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原子核三者若手夏の学校原子核パート研究会 @ 白浜荘 '11 8/16

- •Menu
  - I. What's "Pasta" ?
  - II. Relativistic Mean Field Theory

# III. Result(i) Fixed proton ratio (ii) $\beta$ - equilibrium (iii) Large cell calculation

IV. Conclusion / Future plan

- •Menu
  - I. What's "Pasta" ?

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#### Suggestion of Pasta Structure

1.C Nuclear Physics A175 (1971) 225-271 Not to be reproduced by photoprint or mice	; C North-Holland Publishing Co., Amsterdam crofilm without written permission from the publisher
NEUTRON STA GORDON BAYM <sup>+</sup> , HANS A. BETHE <sup>++</sup> Nordita, Copenhag Received 4 M	AR MATTER and CHRISTOPHER J. PETHICK <sup>†††</sup> gen, Denmark May 1971
which become important here. In fact, it for the nuclei to "turn inside out", that is droplets in a sea of nuclear matter.	might be more favorable, beyond $u = 0.5$ , , for the neutron gas to exist as a lattice of
PHYSICAL REVIEW LETTERS 27 JUNE 1983	Progress of Theoretical Physics, Vol. 71, No. 2, February 1984
D. G. Ravenhall ysics, University of Illinois at Urbana-Champaign, Urbana, Illinois 61801 and	Shape of Nuclei in the Crust of <u>Neutron Star</u> Masa-aki HASHIMOTO, Hironori SEKI and Masami YAMADA*
C. J. Pethick ysics, University of Illinois at Urbana-Champaign, Urbana, Illinois 61801, and NORDITA, DK-2100 Copenhagen Ø, Denmark	Department of Physics and Applied Physics, Waseda University, Tok *Science and Engineering Research Laboratory, Waseda University, To
and J. R. Wilson e Livermore National Laboratory, Livermore, California 94550 (Received 5 May 1983)	(Received August 30, 1983)
It will be interesting t es of these spaghettilike of dense matter. Their have to reflect the group	to explore the consequenc- e and lasagnalike phases physical properties will
that these phases posses	ss. Neutrino scattering

VOLUME 50, NUMBER 26

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## Pasta Structure = Inhomogeneous structures appear in first phase transition, "mixed phase with structure"





(K.Oyamatsu, Nucl.Phys.A561,431(1993))

Pasta Structure = Inhomogeneous structure appear in first phase transition, "mixed phase with structure"

Balance with Surface tension and Coulomb repulsion

Total Energy =  $E_b$ (Volume) +  $E_s$ (surface) +  $E_c$ (Coulomb)

Weizsäcker-Bethe's semi-emperical mass formula

 $\frac{E_{C}}{A} \propto \frac{Z^{2}/A^{1/3}}{A} \propto R^{2} \Rightarrow \frac{E_{C}}{A} = aR^{2}$   $\frac{E_{S}}{A} \propto \frac{A^{2/3}}{A} \propto R^{-1} \Rightarrow \frac{E_{S}}{A} = bR^{-1}$   $\frac{d(E_{S}/A + E_{C}/A)}{dR} = 2aR - bR^{-2} = 0$   $\therefore 2aR^{2} = bR^{-1} 2E_{S} = E_{C}$  E/A Coulomb + Surface Coulomb Surface



#### PHYSICAL REVIEW C 75, 042801(R) (2007)

Impact of nuclear "pasta" on neutrino transport in collapsing stellar cores

Hidetaka Sonoda,<sup>1,2</sup> Gentaro Watanabe,<sup>3,2</sup> Katsuhiko Sato,<sup>1,4</sup> Tomoya Takiwaki,<sup>1</sup> Kenji Yasuoka,<sup>5</sup> and Toshikazu Ebisuzaki<sup>2</sup> <sup>1</sup>Department of Physics, University of Tokyo, Tokyo 113-0033, Japan

<sup>2</sup>The Institute of Chemical and Physical Research (RIKEN), Saitama 351-0198, Japan <sup>3</sup>NORDITA, Blegdamsvej 17, DK-2100 Copenhagen Ø, Denmark
 <sup>4</sup>Research Center for the Early Universe, University of Tokyo, Tokyo 113-0033, Japan <sup>5</sup>Department of Mechanical Engineering, Keio University, Yokohama 223-8522, Japan (Received 11 January 2007; published 6 April 2007)

#### <u>Neutrino differential cross-section</u> <u>in supernova</u>

 Pasta structures of quark-hadron mixed phase in Neutron star core

... etc)



Whole space is divided into equivalent cells. These cells are imposed geometrical symmetry as follow. Sphere : 3D, Cylinder : 2D, Slab : 1D







Existence of other structure

G.Watanabe, H.Sonoda et.al PRL **103**, **121101 (2009)** M.Matsuzaki PRC **73**, **028801 (2006)** 

K.Nakazato, K.Oyamatsu et.al PRL 103, 132501 (2009)





Gyroid

Double-Diamond

Development of computer performance



Full 3D calculation

•Menu

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#### **Relativistic Mean Field Theory**

$$L = \overline{\psi} \left( i \gamma^{\mu} \partial_{\mu} - m - g_{\sigma} \sigma - g_{\omega} \gamma^{\mu} \partial_{\mu} - g_{\rho} \gamma^{\mu} \tau^{a} \rho_{\mu}^{a} \right) \psi \qquad : \text{Neucleon}$$

$$+\frac{1}{2}\partial_{\mu}\sigma\partial^{\mu}\sigma - \frac{1}{2}m_{\sigma}^{2}\sigma^{2} + \frac{1}{3}bm_{\sigma}(g_{\sigma}\sigma)^{3} + \frac{1}{4}c(g_{\sigma}\sigma)^{3} \qquad :\sigma - meson$$

$$-\frac{1}{4}F_{\mu\nu}F^{\mu\nu} + \frac{1}{2}m_{\rho}^{2}\rho_{\mu}^{a}\rho^{a\mu} + \frac{1}{2}\left[\frac{1}{2}\left(\partial^{\mu}\rho^{a\nu} - \partial^{\nu}\rho^{a\mu}\right)\left(\partial_{\mu}\rho_{\nu}^{a} - \partial_{\nu}\rho_{\mu}^{a}\right)\right] \qquad :\rho - \text{meson}$$

$$+\frac{1}{2}m_{\omega}^{2}\omega_{\mu}\omega^{\mu}+\frac{1}{2}\left[\frac{1}{2}\left(\partial^{\mu}\omega^{\nu}-\partial^{\nu}\omega^{\mu}\right)\left(\partial_{\mu}\omega_{\nu}-\partial_{\nu}\omega_{\mu}\right)\right]\qquad\qquad:\omega-\mathrm{meson}$$

mean field approx.



Thomas Fermi approx.zero temperature

$$-\nabla^{2}\sigma(\vec{r}) + m_{\sigma}^{2}\sigma(\vec{r}) = -\frac{dU(\sigma)}{d\sigma} + g_{\sigma N}\left(\rho_{n}^{s}(\vec{r}) + \rho_{p}^{s}(\vec{r})\right)$$
$$-\nabla^{2}\omega_{0}(\vec{r}) + m_{\omega}^{2}\omega_{0}(\vec{r}) = g_{\omega N}\left(\rho_{p}(\vec{r}) + \rho_{n}(\vec{r})\right)$$
$$-\nabla^{2}\rho_{0}(\vec{r}) + m_{\rho}^{2}\rho_{0}(\vec{r}) = g_{\rho N}\left(\rho_{p}(\vec{r}) - \rho_{n}(\vec{r})\right)$$
$$\nabla^{2}V(\vec{r}) = 4\pi e^{2}\left(\rho_{p}(\vec{r}) + \rho_{e}(\vec{r})\right)$$
$$\mu_{p} = \mu_{B} - \mu_{e} + V = \upsilon_{p} + g_{\omega N}\omega_{0} + g_{\rho N}\rho_{0}$$
$$\mu_{n} = \mu_{B} = \upsilon_{n} + g_{\omega N}\omega_{0} - g_{\rho N}\rho_{0}$$
$$\mu_{e} = \left(3\pi^{2}\rho_{e}(\vec{r})\right)^{1/3} + V(\vec{r})$$

#### Parameters

To reproduce the saturation properties of symmetric nuclear matter and the properties of finite nuclei

$$g_{\sigma N} = 6.3935 \ g_{\omega N} = 8.7207 \ g_{\rho N} = 4.2696 \ b = 0.008659 \ c = -0.002421$$
  
 $m_N = 938 [\text{MeV}] \ m_\sigma = 400 [\text{MeV}] \ m_\omega = 783 [\text{MeV}] \ m_\rho = 769 [\text{MeV}]$ 



PHYSICAL REVIEW C 72, 015802 (2005)



## • How to solve $\cdot \cdot ?$

- Introduce a cubic cell with periodic boundary condition and divide it into grids
- · As an initial condition, randomly distribute fermions (n, p, e) over the grid
- We solve coupled differential equations, and simultaneously relax fermions density distributions to attain the uniformity of their chemical potential



•Menu

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**IV. Conclusion / Future plan** 

#### Fixed proton ratio (i)

## Basic structure of "Pasta" : Like "atom / molecule" which construct "Crystal"



Basic structure only appears by the calculation using W-S approximation.

Energy and size of basic structure (R) and cell size  $(R_{\rm w})$ are calculated in detail.

Compare the Energy and cell size of basic structure with W-S approximation.  $R = \begin{cases} R_w \frac{\langle \rho_p \rangle^2}{\langle \rho_p^2 \rangle} (droplet, rod, slab) \\ R_w \left( 1 - \frac{\langle \rho_p \rangle^2}{\langle \rho_p^2 \rangle} \right) (tube, bubble) \end{cases}$ 

#### (i) Fixed proton ratio [Yp(=(A-Z)/A)=0.5]





#### Mechanism of clusterization

- Total Pressure : positive by electron partial pressure
- Baryon partial pressure :
  - negative in  $\rho_{\rm B} < \rho_0$ 
    - $\rightarrow$  unstable
    - $\rightarrow$  clusterization (pasta structure)
- $^{\circ}$   $\rho_0 < \rho_{\rm B}$ : Uniform matter ← Energy loss of Coulomb repulsion and Surface tension

( $\rho_{0:}$  normal nucleus density)



#### (ii) $\beta$ -equilibrium (Crust of neutron star)



#### (iii) Crystal structure

Calculation in small cell: •Without symmetry •Periodic boundary condition



Limitation for crystal structure by cell size



Calculation in Large cell

## Pasta crystal structure



"bubble"



## **Body-centered cubic**



"tube"



(K.Oyamatsu, Nucl.Phys.A561,431(1993))

(iii) Crystal structure ( $Y_p = 0.5$ ) droplet



 $\rho_B$ =0.01 fm<sup>-3</sup> : body-centered cubic

scalars

0.078379

0.06

0.04

0.02



"simple" or "bcc" or "" or "fcc" or ?

"bcc" → "fcc"

(change by baryon density)

 $\rho_B=0.015 \text{ fm}^{-3}$  : face-centered cubic



(iii) Crystal structure ( $Y_p = 0.5$ ) rod



"Simple" or "Honey-cum" or •••?  $\rho_B$ =0.022 fm<sup>-3</sup> : simple cubic



scalars 0.078379 0.06 0.04 0.02 0



 $\rho_B$ =0.028 fm<sup>-3</sup> : Honey-cum

![](_page_21_Picture_8.jpeg)

#### (iii) Crystal structure <u>slab</u> $\rho_{\rm B}$ =0.05 fm<sup>-3</sup>

![](_page_22_Picture_1.jpeg)

![](_page_22_Picture_2.jpeg)

![](_page_22_Figure_3.jpeg)

### <u>tube</u>

![](_page_22_Picture_5.jpeg)

"Simple" or "Honeycomb" or •••?

"Honeycomb"

#### $\rho_B=0.075 \text{ fm}^{-3}$ : Honey-cum

![](_page_22_Picture_9.jpeg)

#### (iii) Crystal structure

## <u>bubble</u>

![](_page_23_Picture_2.jpeg)

"Simple" or "fcc" or "bcc" or •••?

"bcc"  $\rightarrow$  "fcc" (change by baryon density)

#### $\rho_B$ =0.085 fm<sup>-3</sup> : face-centered cubic

![](_page_23_Picture_6.jpeg)

scalars 0.078379 0.06 0.04 0.02

 $\rho_B=0.09 fm^{-3}$  : body-centered cubic

![](_page_23_Picture_9.jpeg)

#### (iii) Crystal structure

#### **Complex Pasta structure**

 $\rho_B = 0.015 \text{ fm}^{-3}$  : dumbell

![](_page_24_Picture_3.jpeg)

![](_page_24_Picture_4.jpeg)

![](_page_24_Picture_5.jpeg)

小さいセルでの計算との比較 : 0.14MeV低い 大きいセルでの比較 : 未実行

プログラムミス(?)

# IV. Conclusion / Future

We demonstrate 3D calculation of non-uniform low-density nuclear matter based on Relativistic mean field theory and Thomas-Fermi approximation.

- For fixed proton ratio calculation, we perform same cell size calculation of W-S approximation and get almost same pasta structures and baryon density dependence of binding energy.
- For β-equilibrium calculation, only sphere shape appears.
  It is similar result with W-S cell approximation.
- We perform large size cell calculation in which some basic structure appear.

# IV. Conclusion / Future

- Expansion of cell size
  - What crystal structure is the most stable state?
  - EOS for various proton fraction ratio
- > Extension to high density and finite temperature nuclear matter

Comparing with QMD calculation result for supernova compression process and local minimum states of our calculation.

![](_page_27_Picture_0.jpeg)

![](_page_28_Picture_0.jpeg)

Yp=0.1

#### Yp=0.3

![](_page_29_Figure_2.jpeg)

Yp=0.5

![](_page_29_Figure_4.jpeg)

![](_page_30_Figure_0.jpeg)

#### Complex Pasta structure

![](_page_31_Picture_1.jpeg)

![](_page_31_Picture_2.jpeg)

![](_page_31_Picture_3.jpeg)

## 核物質の結合エネルギー

![](_page_32_Figure_1.jpeg)

#### 核物質の結合エネルギーのモデルの1例。 密度ρ<sub>0</sub>(約0.16fm<sup>-3</sup>)の対称核物質 がもっとも安定。

核物質の固さ(incompressibility)

$$K = p_F^2 \frac{d^2 \varepsilon}{dp_F^2} = 9\rho^2 \frac{d^2 \varepsilon}{d\rho^2} = 9\frac{dP}{d\rho}$$

は重要な量だが、まだ決まっていない。 図の赤線は*K* = 240 MeVの2次曲線。

#### パスタ構造 = 一次相転移に伴う物質の非一様構造 「構造を持った混合相」

## Total Energy = (bulk) + (Surface) + (Coulomb)

クーロン斥力と表面張力の釣合いによる規則的な構造

![](_page_33_Figure_3.jpeg)

![](_page_34_Picture_0.jpeg)

### どの近似でもパスタ構造は現れる

## (1)Liquid-drop model (Macroscopic)

- Ravenhall, Pethick&Willson, PRL 50, 2066(1983)
- ·Hashomoto, Seki & Yamada, PTP 71, 320 (1984)
- Lorentz, Ravenhall & Pethick, PRL 70,379 (1993)

## (2)Thomas-Fermi (Semi-classical)

- •Williams&Koonin, NPA 435, 844 (1985)
- •Oyamatsu, NPA 561,431 (1993)
- ·Sumiyoshi&Oyamatsu&Toki ,NPA 595, 323 (1995)

## (3)Hartree-Fock (Quantum)

- •Magierski&Heenen, PRC 65, 045804 (2002)
- •Gögelein & Müther, PRC 76, 024312 (2007)
- •Newton & Stone, PRC 79, 055801 (2009)

![](_page_35_Figure_13.jpeg)

![](_page_35_Picture_14.jpeg)

![](_page_35_Picture_15.jpeg)

Newton (2009)

非対称核物質 Y<sub>p</sub>=0.1

![](_page_36_Figure_1.jpeg)

T. Maruyama, T. Tatsumi Recent Res. Devel. in Phys. 7 (2006) 1.

対称核物質 Y<sub>p</sub>=0.5

![](_page_36_Figure_4.jpeg)

0.10 0.05 3D, 0.013  $\begin{array}{c} 0.00\\ 0.10\end{array}$ 0.05 20, 0.030  $\begin{array}{c} 0.00\\ 0.10\end{array}$ 0.05 1D, 0.050  $\begin{array}{c} 0.00\\ 0.10\end{array}$ 0.05 2D, 0.070  $\begin{array}{c} 0.00\\ 0.10\end{array}$ 0.05 3D, 0.090 0.00 10 20

非対称核物質 Y<sub>p</sub>=0.3