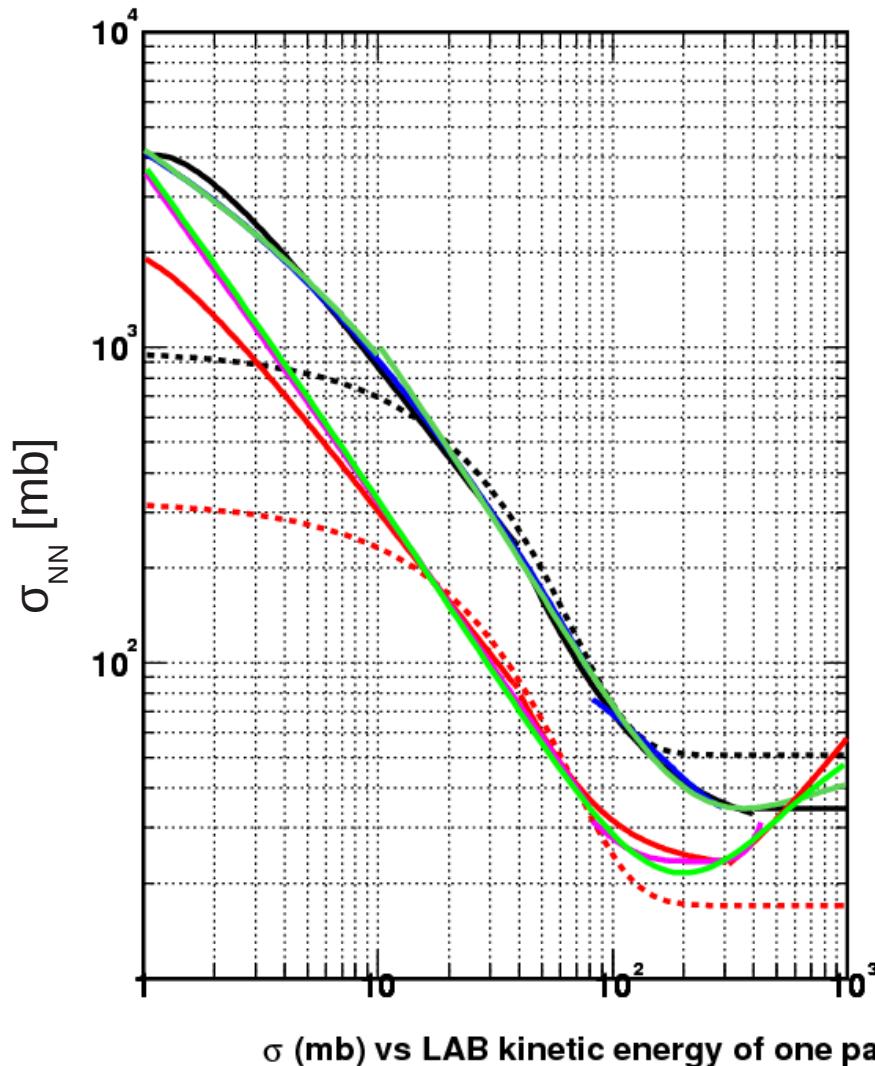


Giordano et al. PRC 81(2010)044611 (SMF)
Au+Au @ 400 AMeV, ASY-STIFF



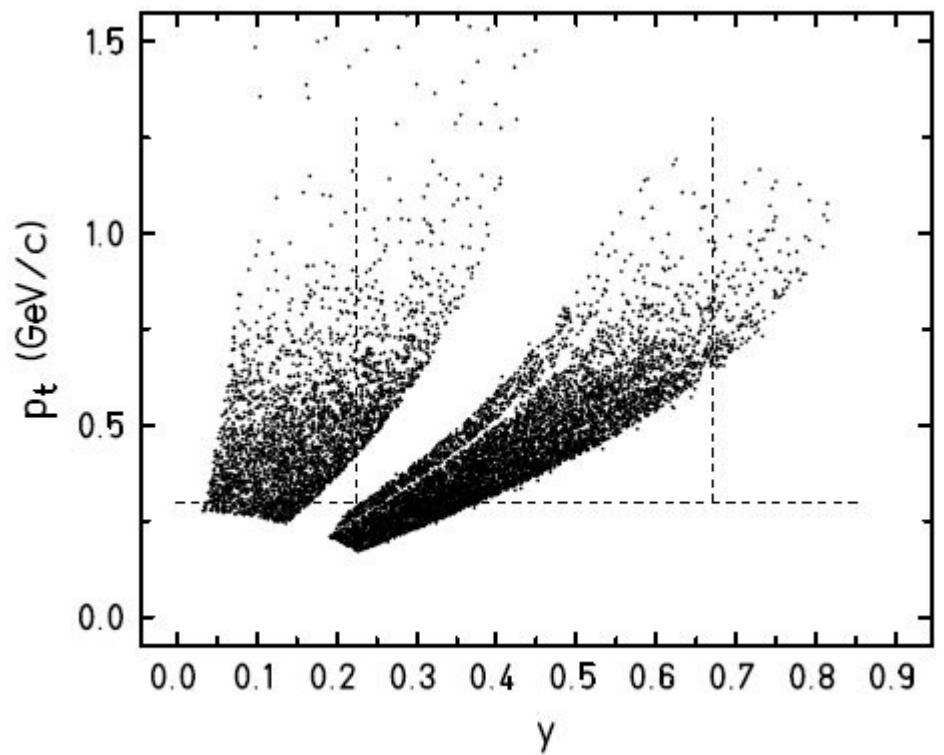
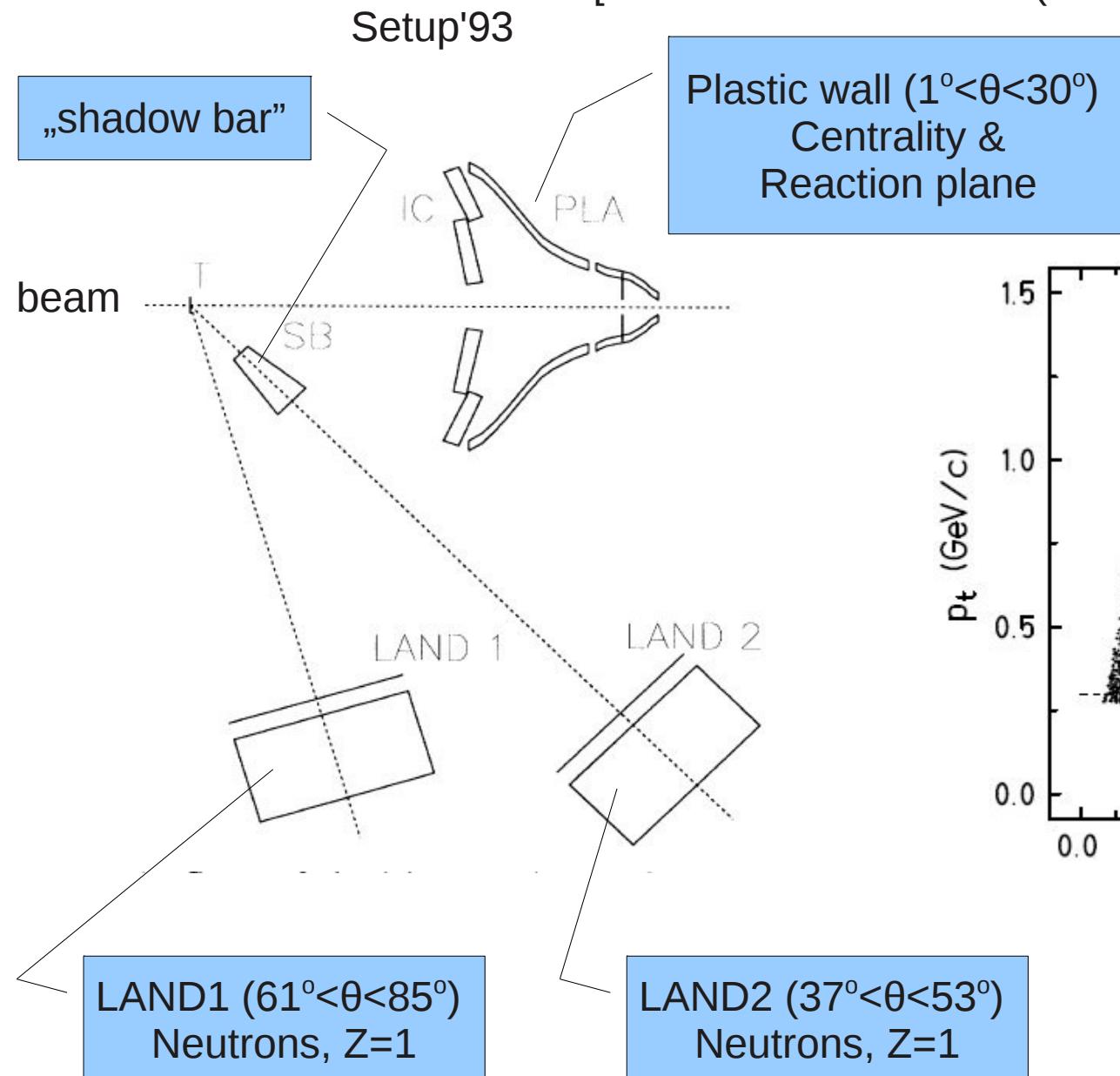
It is possible that at high p_T one may probe the momentum dependence of the E_{sym} instead of its density dependence.
Some effects may be due to the σ_{NN} isospin dependence

Free σ_{np} and σ_{pp} as a function of energy



Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

[Y. Leifels et al. PRL 71 (1993) 963]



Acceptance for neutrons
and cuts for analysis (dashed)

Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

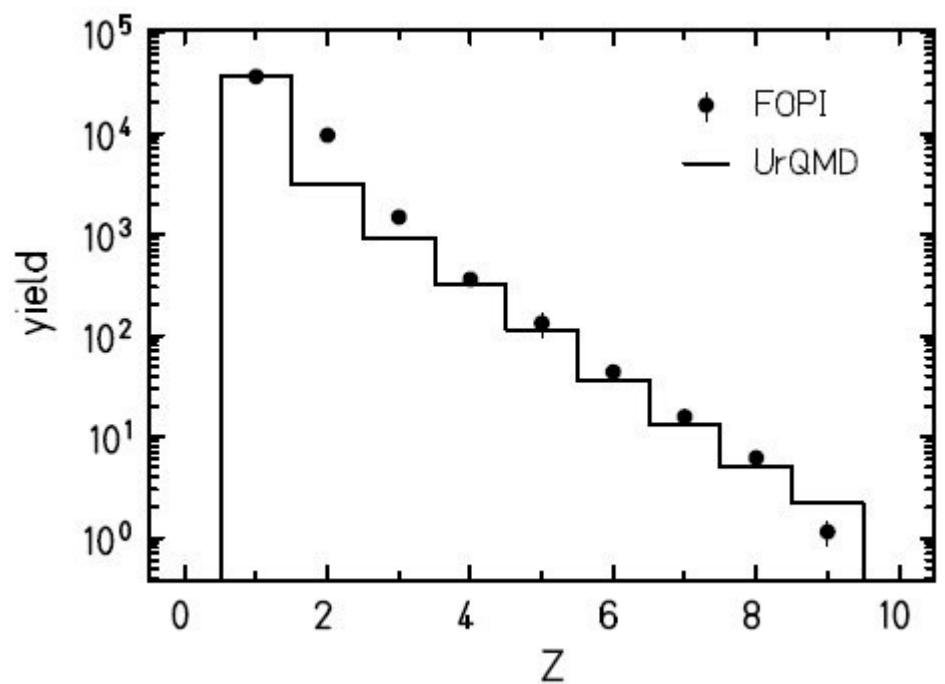
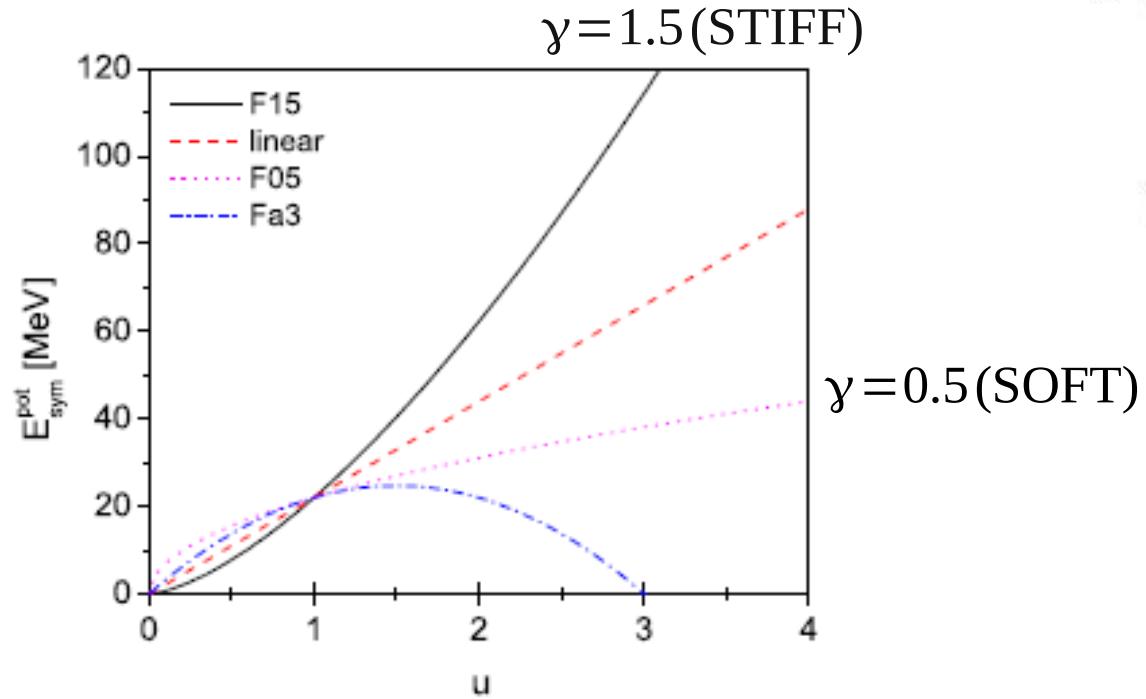
[Y. Leifels et al. PRL 71 (1993) 963] -> [P. Russotto et al. PLB 697 (2011) 471]

UrQMD, Q. Li, J.Phys. G 31(2005)1359

P. Russotto et al. PLB 697 (2011) 471
Data: W. Reisdorf, et al., NPA 612 (1997) 493
Central collisions, Au+Au @ 400 AMeV

„Fermi-gas“ parametrization of the symmetry term:

$$E_{sym} = E_{sym}^{pot} + E_{sym}^{kin} = 22 \text{ MeV} \left(\frac{\rho}{\rho_o} \right)^{\gamma} + 12 \text{ MeV} \left(\frac{\rho}{\rho_o} \right)^{2/3}$$

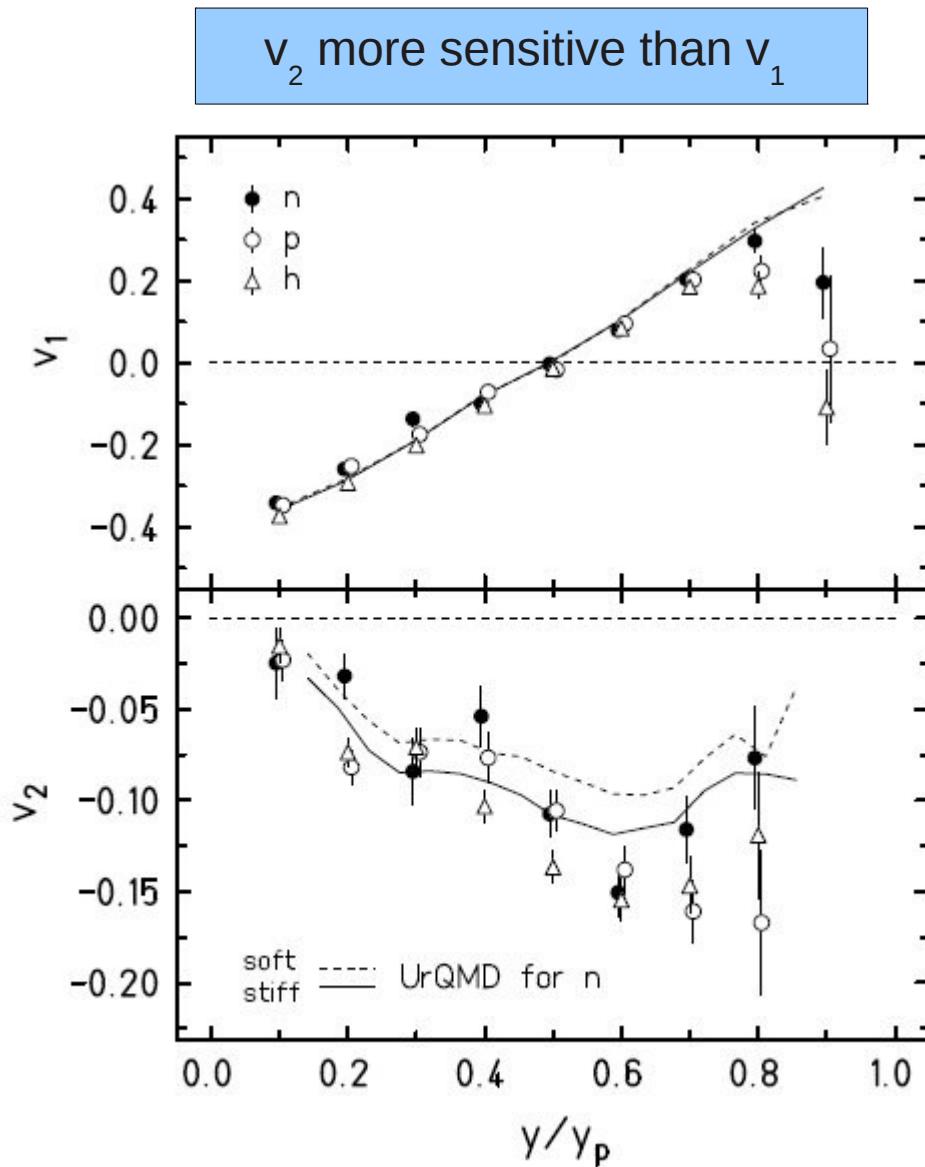
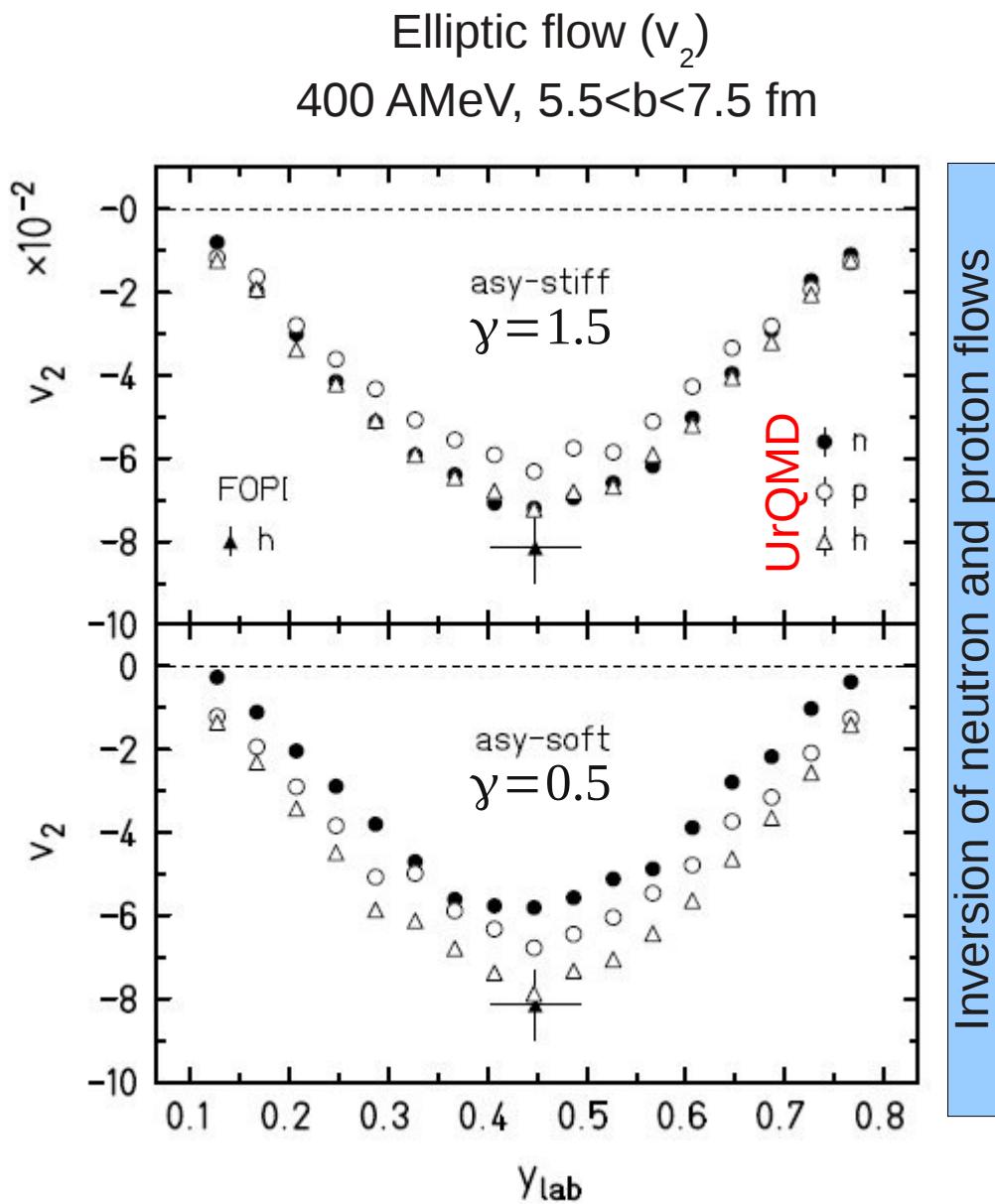


Fragment yields
(test of clustering procedure
with $\Delta r=3$ fm and $\Delta p=275$ MeV/c)
Normalization at $Z=1$

d,t, α underpredicted (x 2-3)

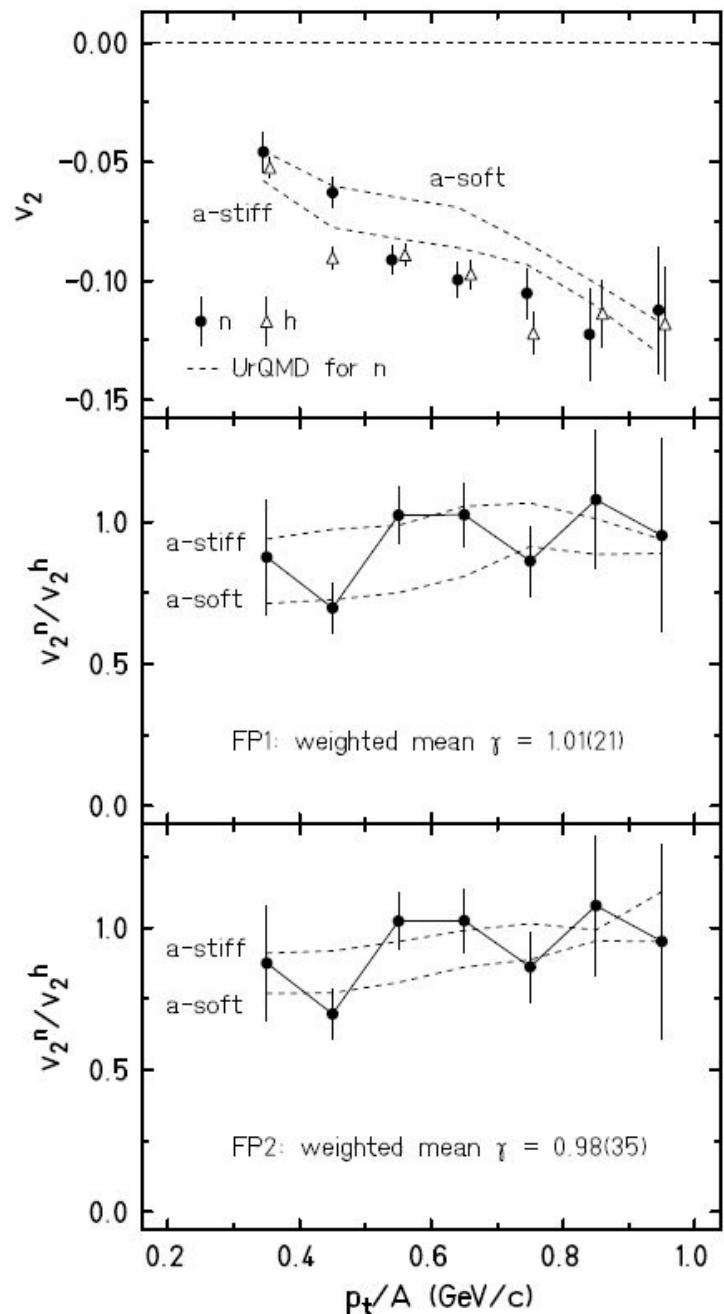
Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

[P. Russotto et al. PLB 697 (2011) 471]



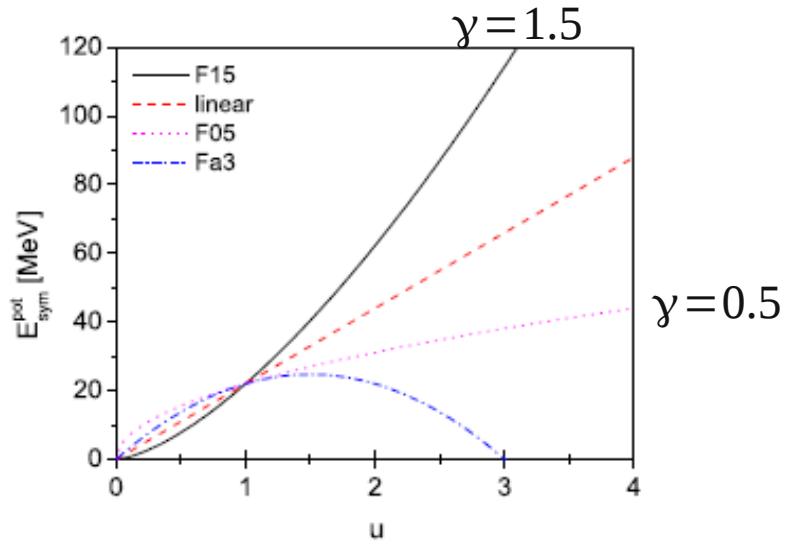
Reanalysis of FOPI/LAND Au+Au @ 400, 600, 800 AMeV

[P. Russotto et al. PLB 697 (2011) 471]



v_2^n / v_2^H ratio vs p_T sensitive to the symmetry term
and not so much to the momentum dependence of the σ_{NN}

neutron/hydrogen
FP1: $\gamma = 1.01 \pm 0.21$
FP2: $\gamma = 0.98 \pm 0.35$
neutron/proton
FP1: $\gamma = 0.99 \pm 0.28$
FP2: $\gamma = 0.85 \pm 0.47$
adopted: $\gamma = 0.9 \pm 0.4$

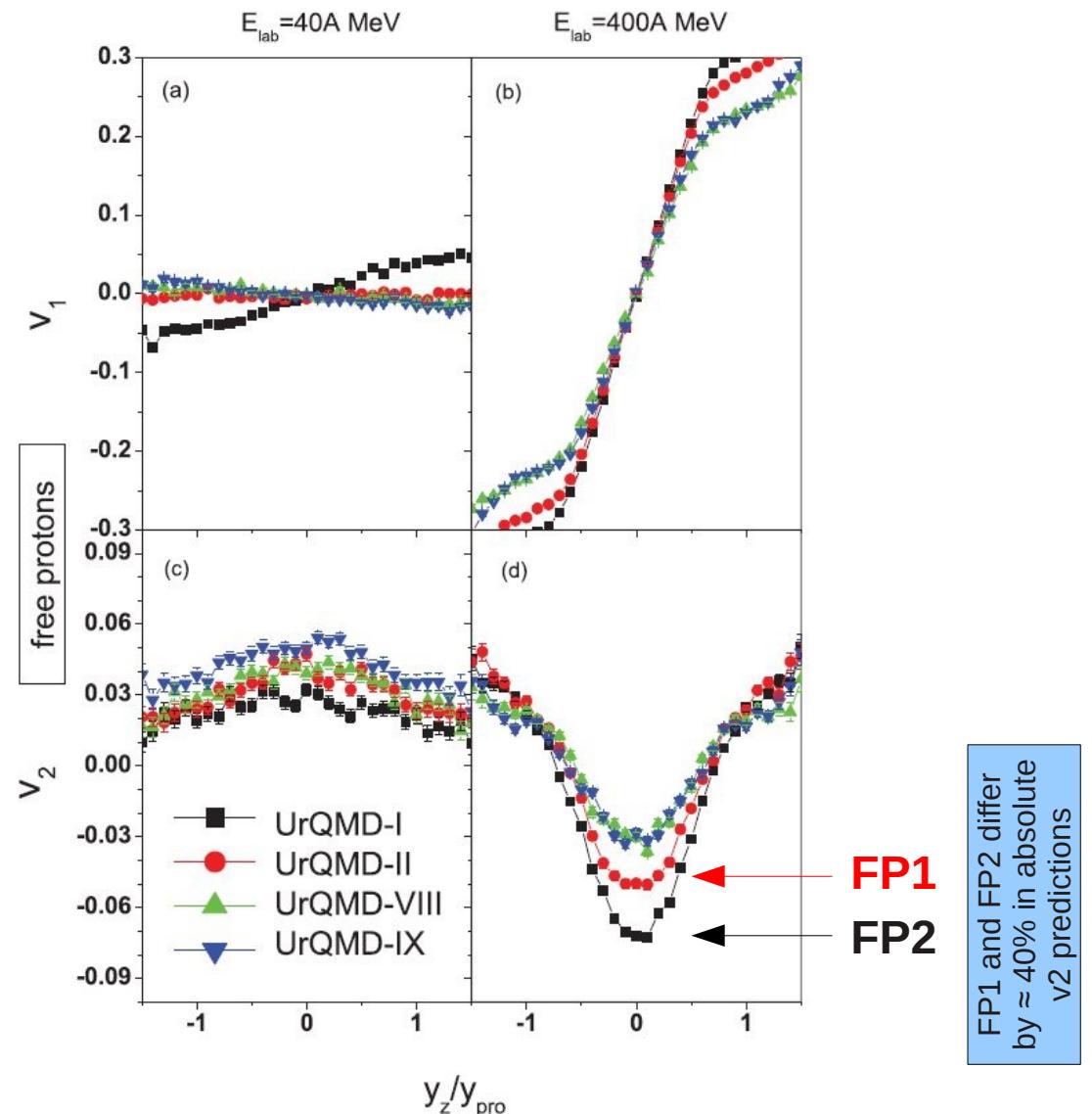
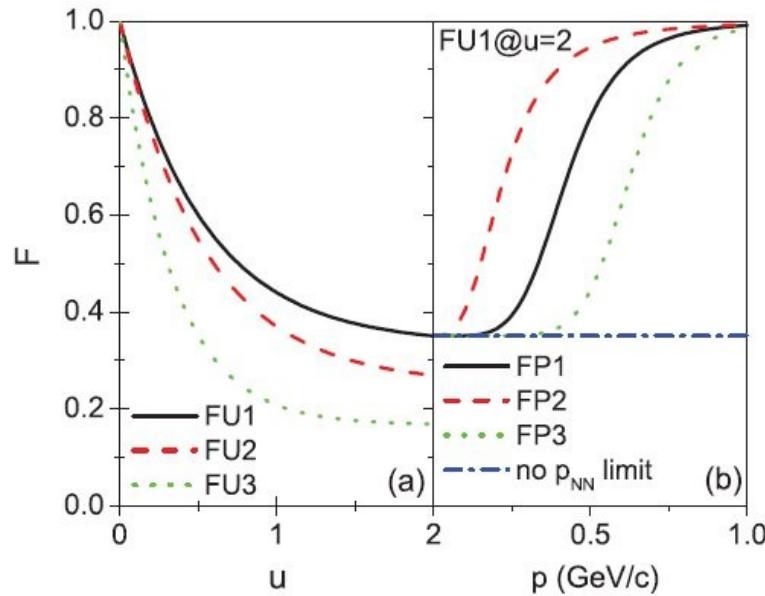


FP1, FP2: different parametrizations of the
momentum dependence of the elastic
nucleon-nucleon cross section
[Q. Li et al. PRC 83(2011)044617]

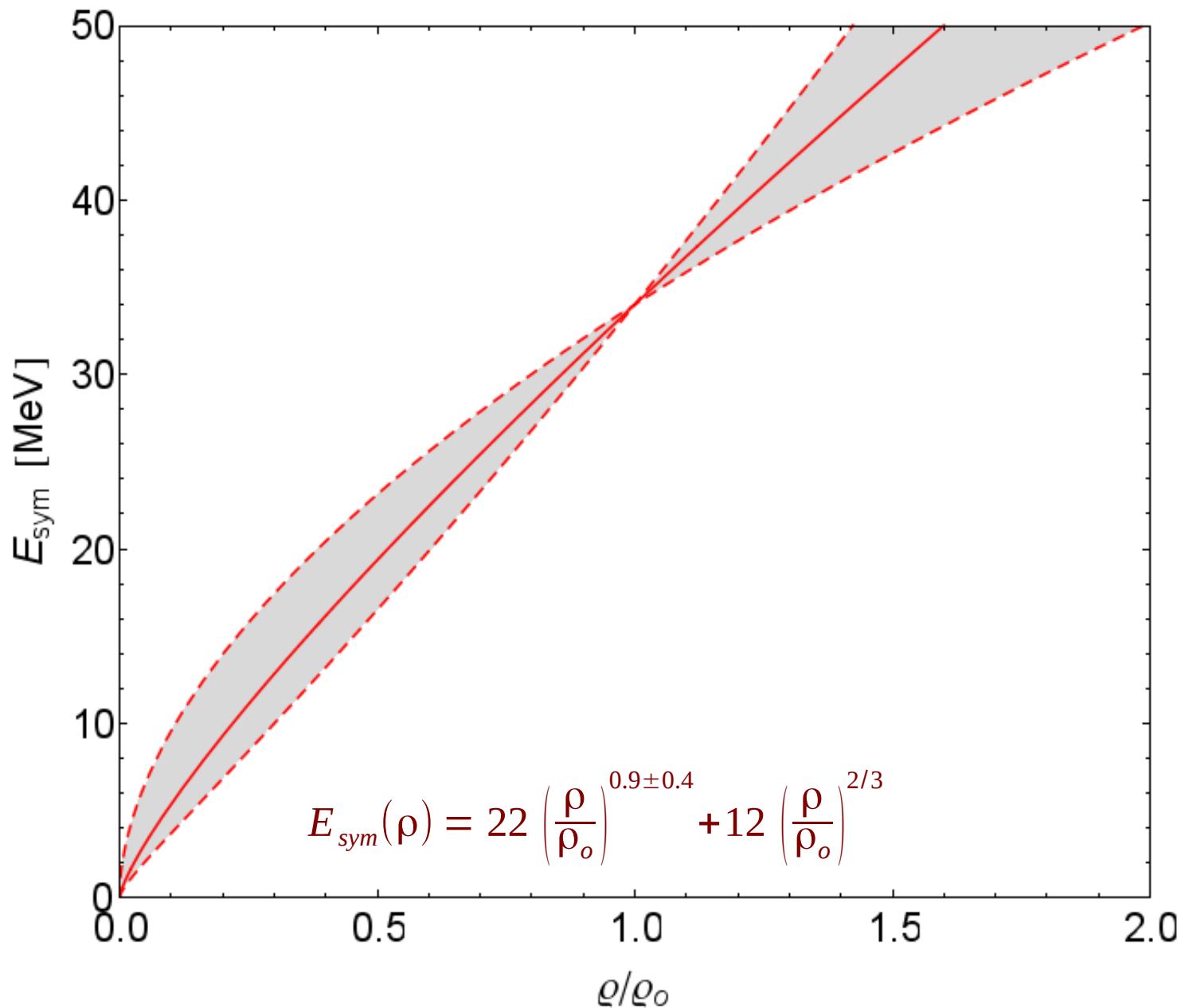
Medium correction factors to the elastic σ_{NN}

[UrQMD, Q. Li et al., PRC 83(2011)044616]

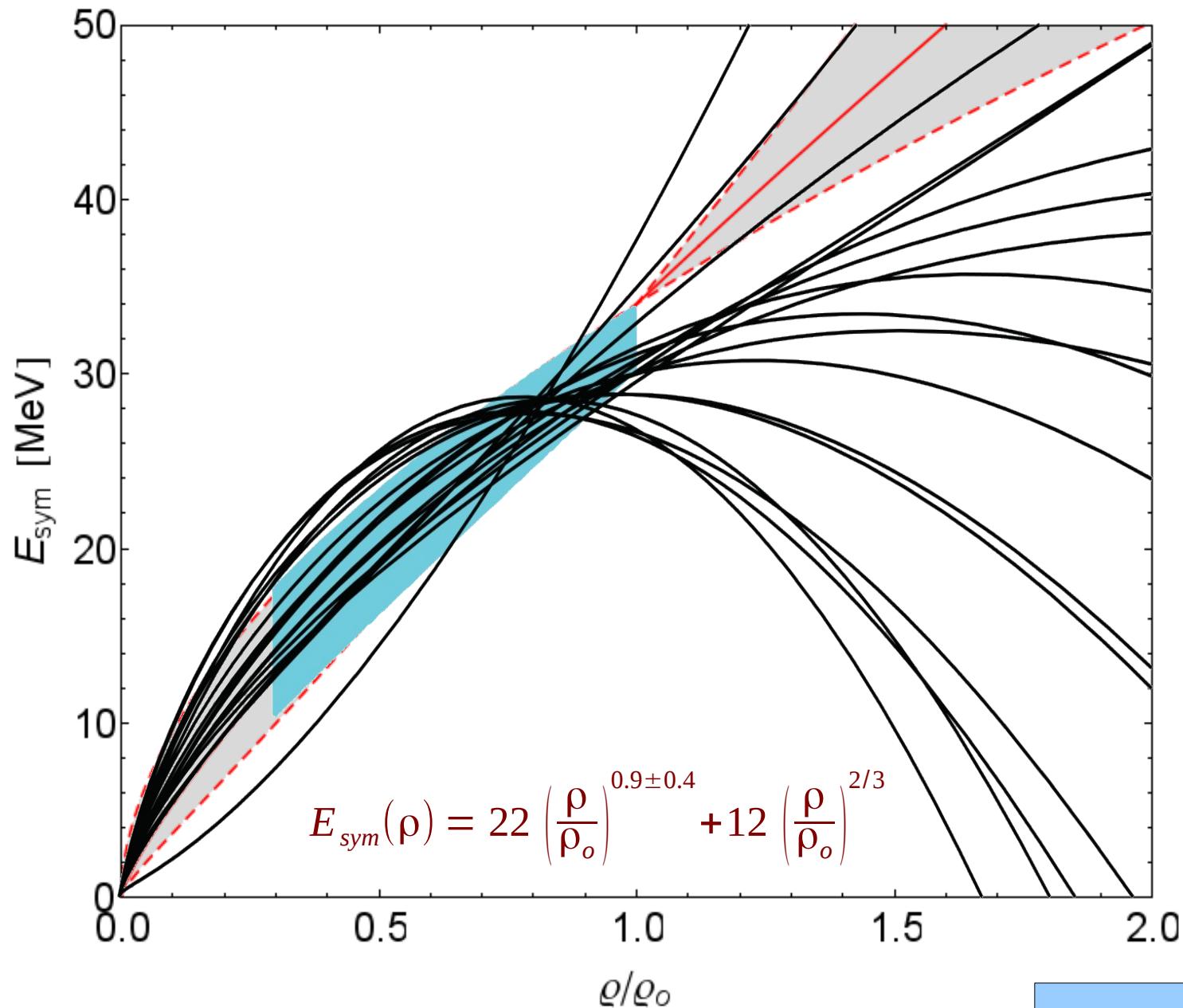
$$\sigma_{\text{tot}}^* = \sigma_{\text{in}} + \sigma_{\text{el}}^* = \sigma_{\text{in}}^{\text{free}} + F(\rho, p) \sigma_{\text{el}}^{\text{free}}$$



Esym from FOPI + UrQMD



E_{sym} from FOPI + UrQMD + Tsang(2012)



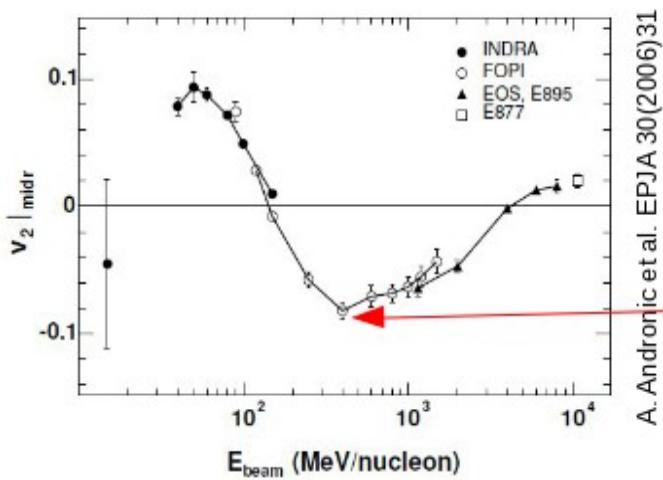
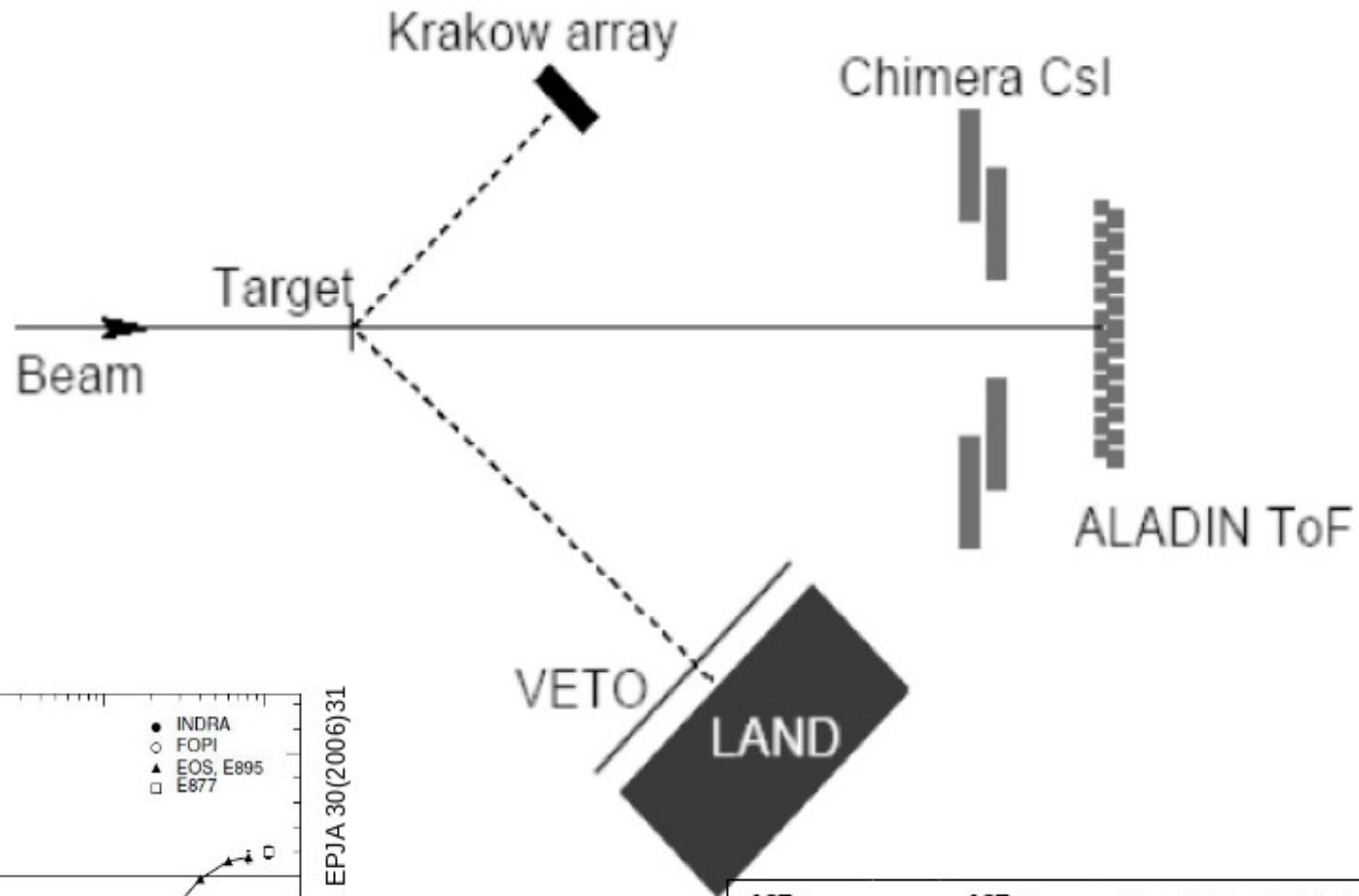
P. Russotto et al., PLB 697 (2011) 471
M.B. Tsang et al., PRC 86(2012)015803

See D. Cozma

ASY-EOS experimental setup

May 2011

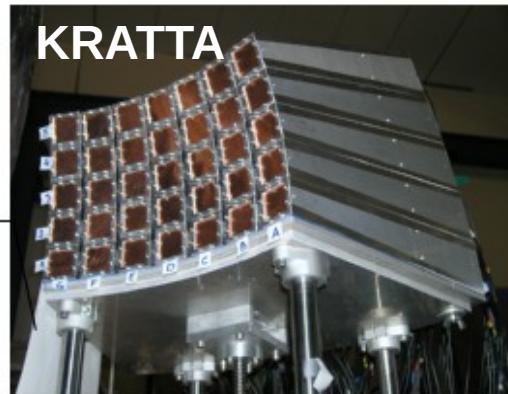
Setup from the proposal of 2009



$^{197}\text{Au} + ^{197}\text{Au}$ @ 400 AMeV	$\delta^2 = 0.039$
$^{96}\text{Zr} + ^{96}\text{Zr}$ @ 400 AMeV	$\delta^2 = 0.028$
$^{96}\text{Ru} + ^{96}\text{Ru}$ @ 400 AMeV	$\delta^2 = 0.007$

ASY-EOS

5x7 triple telescopes, $20^\circ < \theta < 60^\circ$
 Si-CsI-CsI
 Midrapidity pdt +
 Isotopes of $Z < 9$

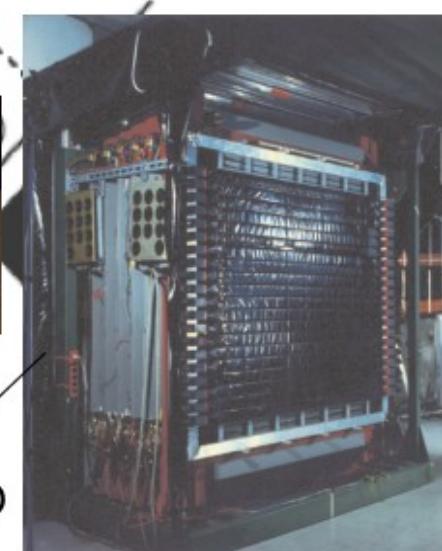


Start + ROLU



μ -Ball + Halo

4 rings, $\theta > 60^\circ$, CsI(Tl)
 Discriminate target vs air
 interactions, remove halo,
 possibly centrality + rpl



ALADIN ToF

2x2x1 m³ plastic/Fe sandwich
 + plastic veto wall
 Midrapidity neutrons & $Z=1$

LAND+VETO

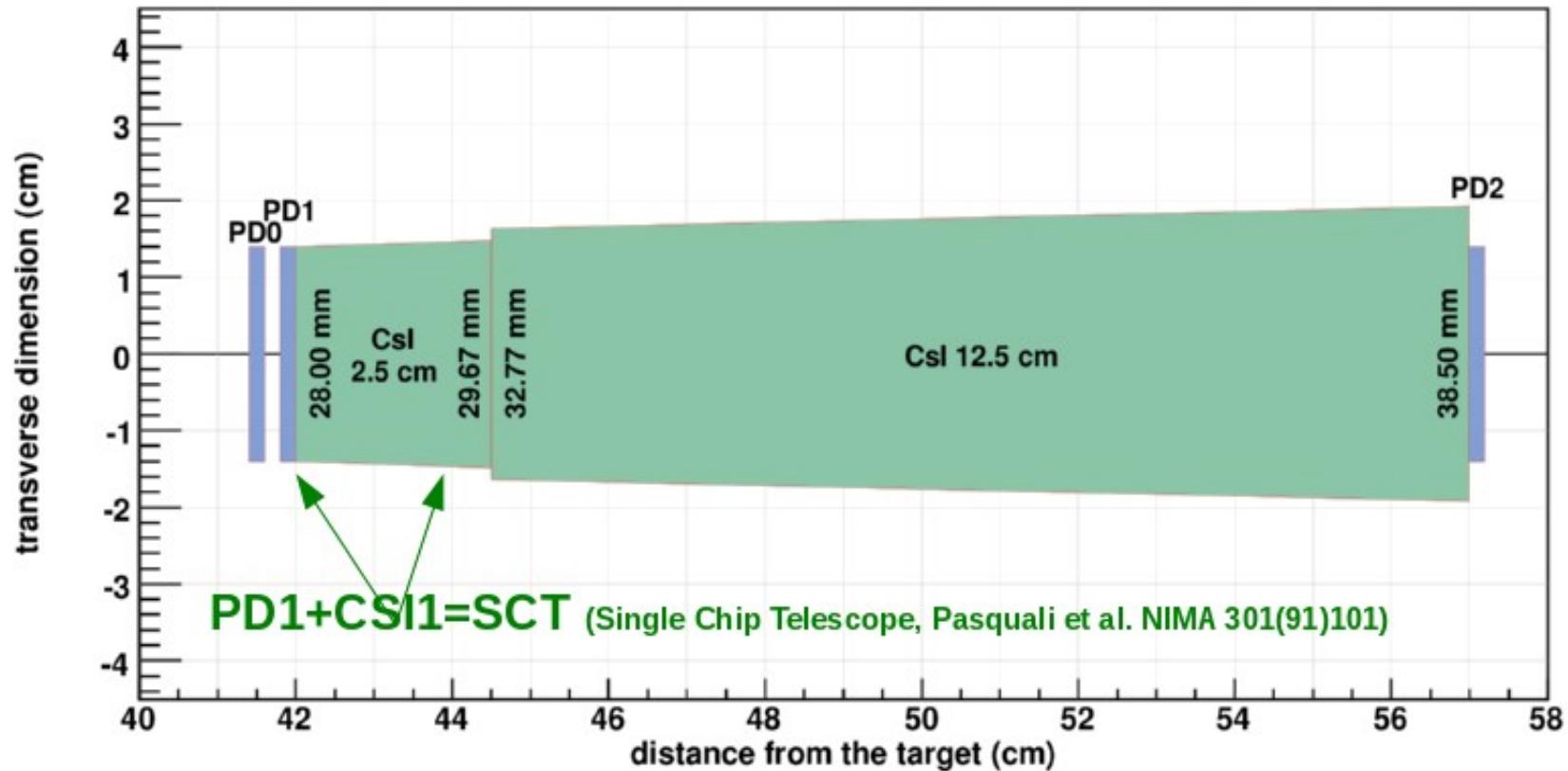
setup

4 rings, 352 CsI(Tl), $7^\circ < \theta < 20^\circ$
 Centrality
 &
 Reaction plane



96 plastic bars
 x-y positions, centrality,
 reaction plane, trigger

Active elements



Photodiodes: HAMAMATSU S5377-02

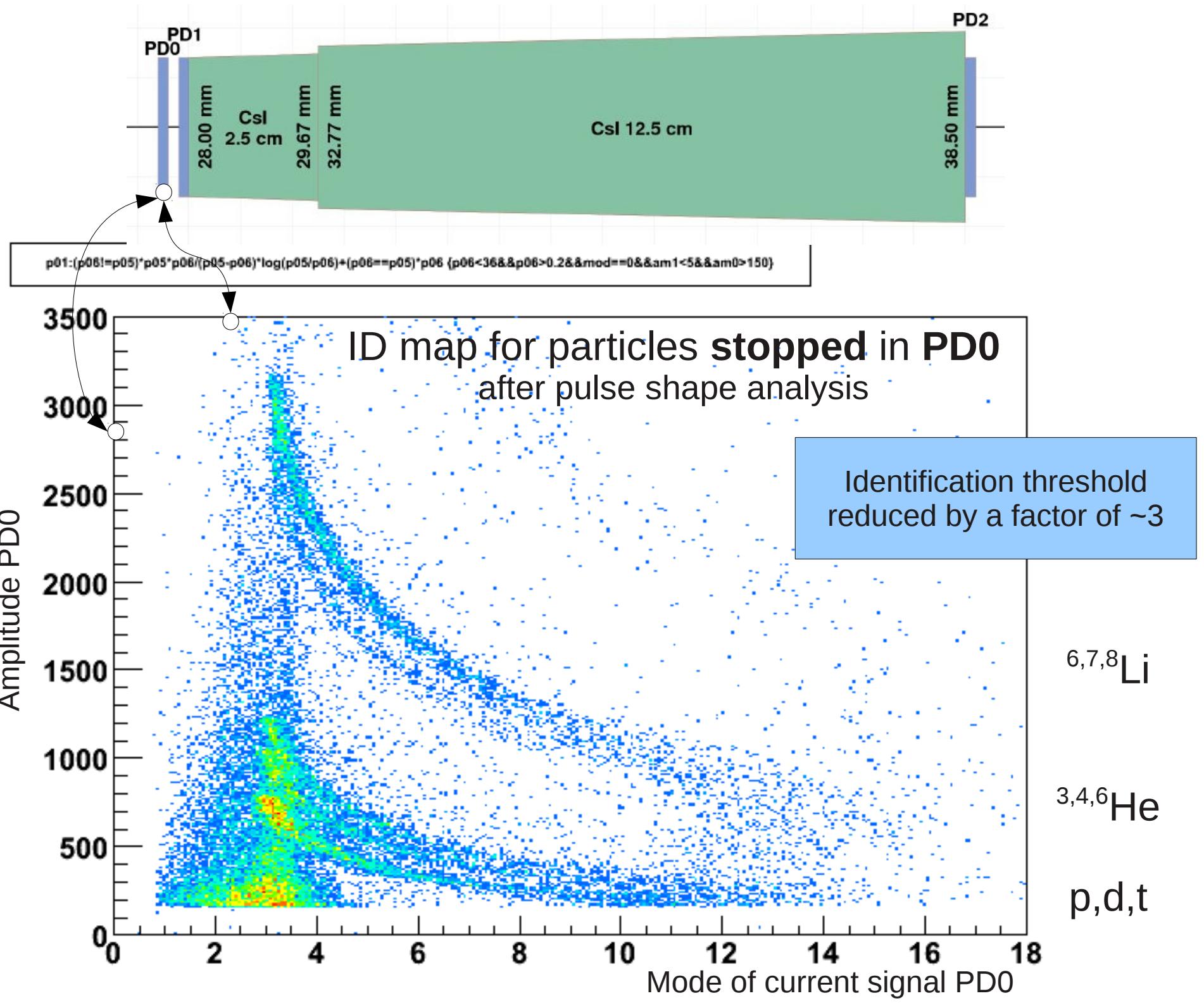
- Active Area: 28x28 mm²
- Thickness: 500 ± 15 µm
- Orientation: (111)
- Dead Layers: 1.5 µm front, 20 µm rear
- Full Depletion: ~170 V
- Dark Current: 30 nA, (Max. 150 nA)
- Rise Time: 40 ns
- Capacitance: 200 pF

CsI(Tl): IMP-CAS, Lanzhou, China

- Tl concentration: 1500 ppm
- LO non-uniformity: <7%
- Shape: Truncated pyramids
- Tolerance: ± 0.1 mm

Wrapping: 3M Vikuiti™ ESR foil

- Reflectance: >98%
- Thickness: 65 µm

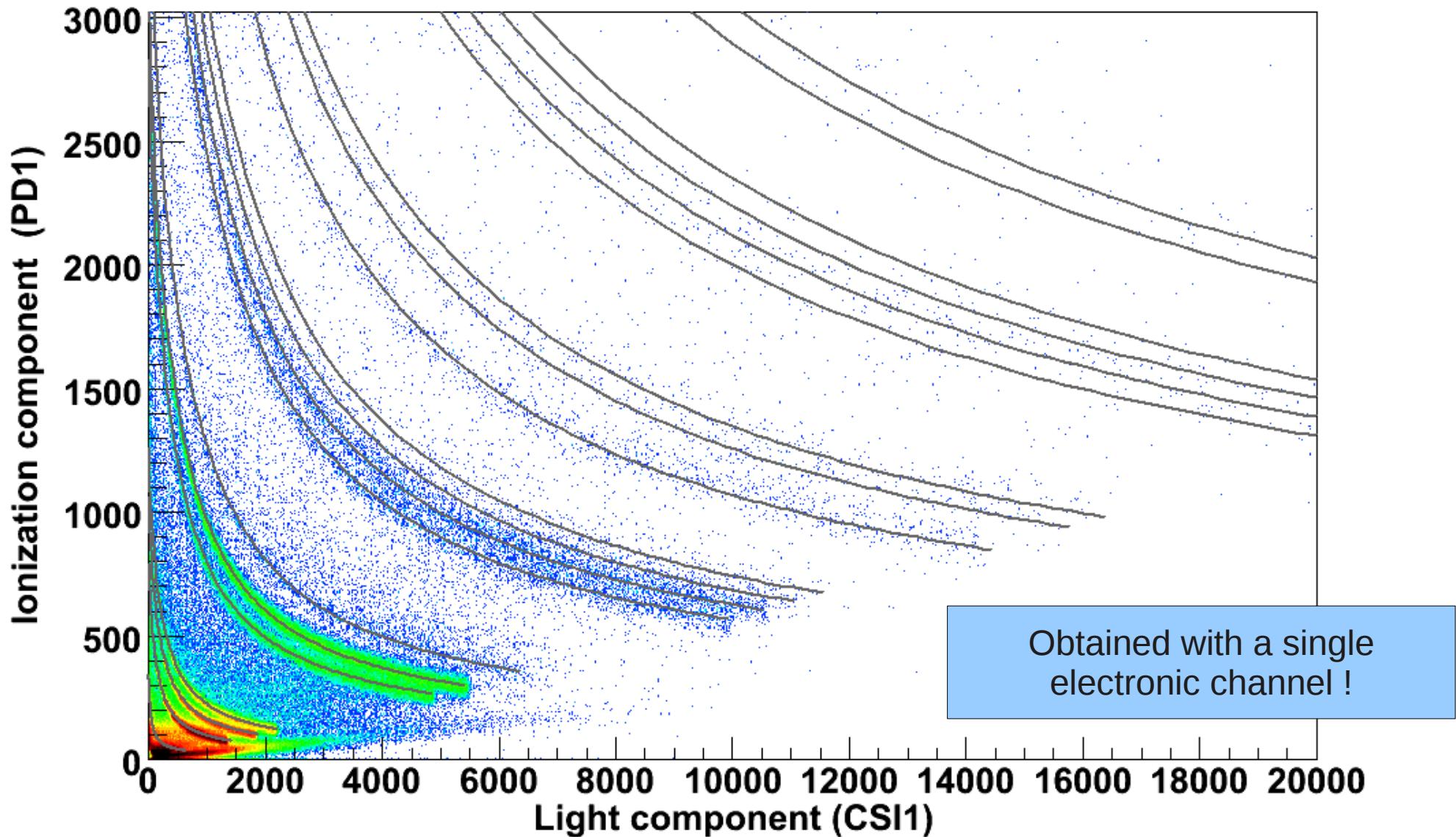


SCT decomposed (non-trivial)

(lines from the ATIMA range-energy tables)

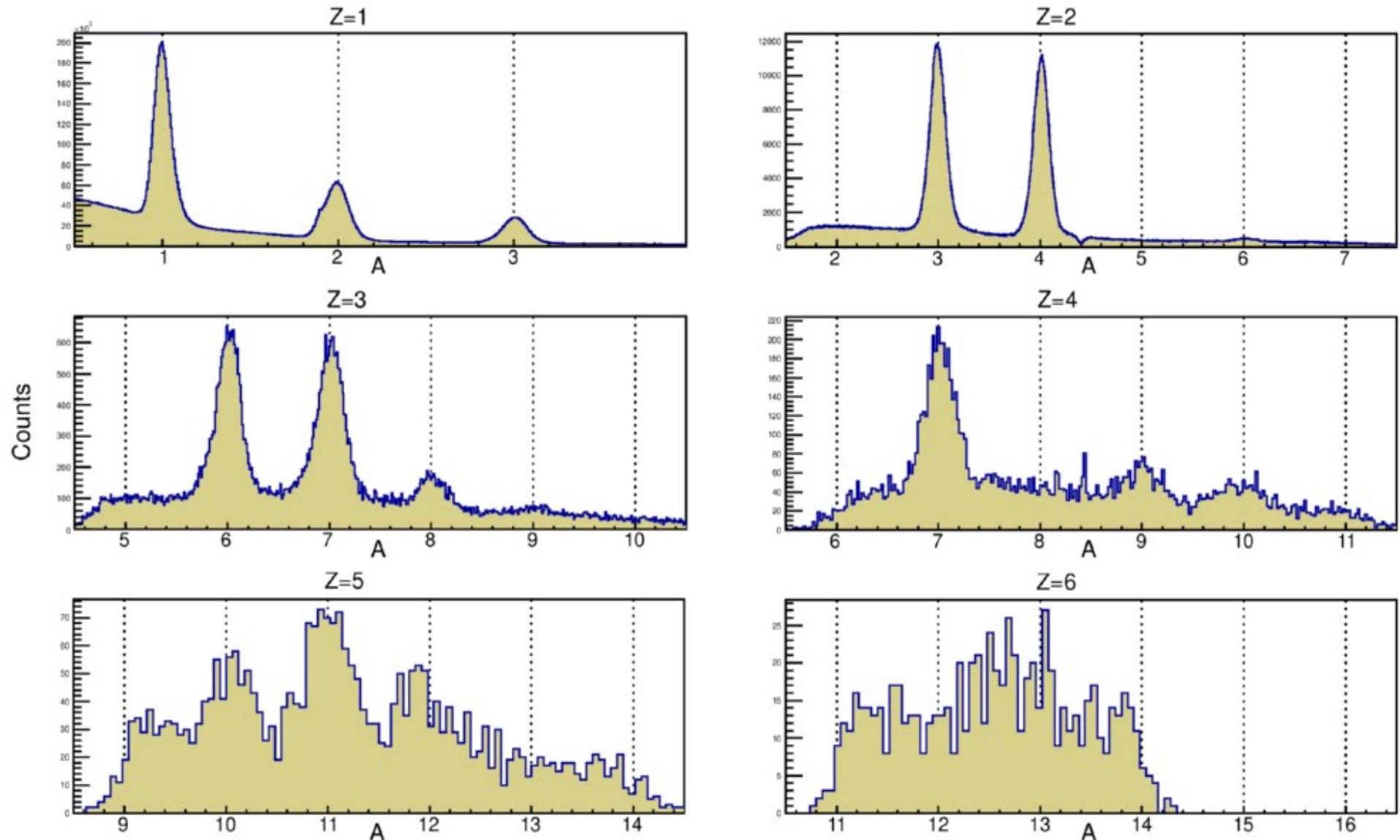
p10+p10/33.5:p12+p13-p10/33.5

Entries 1072173

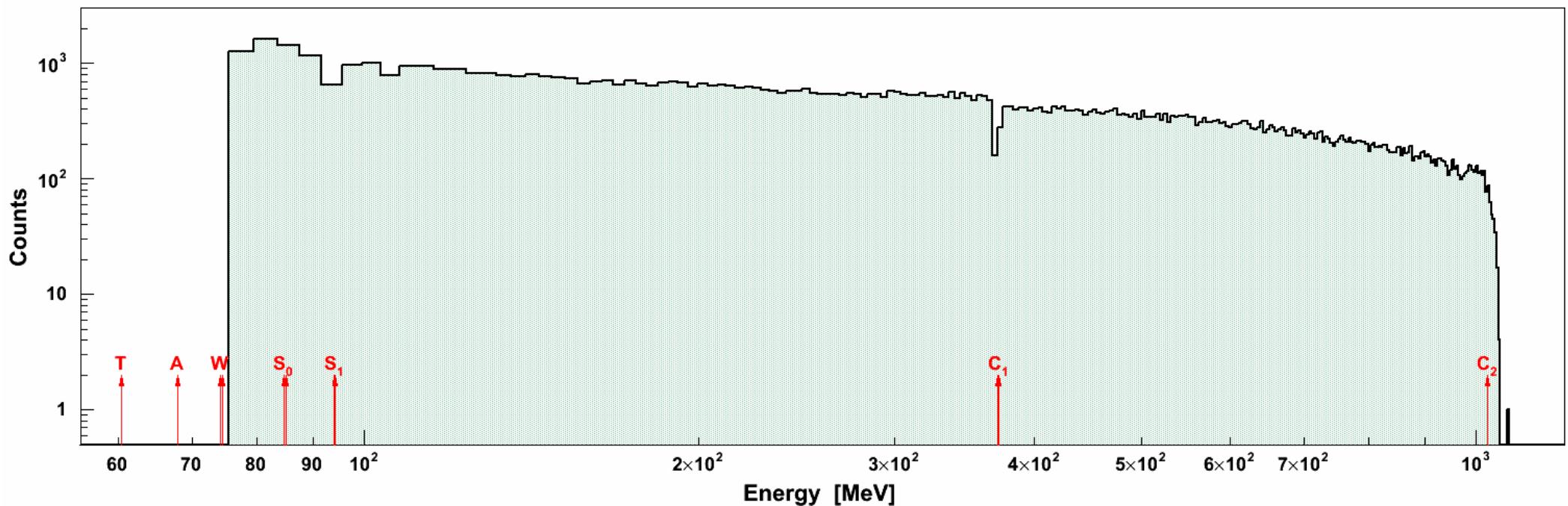


Mass distributions

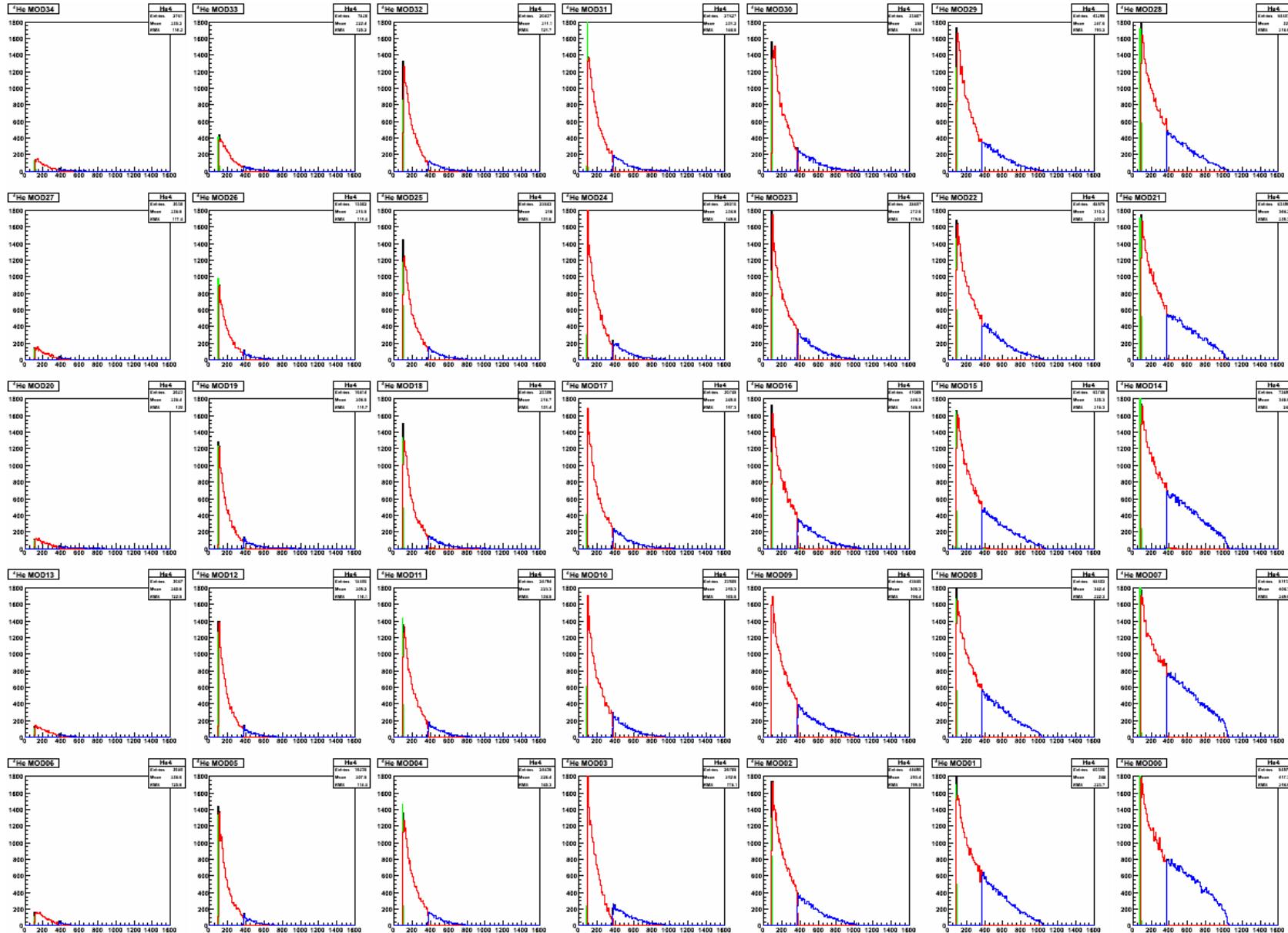
Au+Au @400 AMeV



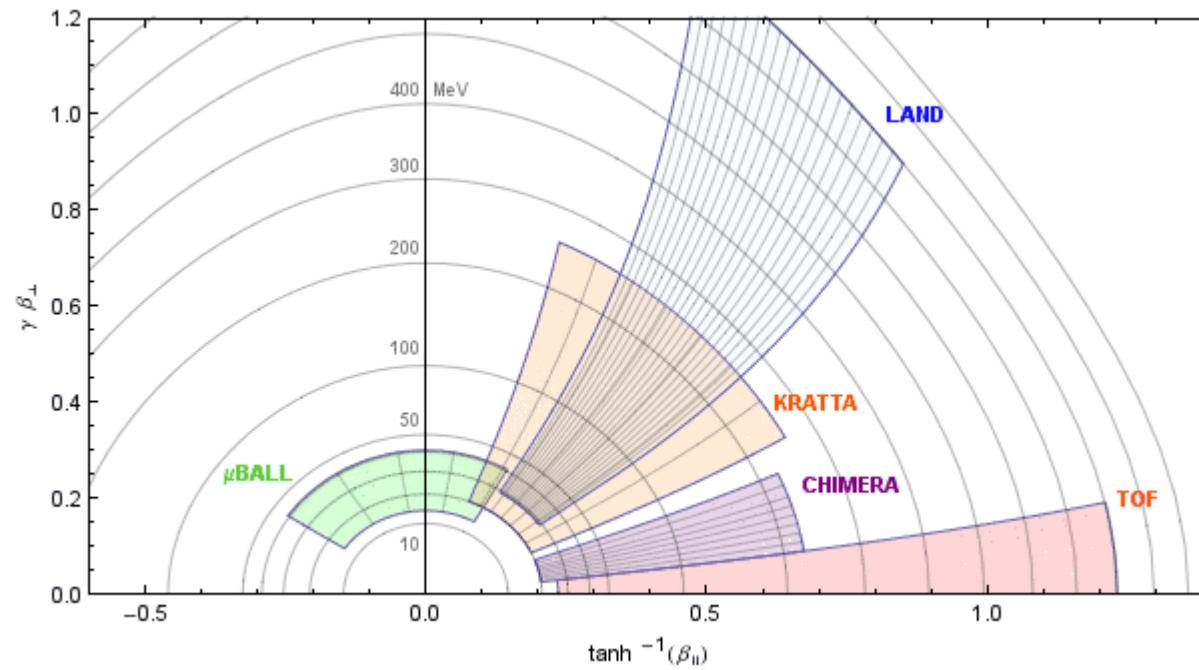
Alpha particle energy spectrum (MOD07) for Au+Au @400 AMeV (log-log scale)



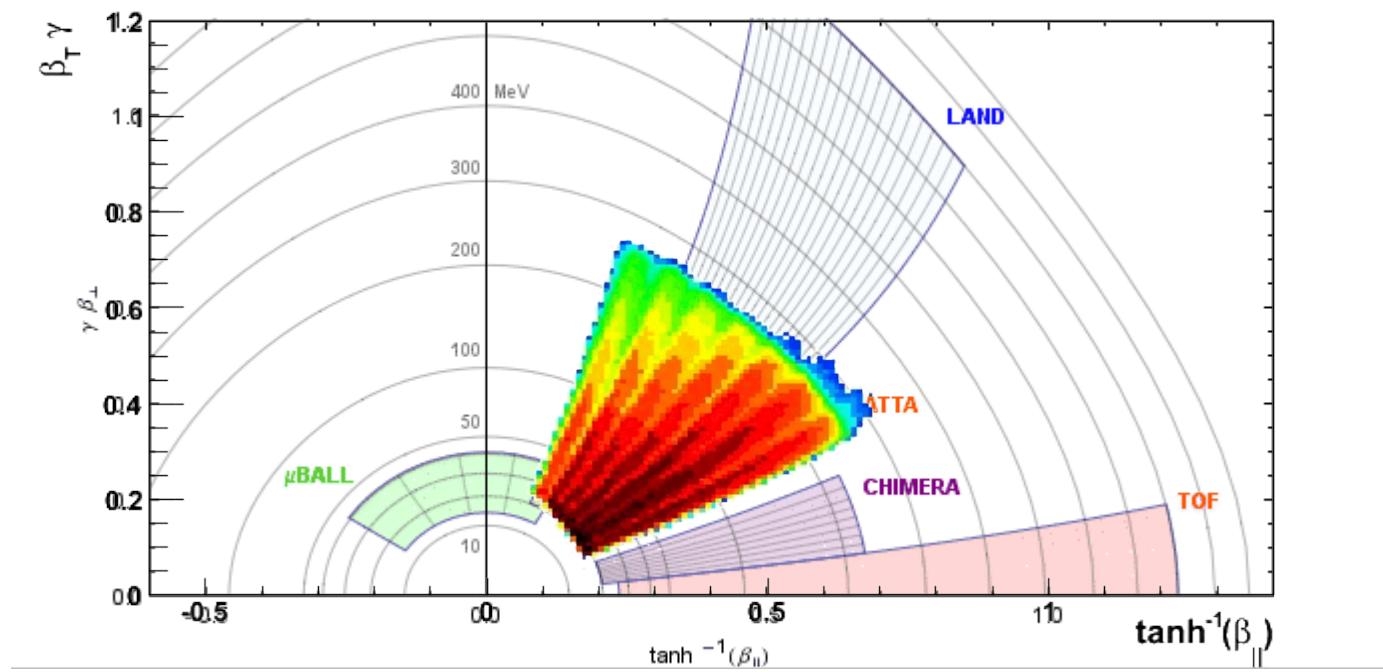
Alpha particles, all modules



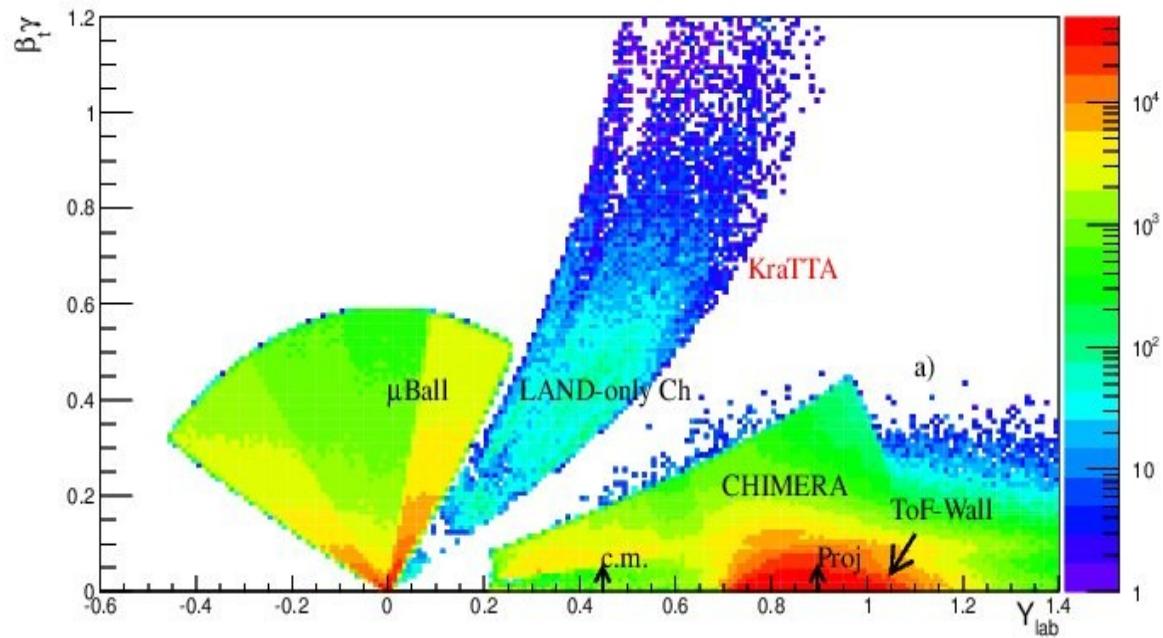
Acceptance of the setup (protons)



Acceptance of the setup (protons)

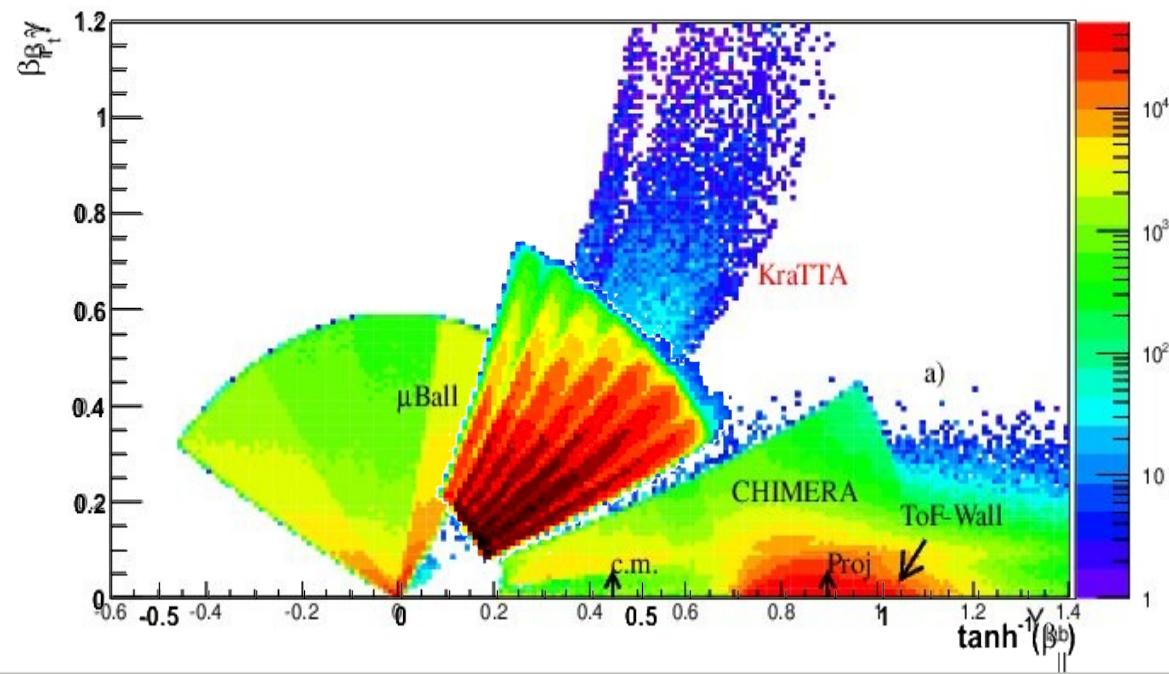


Acceptance of the setup (protons)



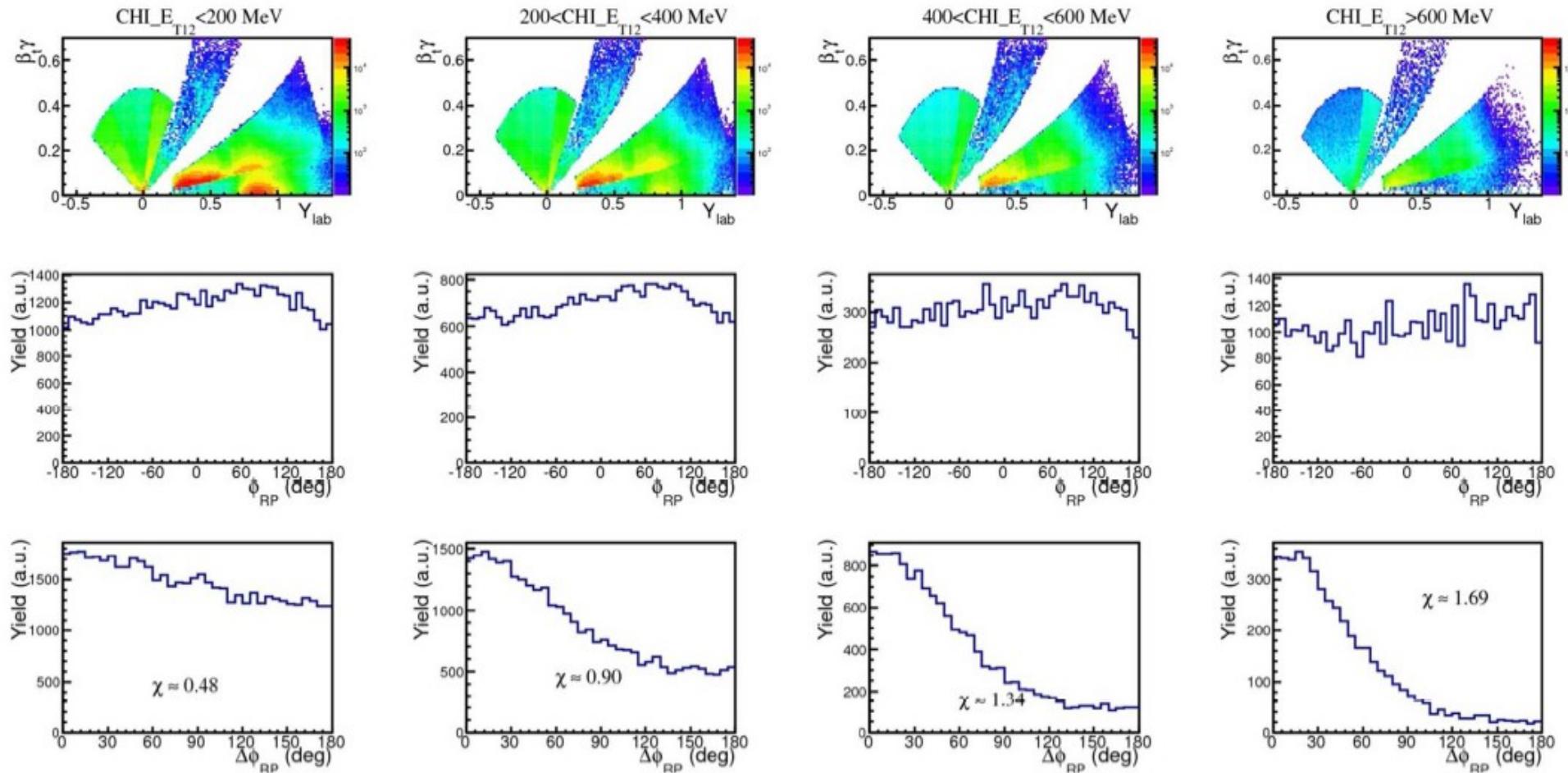
P. Russotto, INPC2013, Firenze, Italy 2-7.06.2013

Acceptance of the setup (protons)



Centrality selection and Reaction plane orientation Au+Au @ 400 AMeV

preliminary



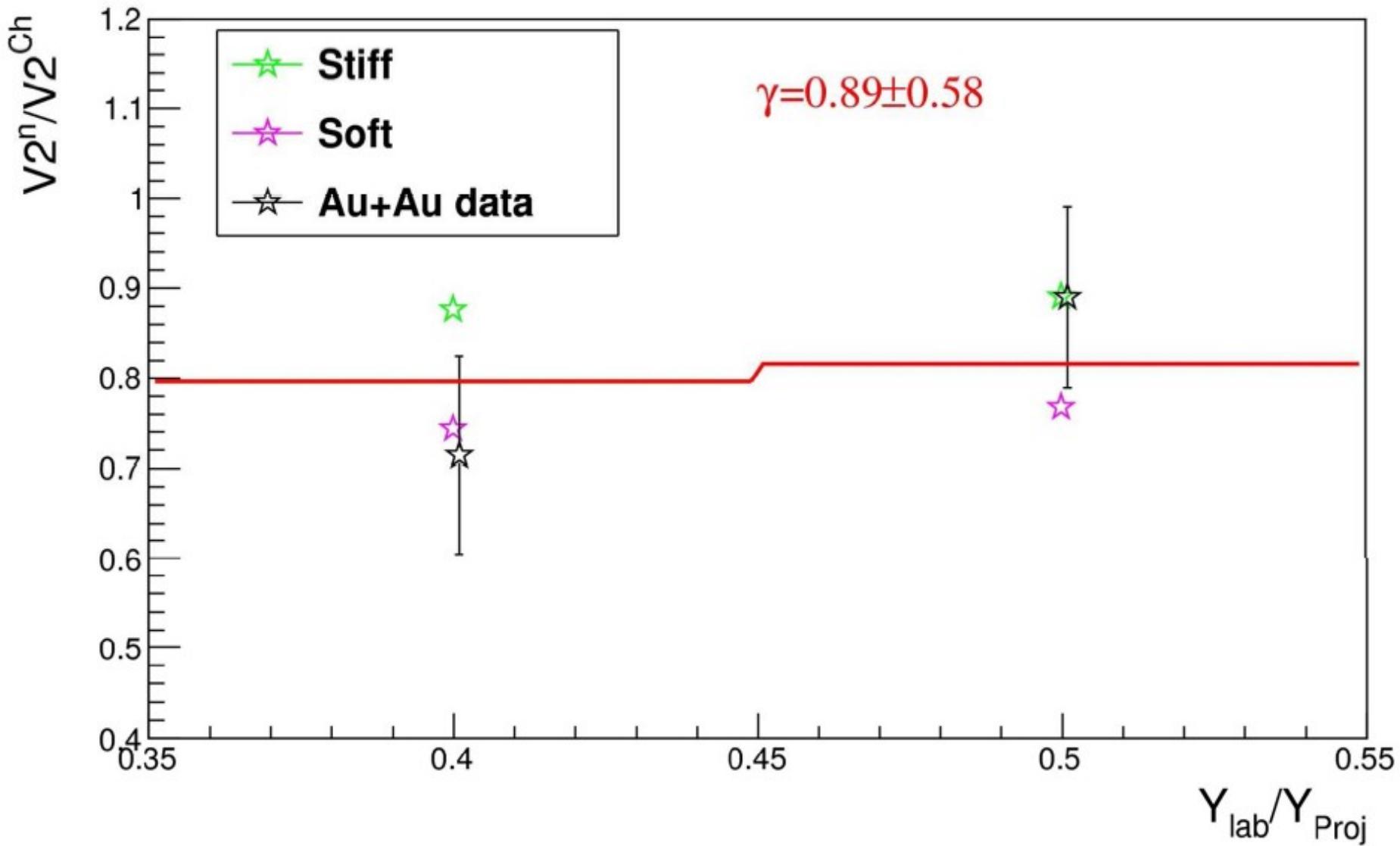
$$\vec{Q} = \sum_{v=1}^M Z_v \hat{p}_v w_v \quad w_v = \begin{cases} 1 & \text{for } Y_{c.m.} > 0.1 \\ 0 & \text{for } Y_{c.m.} < 0.1 \end{cases}$$

ad. from P. Danielewicz et al., PLB 1985

$E_{trans_{12}} (\text{MeV})$	$x \sim$	$\langle \cos(2\Delta\Phi) \rangle \sim$	$\Delta\Phi (\text{deg}) \sim$
<200	0.48	0.12	41
200-400	0.90	0.34	35
400-600	1.34	0.52	29
>600	1.69	0.66	24

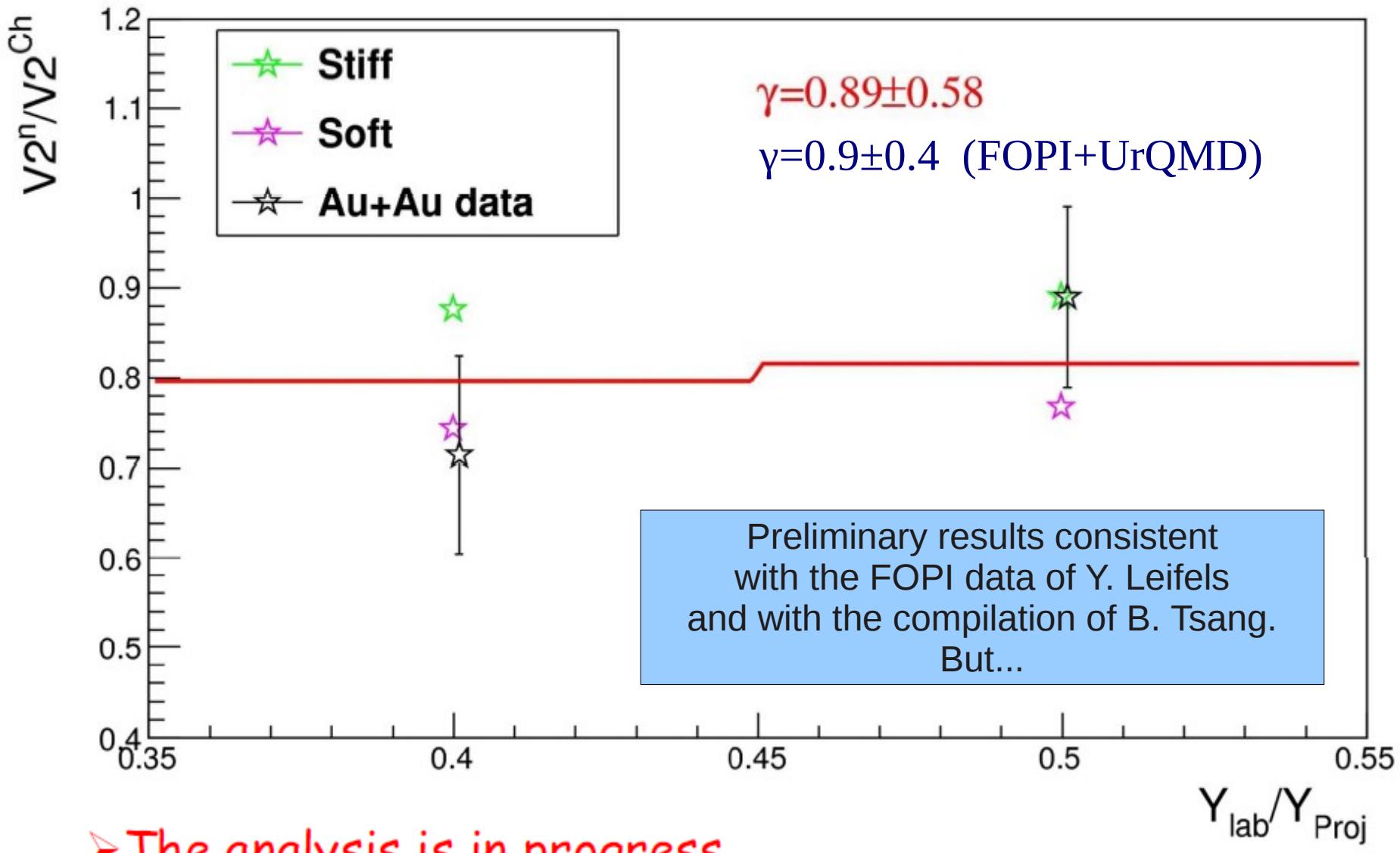
Preliminary gamma extraction

Au+Au @ 400 AMeV
 $b < 7.5$ fm



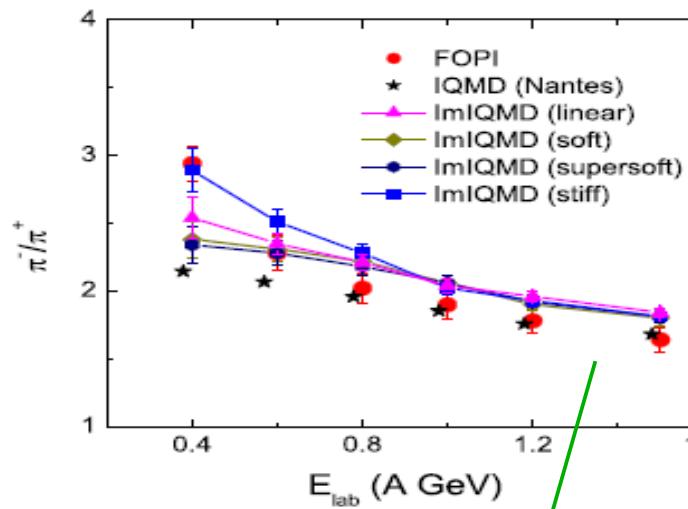
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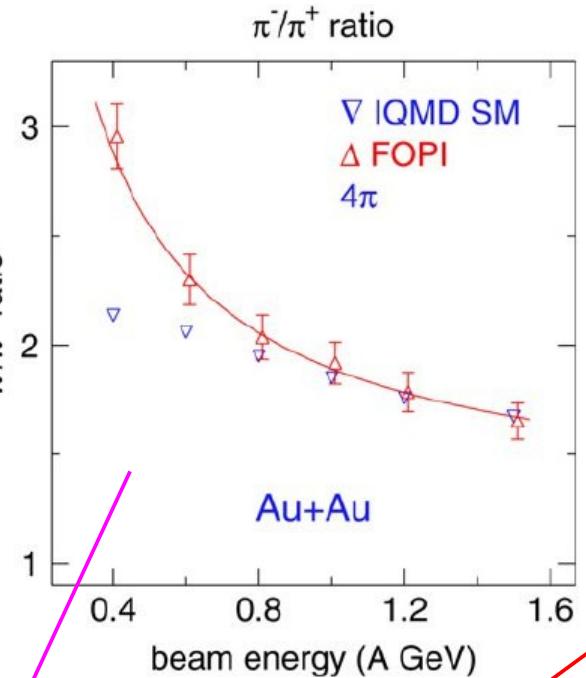


$$\frac{\pi^-}{\pi^+} = \frac{5n^2 + np}{5p^2 + np} \approx \left(\frac{n}{p}\right)^2$$

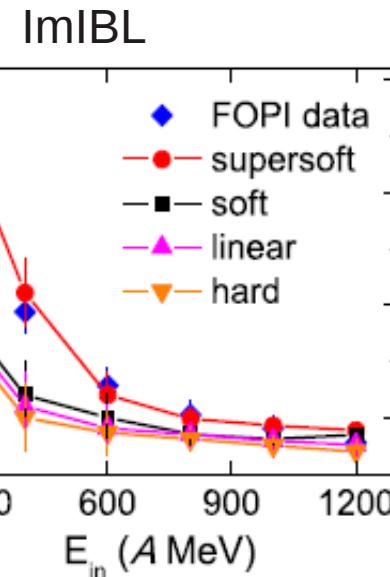
Pion ratio puzzle (FOPI, Au+Au)



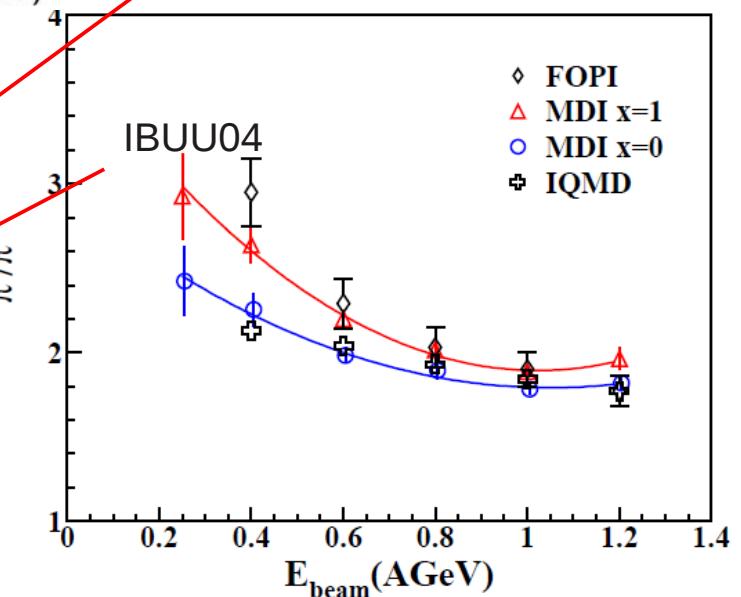
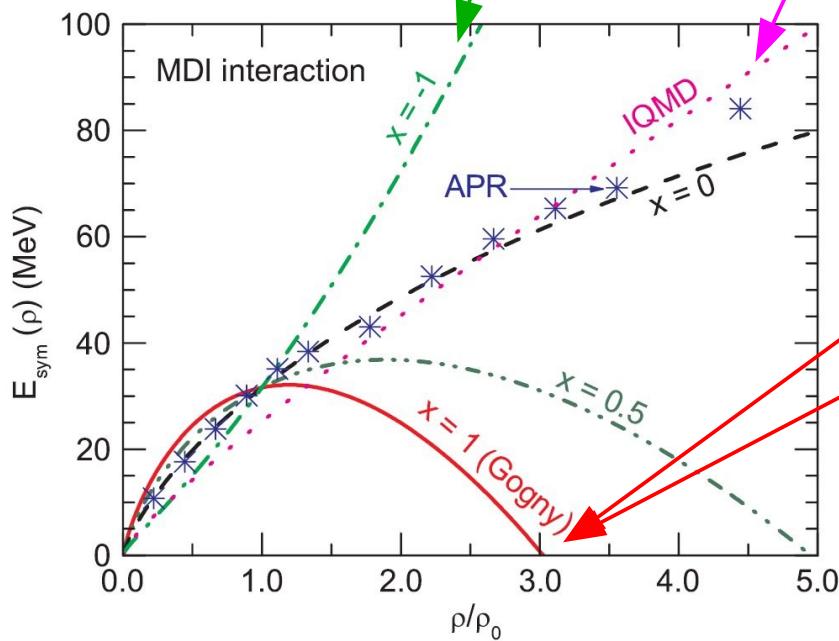
Feng, Jin, PLB 683 (2010) 140



W.Reisdorf et al. NPA 781 (2007) 459



Z. Xiao et al., PRL 102 (2009) 062502



See also W-M..Guo et al. ArXiv:1306.4783v1 [nucl-th] for UrQMD vs IBUU04

W.-J. Xie et al., PLB 718 (2013) 1510

Difficulties in measuring the $E_{\text{sym}}(\rho)$

Experiment

- Mixture of density, temperature and time dependent processes
- Detection of neutrons and protons simultaneously
- Tiny effects – high precision and statistics needed
- Observables minimizing the influence of the isoscalar part
- Correlations of many observables needed to disentangle competing effects
- Exotic beams, asymmetric systems (with larger δ)

Model

- In-medium cross sections, treatment of Δ resonance dynamics (π^-/π^+)
- Momentum dependence of the mean-field, effective masses
- Control the competition between the mean-field and collisions
- Realistic description of cluster formation (at least $t/{}^3\text{He}$)
- Ability to describe „hot” and „cold” observables. Often “hot” model observables are compared to “cold” experimental data.

What to measure?

- High energy tails (high p_T with high precision) of π^-/π^+ , n/p, t/ ${}^3\text{He}$, ..., → messengers of the high density first chance collisions, to resolve the influence of the competing effective mass splitting effect [Giordano et al.] and enhance sensitivity to the Esym [Q-F Li et al. JPG 32(2006) 151].
- Simultaneously π^-/π^+ , n/p, t/ ${}^3\text{He}$, ... to constrain the Δ properties and dynamics.
- Observables minimizing the influence of isoscalar part of the EoS and of the in-medium cross sections (ratios [Russotto et al.], double-ratios, differences [Cozma et al.] of flow observables, differential flows [B-A.Li PRL 2002]).
- Two or more beam energies to account for the competing isospin dependence of the N-N cross sections and to vary the densities. Construct excitation functions, measure e.g. the balance energy of neutrons (low energy) [Guo et al. SCP 55(2012)252]
- Observable trends as a function of centrality, rapidity and p_T .
- High energies ($E_{\text{beam}} \gg 100 \text{ AMeV}$), heavy systems ($A_{\text{sys}} > 100$), non-central collisions, heavy clusters [W. Reisdorf (FOPI) arXiv:1307.4210 [nucl-ex]]

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