

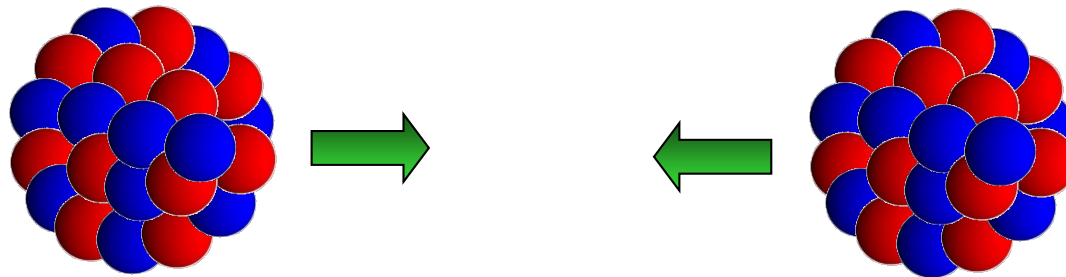
Perspectives on nuclear reaction theory and superheavy elements

Kouichi Hagino

Tohoku University, Sendai, Japan



TOHOKU
UNIVERSITY



1. Nuclear Reactions: overview
2. Heavy-ion fusion reactions
3. Fusion for superheavy elements
4. Summary

Introduction: low-energy nuclear physics

□ behaviors of atomic nuclei as a quantum many-body systems

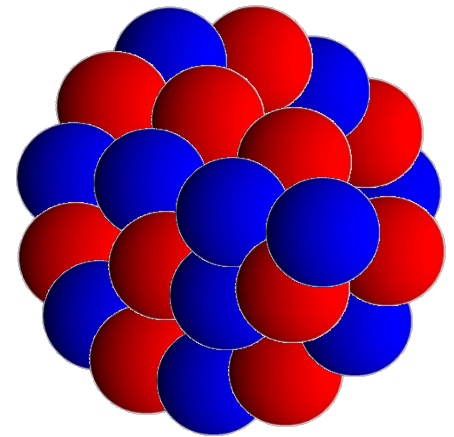
← understanding based on strong interaction

➤ static properties: nuclear structure

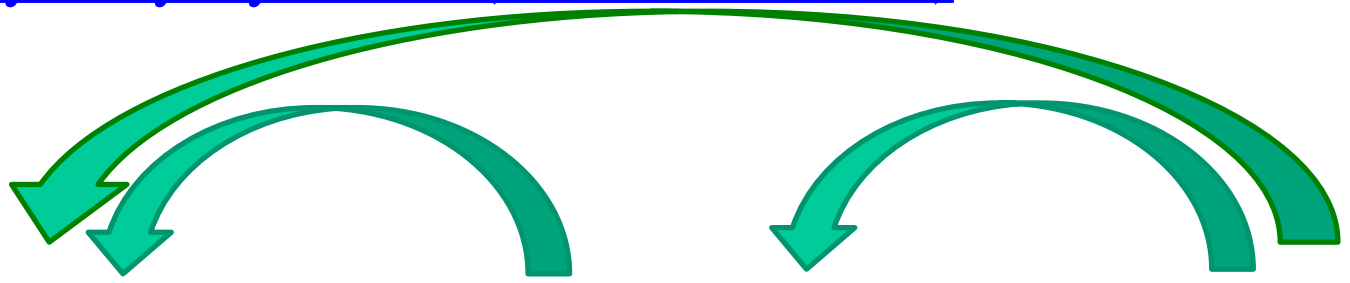
- ✓ ground state properties
(mass, size, shape,....)
- ✓ excitations
- ✓ nuclear matter

➤ dynamics: nuclear reactions

an interplay between these



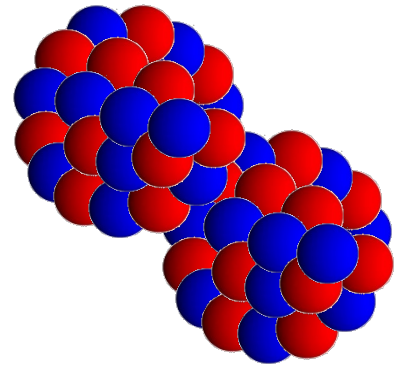
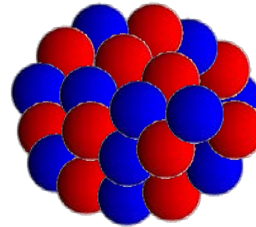
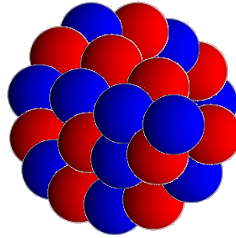
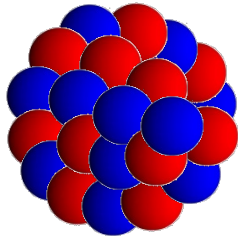
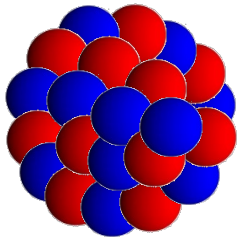
Quantum Many-body Dynamics (nuclear reactions)



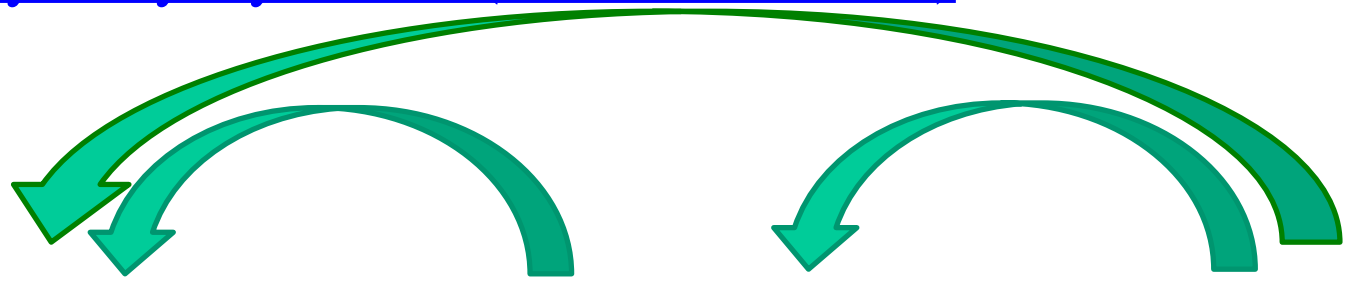
elastic scattering

inel. scattering

fusion



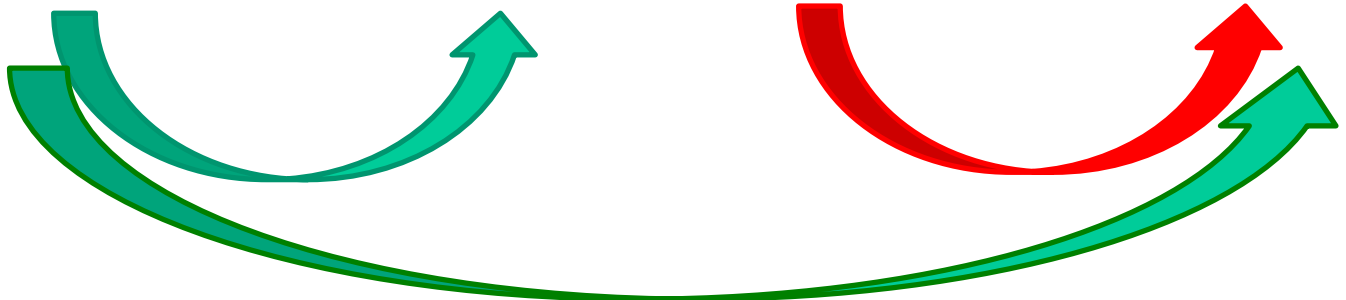
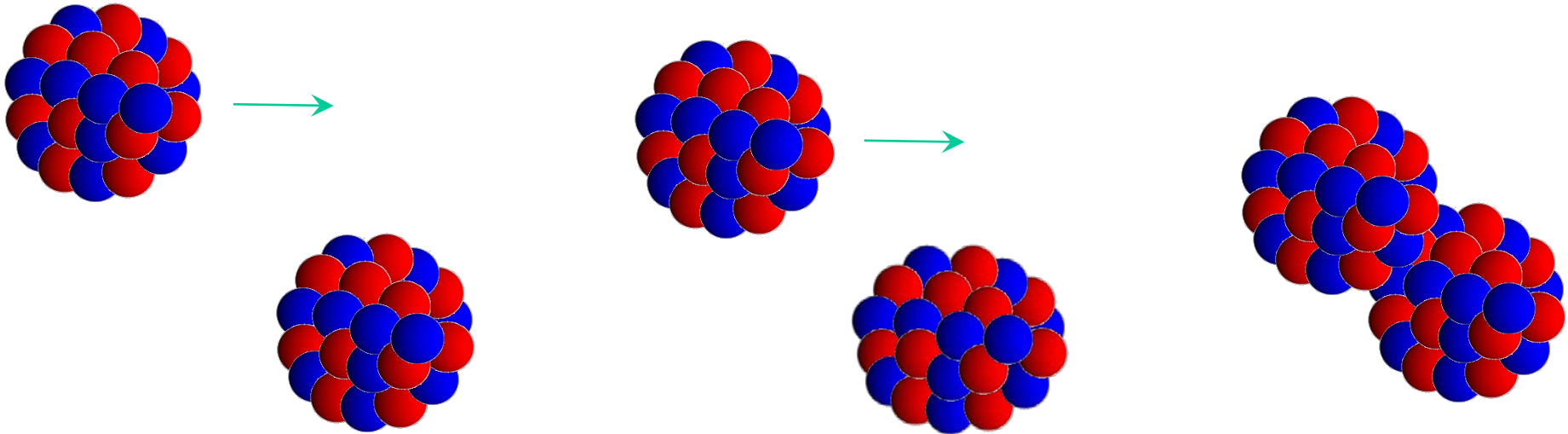
Quantum Many-body Dynamics (nuclear reactions)



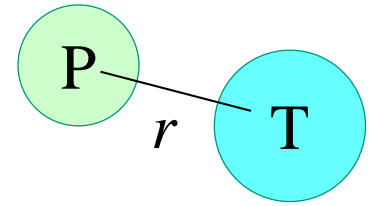
elastic scattering

inel. scattering

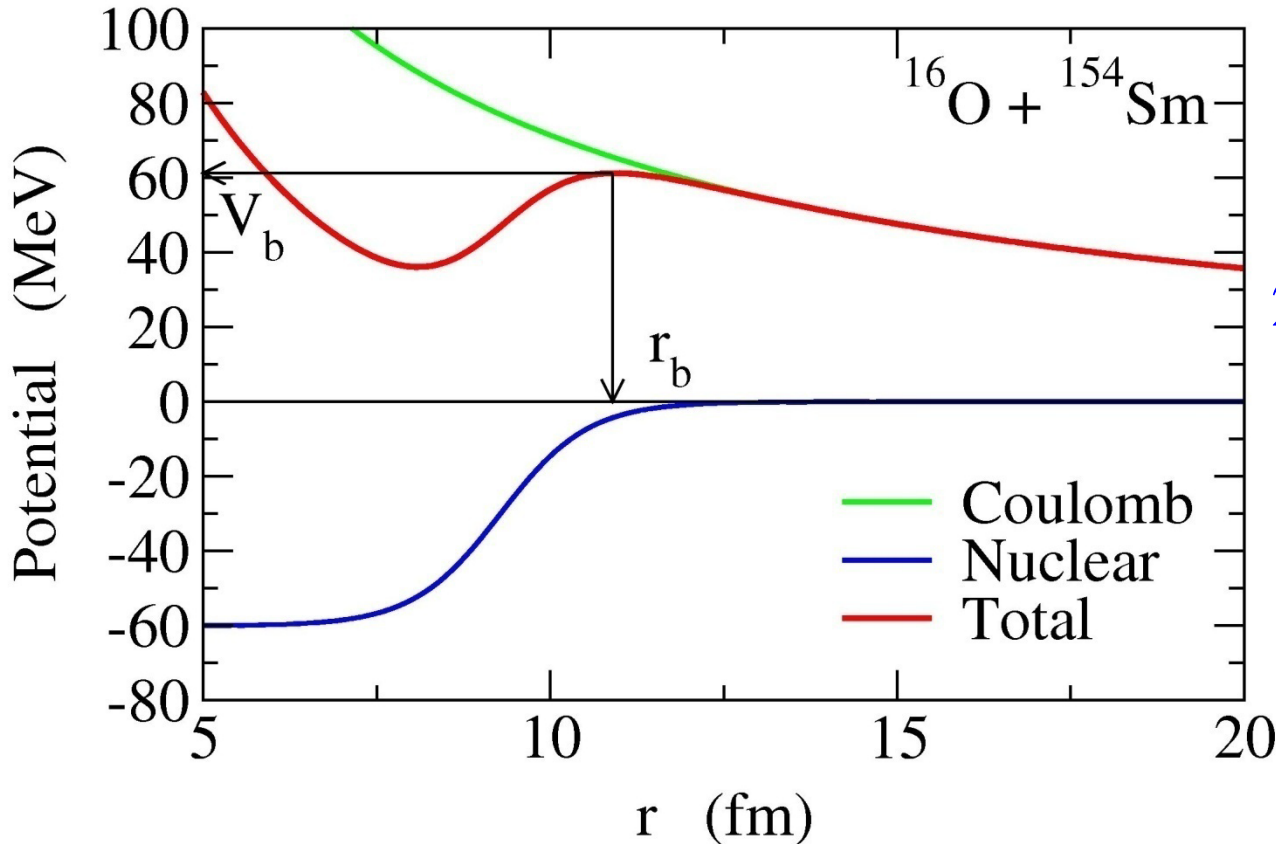
fusion



Coulomb barrier



Coulomb barrier



1. Coulomb interaction
long range
repulsion

2. Nuclear interaction
short range
attraction



Potential barrier
(Coulomb barrier)

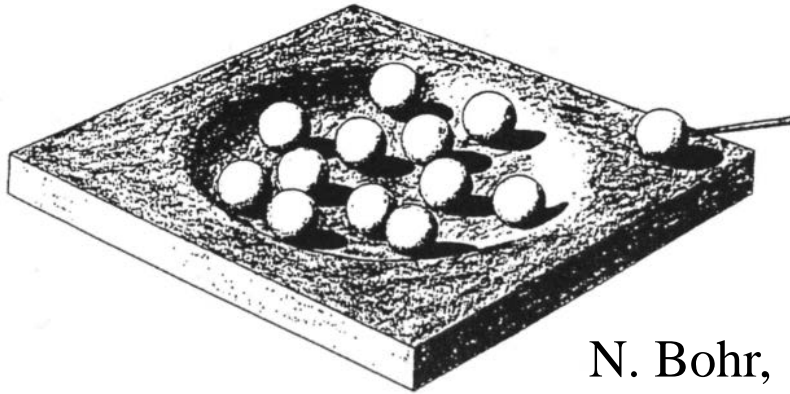
the barrier height \rightarrow defines the energy scale of a system

Fusion reactions at energies around the Coulomb barrier

Fusion reactions: compound nucleus formation

Niels Bohr (1936)

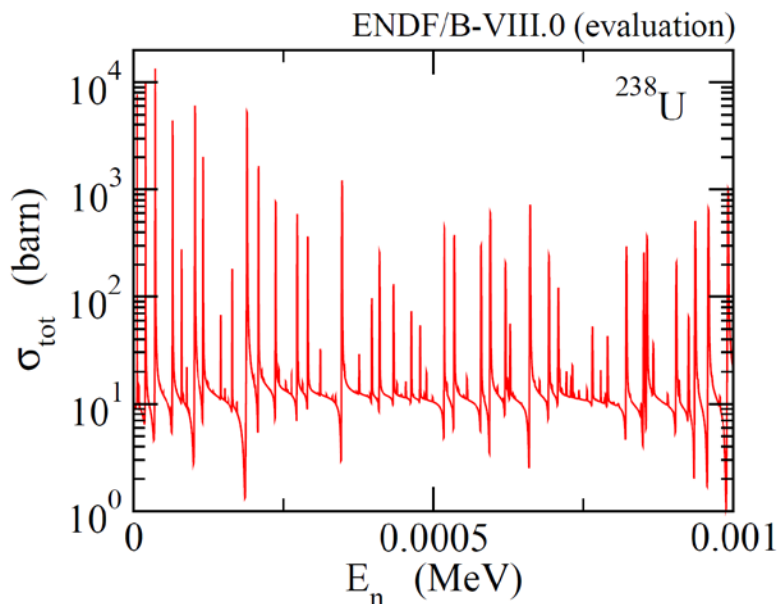
Neutron capture of nuclei \rightarrow **compound nucleus**



N. Bohr,
Nature 137 ('36) 351



Wikipedia

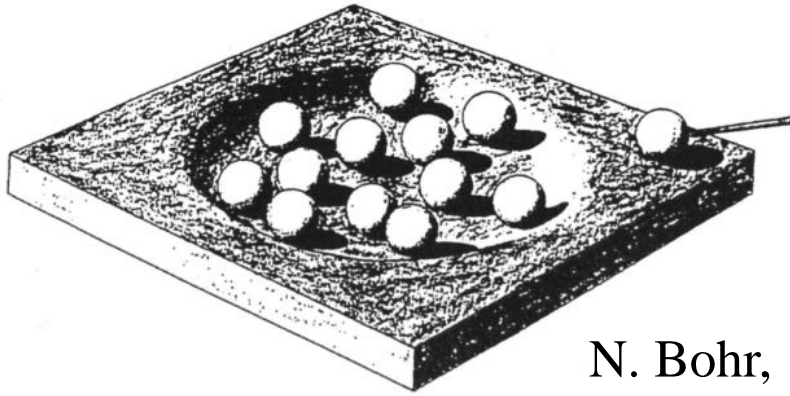


cf. Experiment of Enrico Fermi (1935)
many very narrow (=long life-time)
resonances (width \sim eV)

Fusion reactions: compound nucleus formation

Niels Bohr (1936)

Neutron capture of nuclei \rightarrow compound nucleus

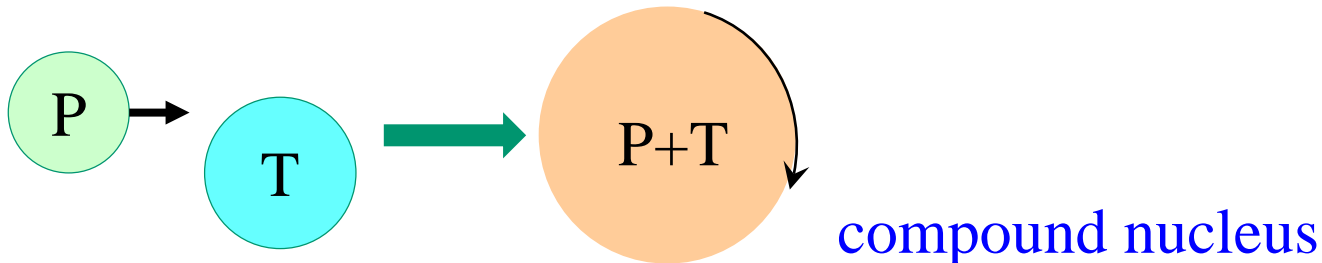


N. Bohr,
Nature 137 ('36) 351

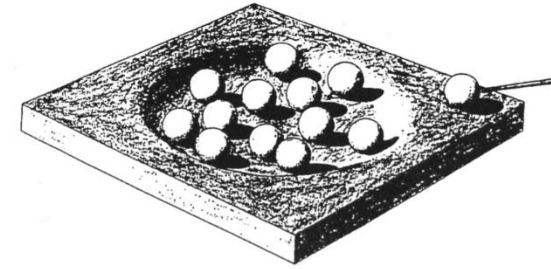
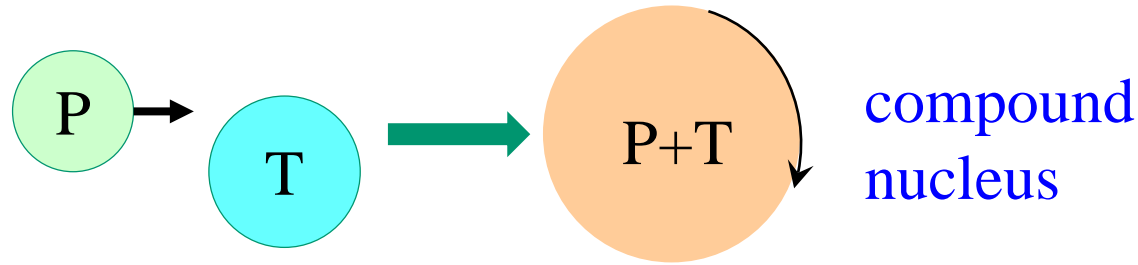


Wikipedia

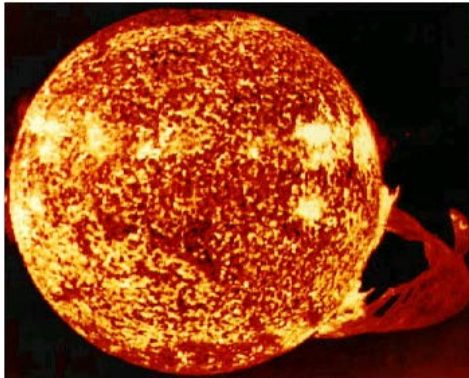
forming a compound nucleus with heavy-ion reactions = H.I. fusion



Fusion reactions: compound nucleus formation

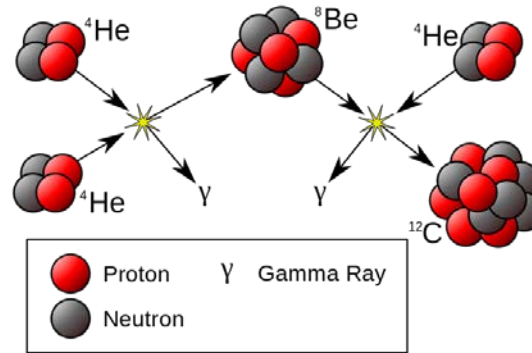


cf. Bohr '36

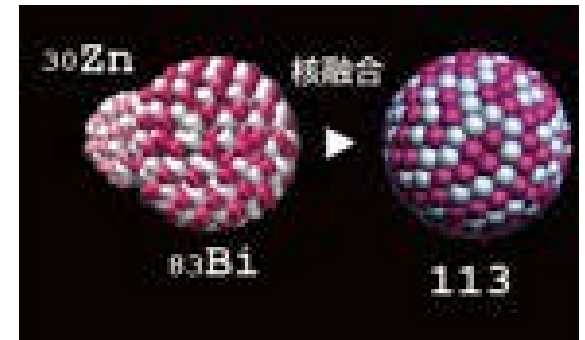


NASA, Skylab space station on December 19, 1973, solar flare reaching 589 000 km off solar surface

energy production
in stars (Bethe '39)



nucleosynthesis



superheavy elements

Fusion and fission: large amplitude motions of quantum many-body systems with strong interaction
 ← microscopic understanding: **an ultimate goal of nuclear physics**

Low-energy heavy-ion fusion reactions and quantum tunneling

✓ Reaction dynamics

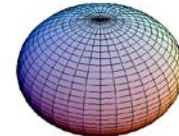
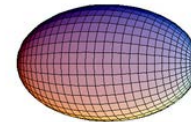
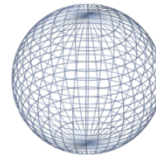
strong interplay between reaction and structure

cf. high E reactions: much simpler reaction mechanisms

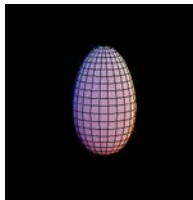
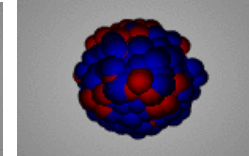
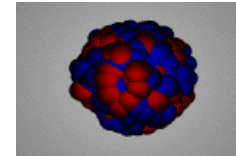
✓ Many-particle tunneling

cf. ▪ rich intrinsic motions

- several nuclear shapes



- several surface vibrations



several modes
and adiabaticities

Low-energy heavy-ion fusion reactions and quantum tunneling

✓ Reaction dynamics

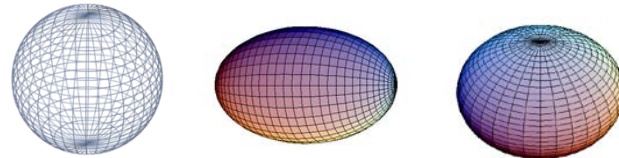
strong interplay between reaction and structure

cf. high E reactions: much simpler reaction mechanisms

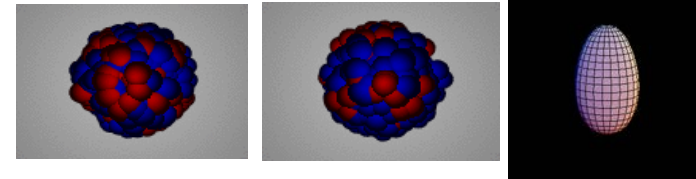
✓ Many-particle tunneling

cf. ▪ rich intrinsic motions

- several nuclear shapes



- several surface vibrations



- several types of nucleon transfers

“environment” can be changed relatively freely

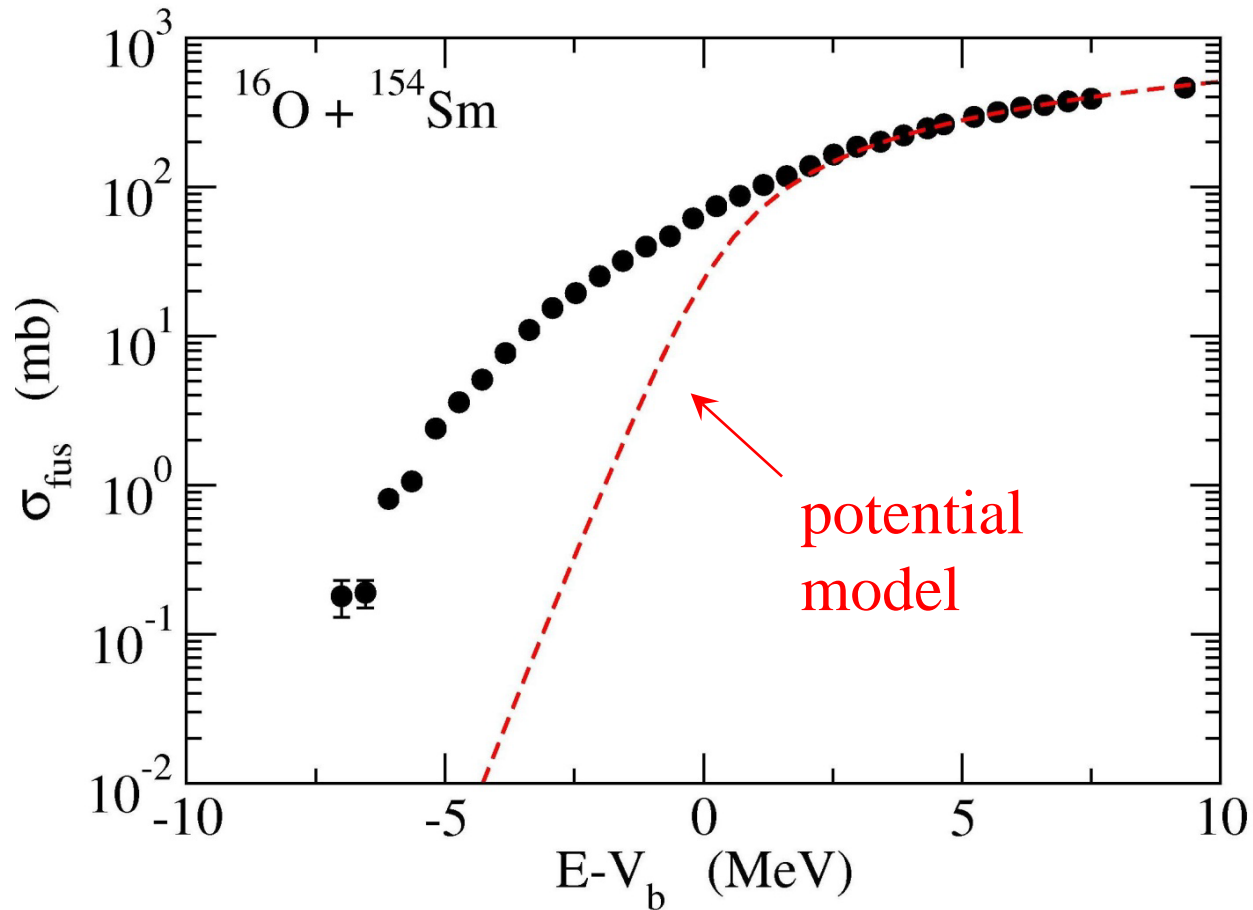
▪ E : variable cf. α decays: fixed energy

H.I. fusion reaction = an ideal playground to study quantum tunneling with many degrees of freedom

Discovery of large sub-barrier enhancement of σ_{fus} (~ 80 's)

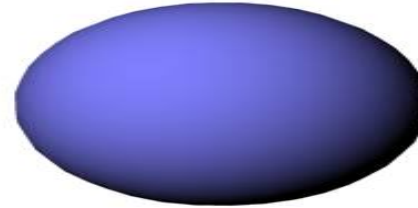
potential model: inert nuclei (no structure)

$$\sigma_{\text{fus}} = \frac{\pi}{k^2} \sum_l (2l + 1)(1 - |S_l|^2)$$

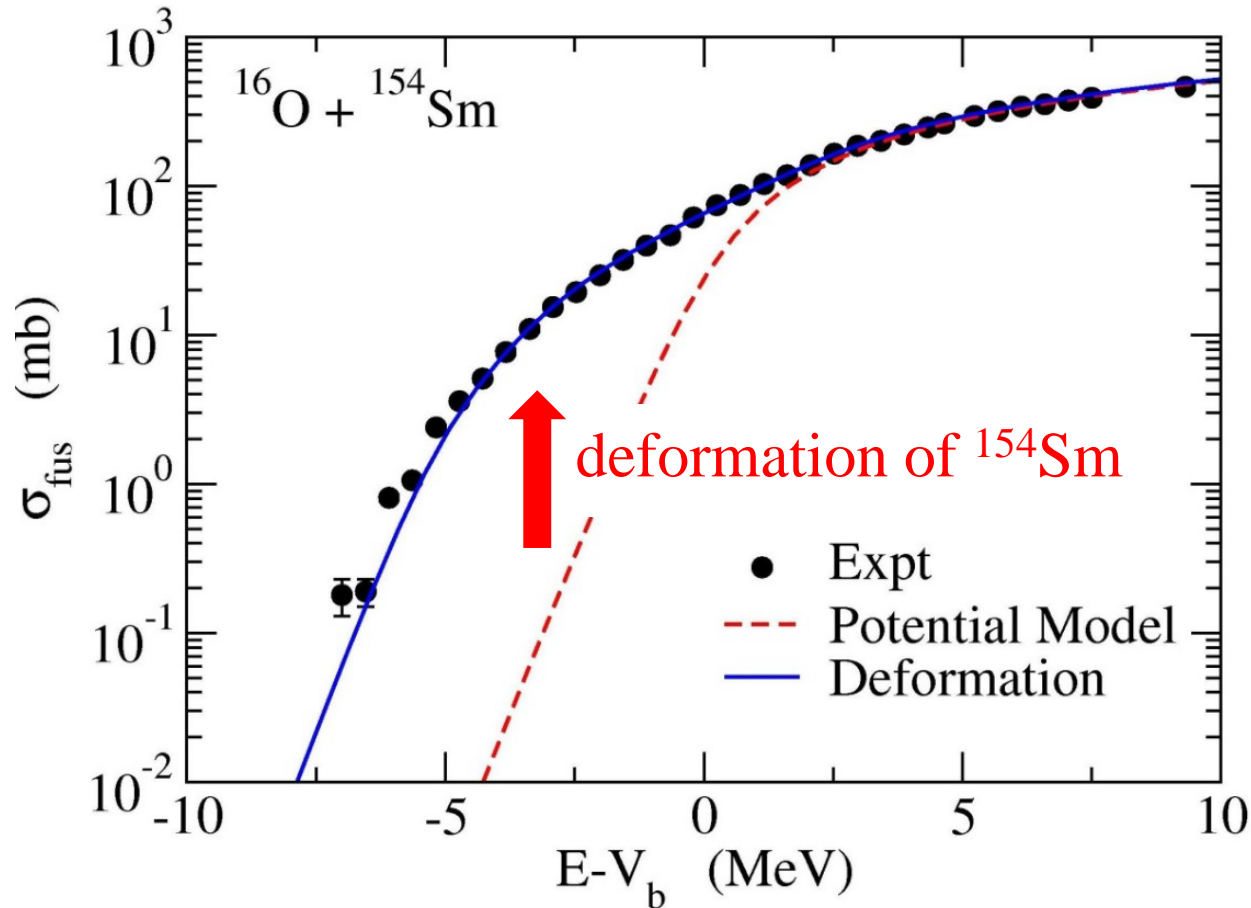


Discovery of large sub-barrier enhancement of σ_{fus} (~ 80 's)

^{154}Sm : a typical deformed nucleus

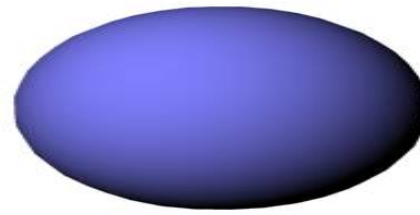


^{154}Sm

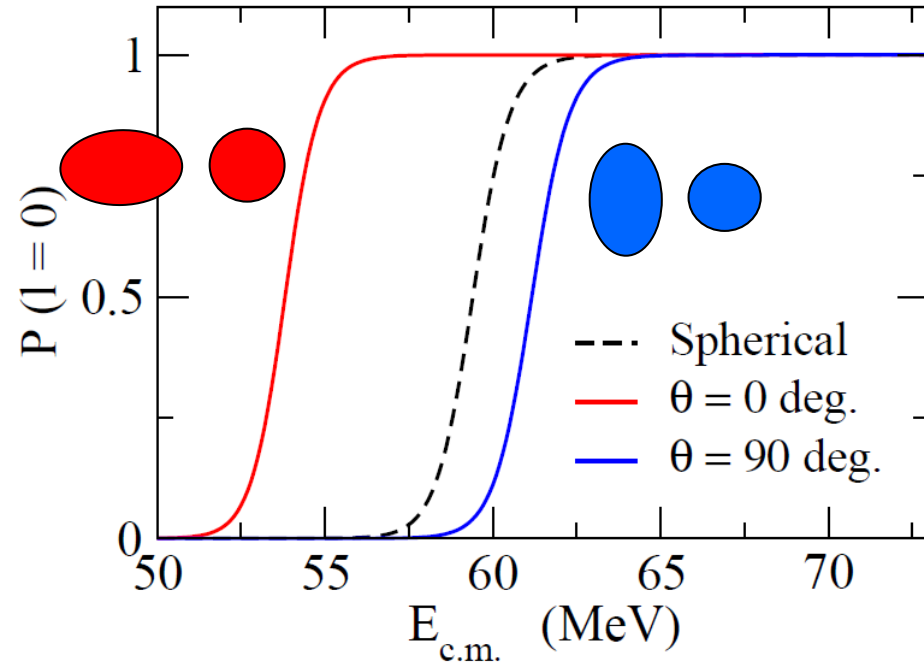
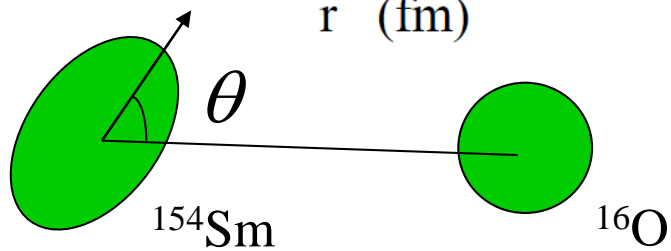
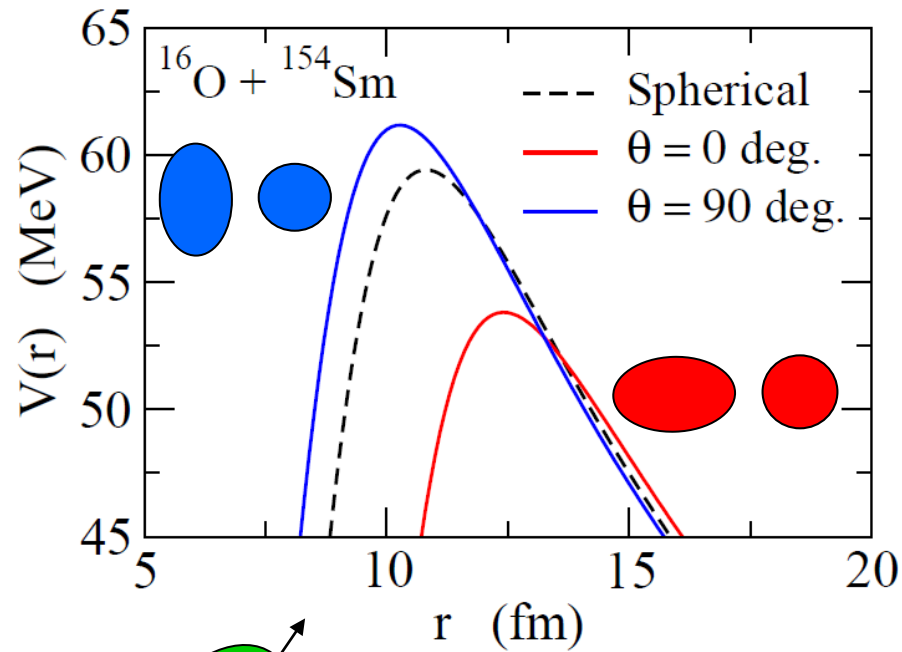


Effects of nuclear deformation

^{154}Sm : a typical deformed nucleus

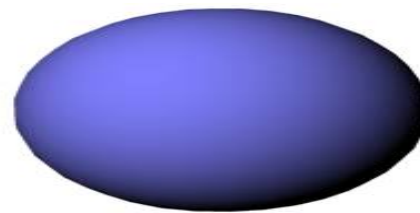


^{154}Sm

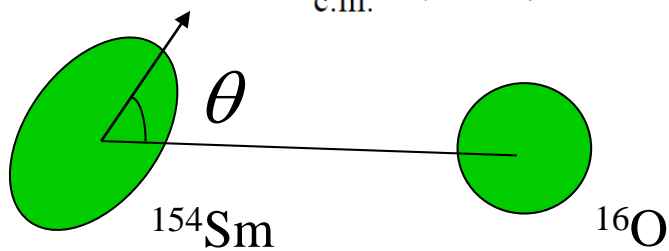
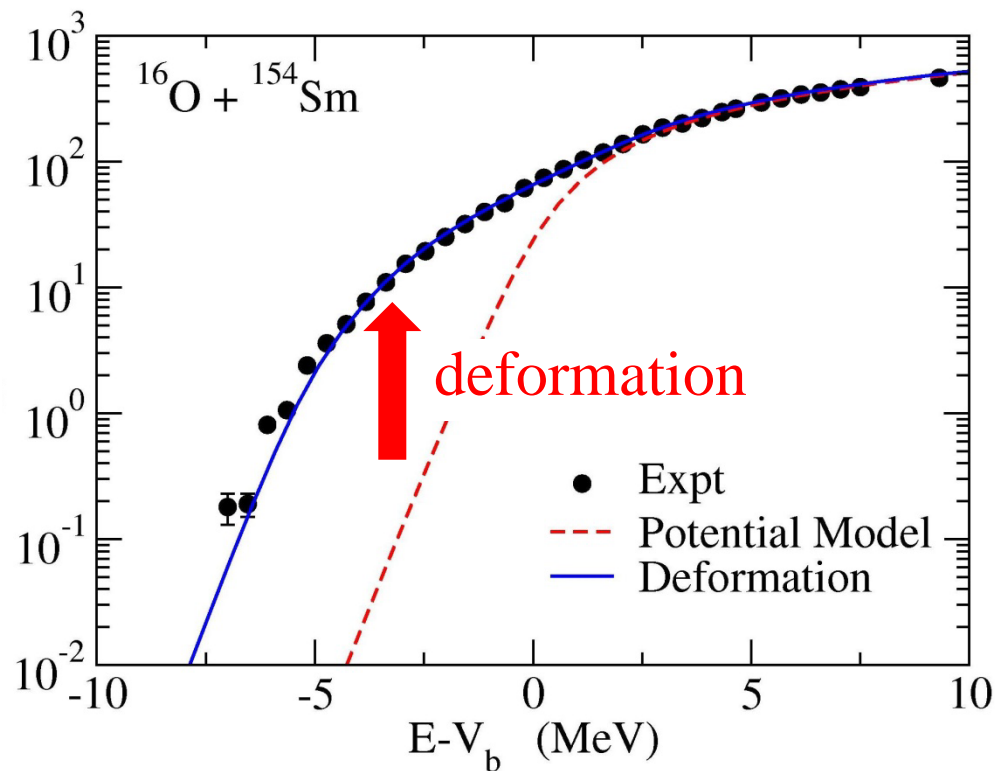
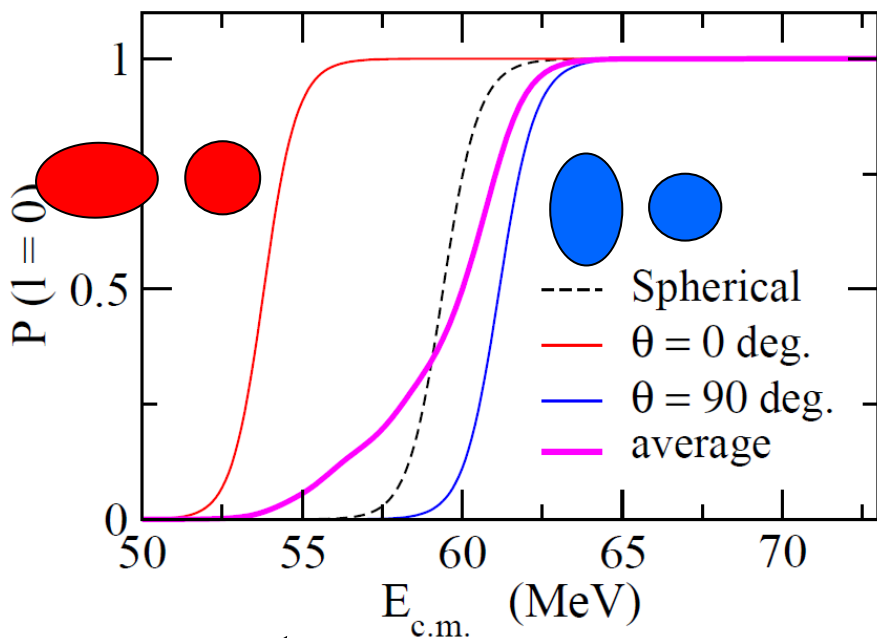


Effects of nuclear deformation

^{154}Sm : a typical deformed nucleus

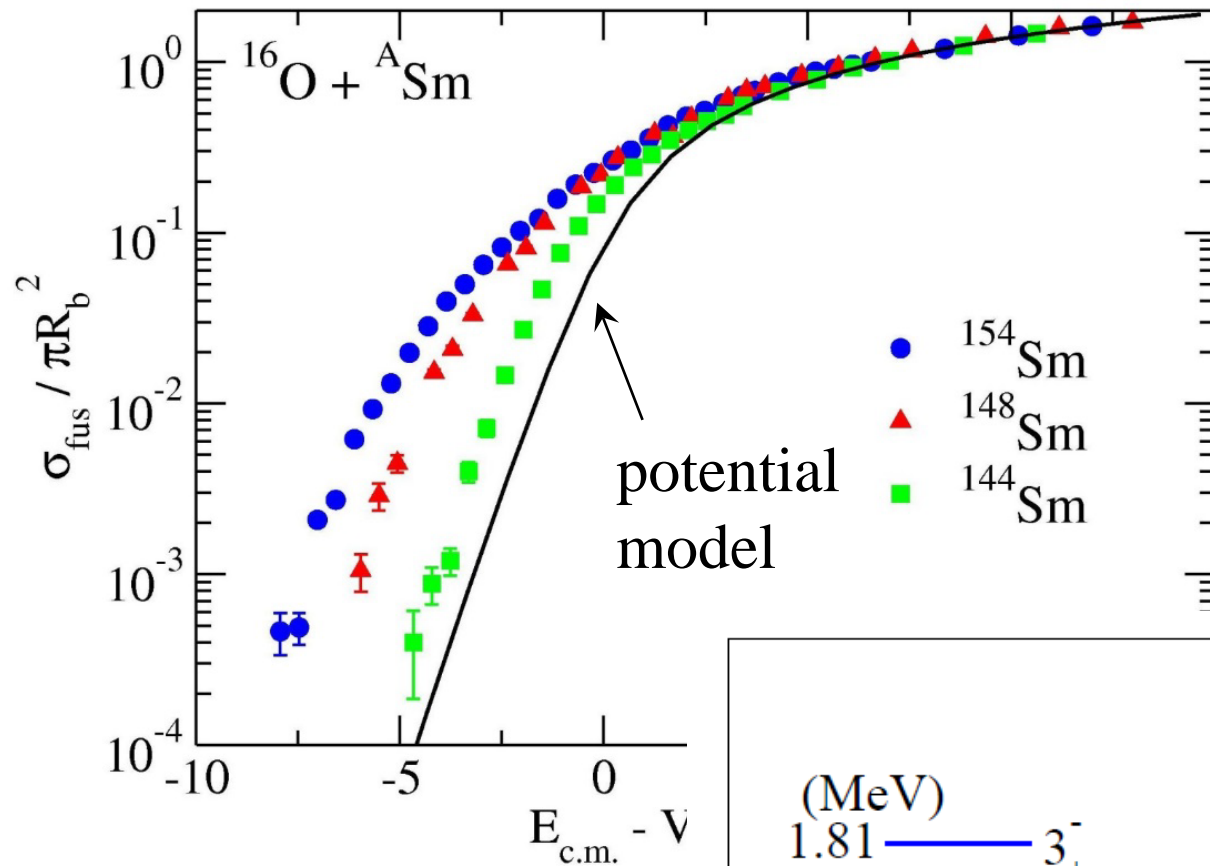


^{154}Sm



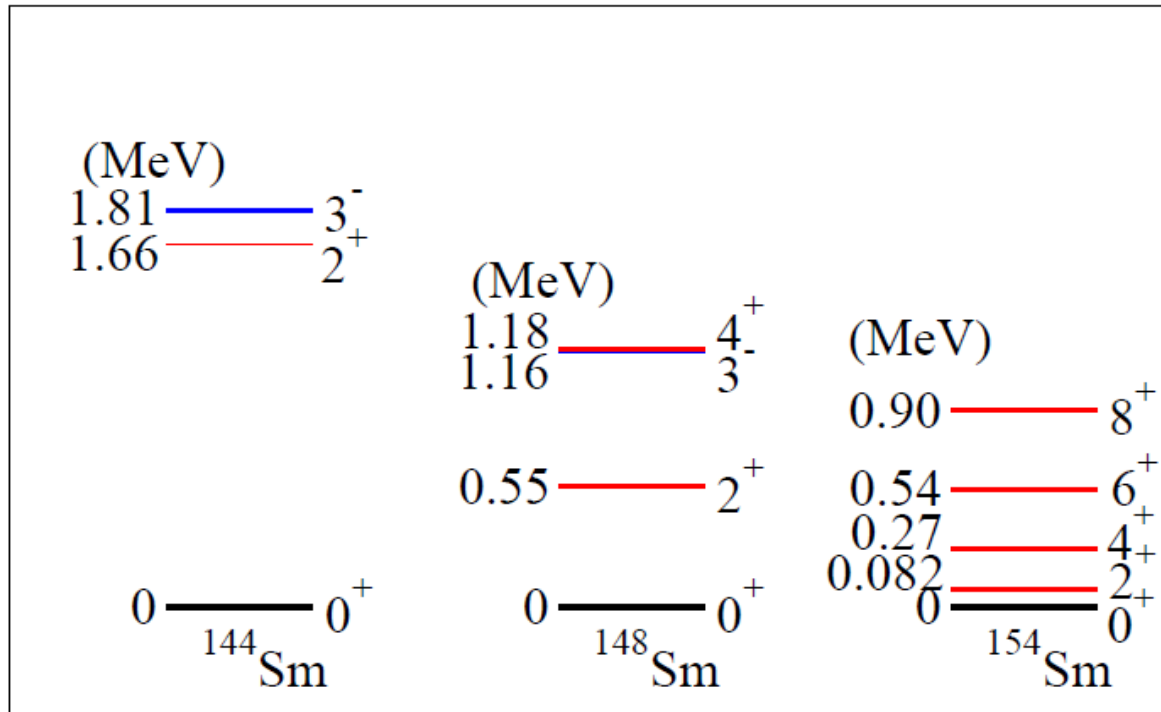
$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta) \sigma_{\text{fus}}(E; \theta)$$

Fusion: strong interplay between nuclear structure and reaction



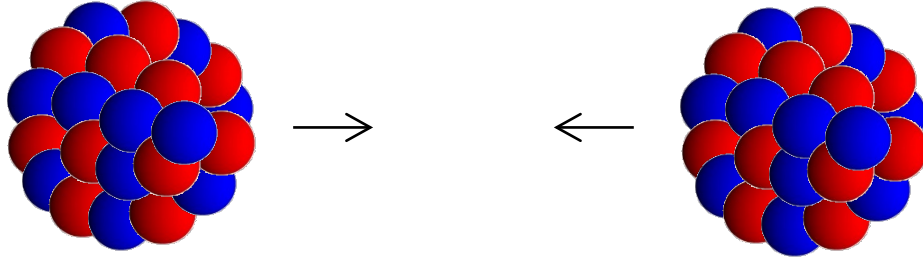
enhancement of fusion cross sections
: a general phenomenon

strong correlation with nuclear spectrum
→ coupling assisted tunneling



Coupled-channels method: a quantal scattering theory with excitations

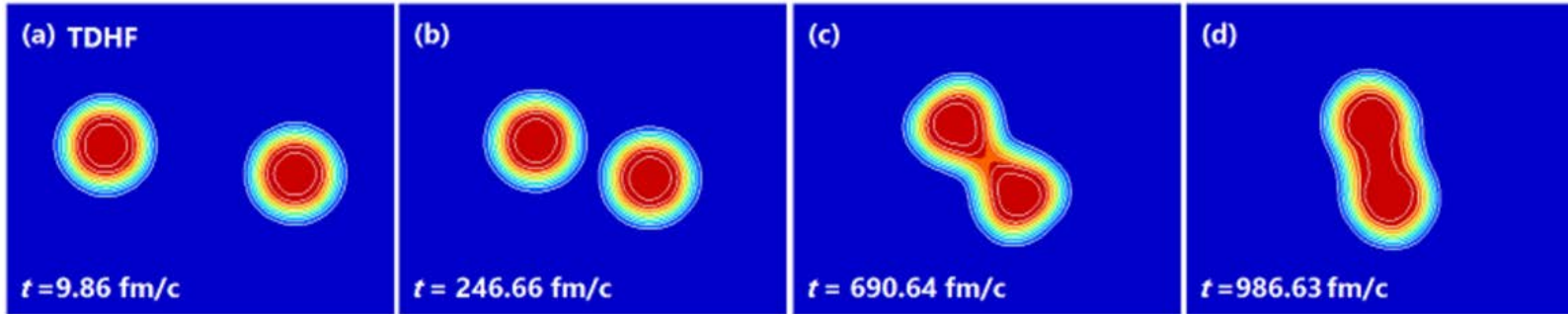
many-body problem



still very challenging

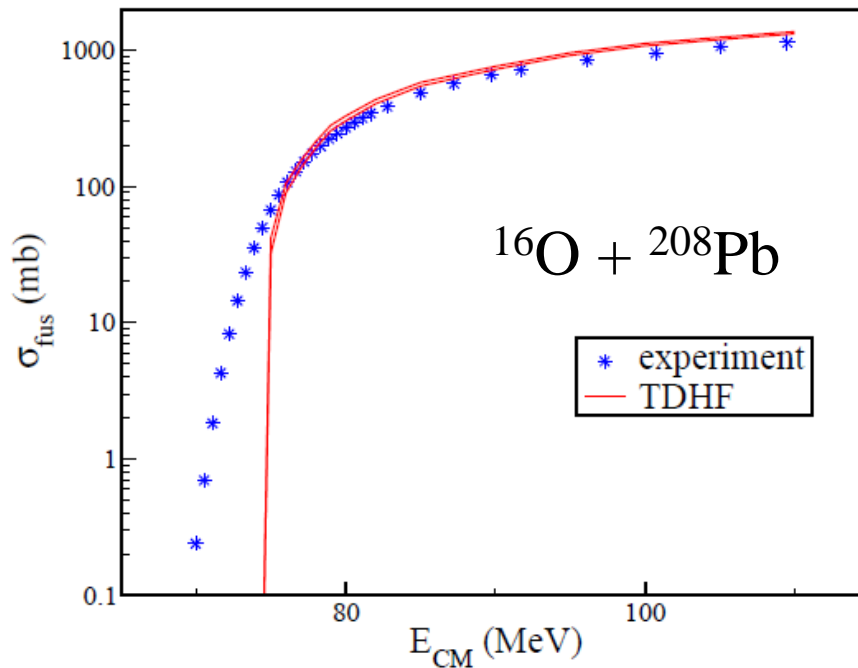
TDHF simulation

TDHF = Time Dependent Hartree-Fock
(a single Slater determinant)



S. Ebata, T. Nakatsukasa, JPC Conf. Proc. 6 ('15) 020056

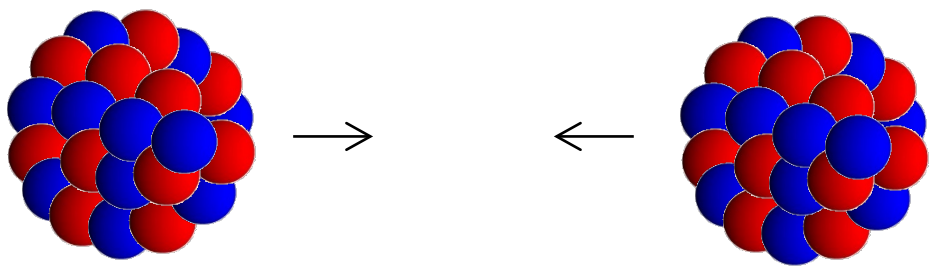
ab-initio, but no tunneling



C. Simenel,
EPJA48 ('12) 152

Coupled-channels method: a quantal scattering theory with excitations

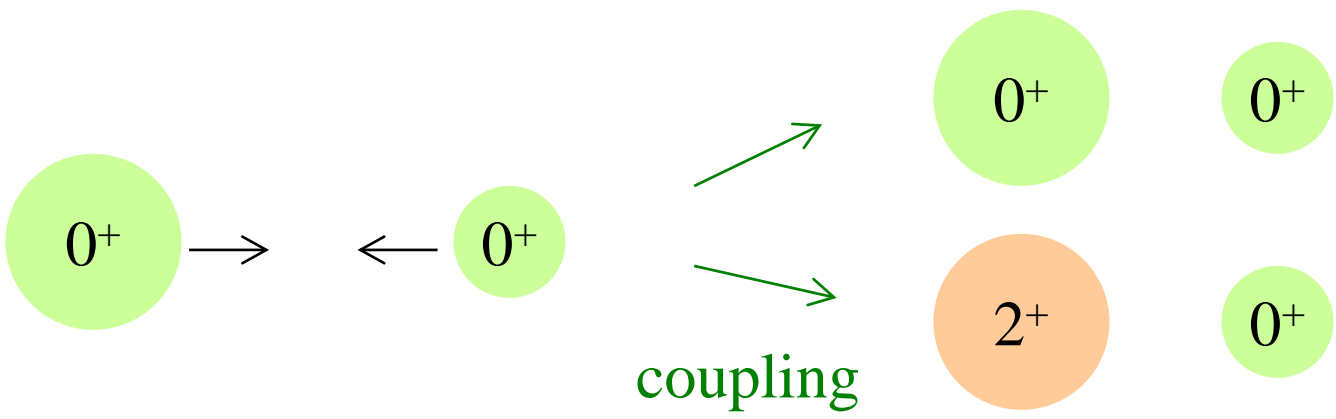
many-body problem



still very challenging

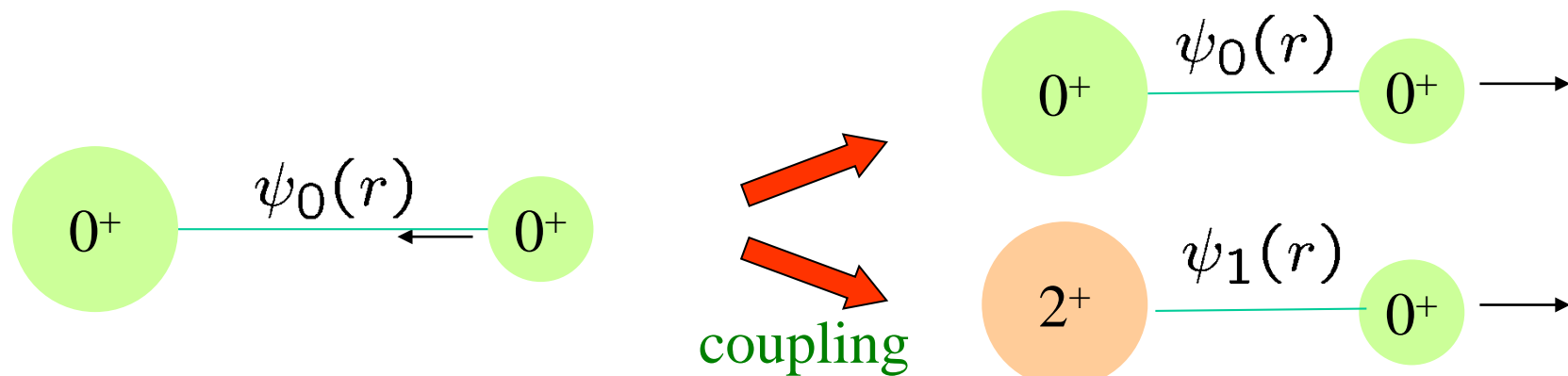


two-body problem, but with excitations
(coupled-channels approach)

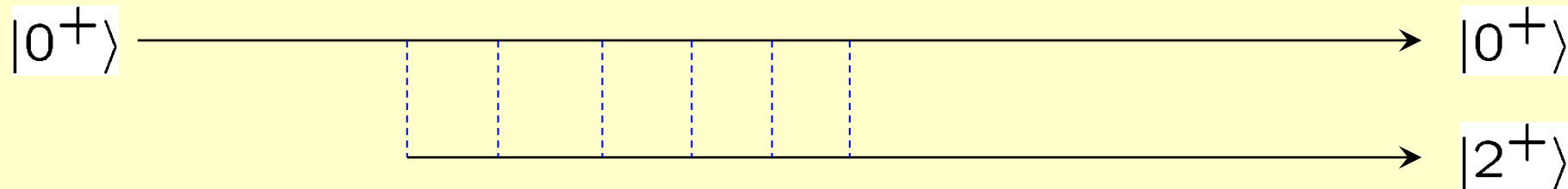


scattering theory with excitations

Coupled-channels method: a quantal scattering theory with excitations



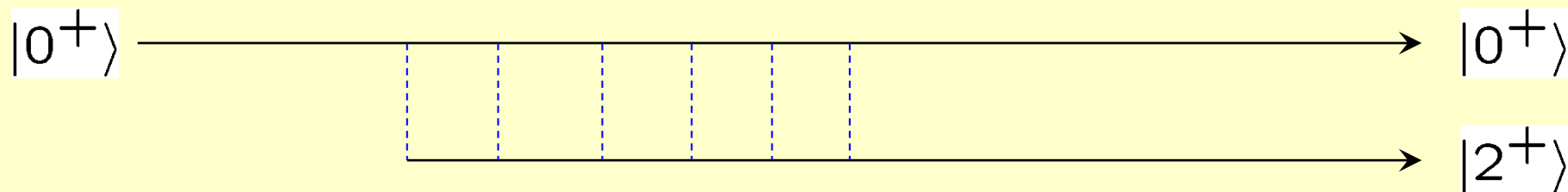
$$\left[-\frac{\hbar^2}{2\mu} \nabla^2 + \overleftarrow{V}(r) - \overleftarrow{E} \right] \overrightarrow{\psi}(r) = 0$$



dynamics of excitations/de-excitations during reaction

- ✓ Non-perturbative (full order)
- ✓ Non-adiabatic (excitation energy)

Coupled-channels method: a quantal scattering theory with excitations



dynamics of excitations/de-excitations during reaction

- ✓ Non-perturbative (full order)
- ✓ Non-adiabatic (excitation energy)

in the past, the linear coupling approximation in a Hamiltonian :

$$H = -\frac{\hbar^2}{2\mu} \nabla^2 + V_0(r) + \beta f(r) \hat{O}$$

full order treatment

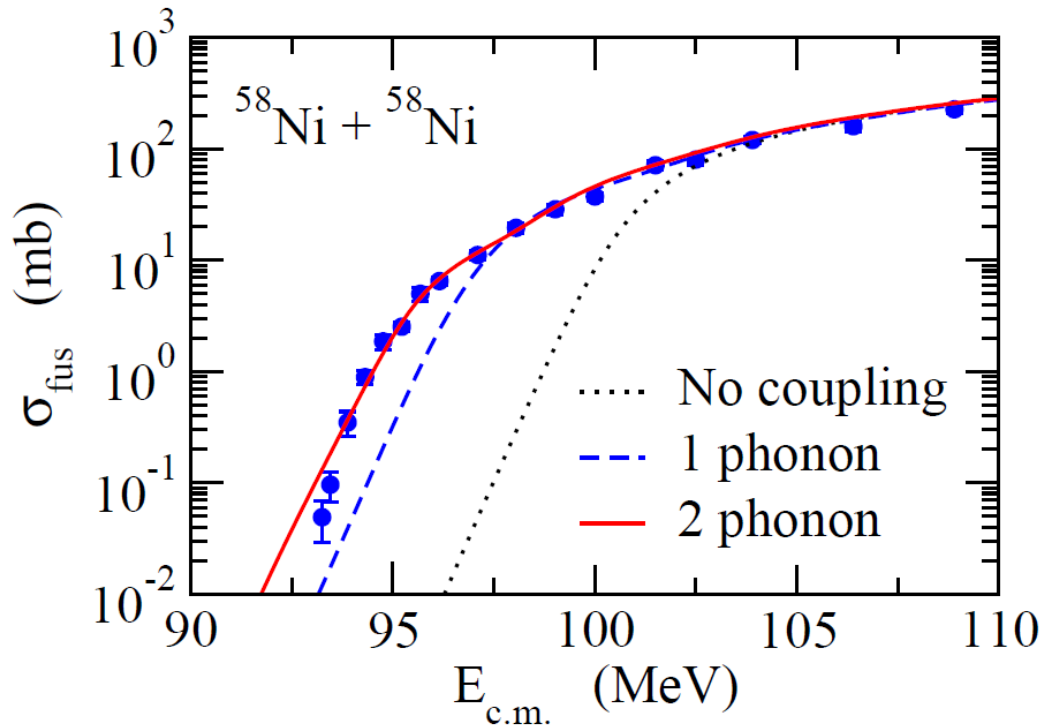
$$V(r, \beta \hat{O})$$

K.H., N. Rowley, and A.T. Kruppa,
Comp. Phys. Comm. 123('99) 143.

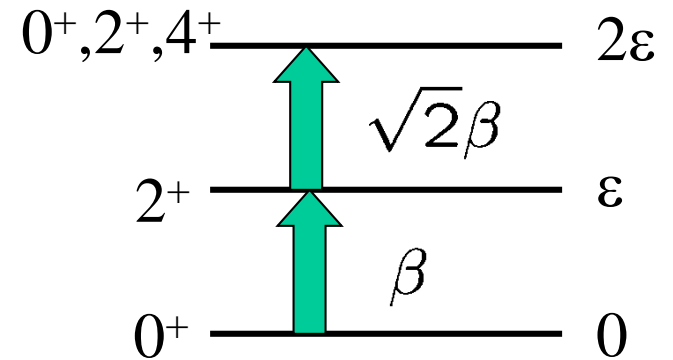
a standard
method

modellings of coupled-channels calculations

- K.H., N. Rowley, A.T. Kruppa, CPC123 ('99) 143
- M. Zamrun, K.H., S. Mitsuoka, H. Ikezoe, PRC77 ('08) 034604
- T. Ichikawa, K.H., A. Iwamoto, PRL103 ('09) 202701
- S. Yusa, K.H., and N. Rowley, PRC88 ('13) 044620
- J.M. Yao and K.H., PRC94 ('16) 11303(R) etc.



harmonic oscillations

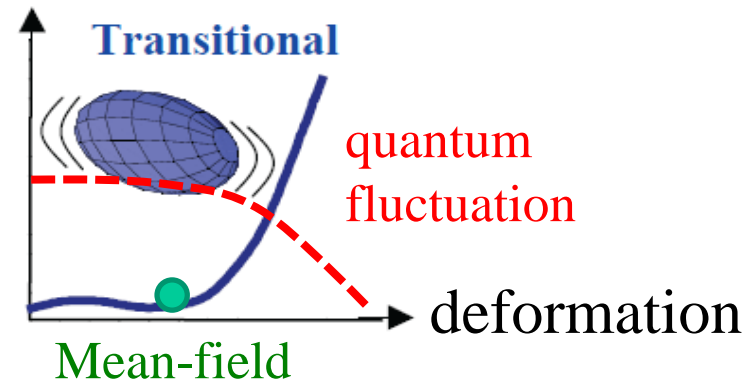


review : K. Hagino and N. Takigawa, Prog. Theo. Phys.128 ('12)1061.

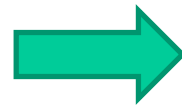
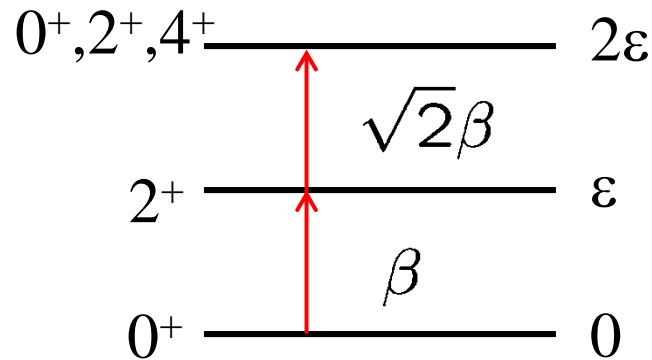
Further development: semi-microscopic modelling

K.H. and J.M. Yao, PRC91('15) 064606

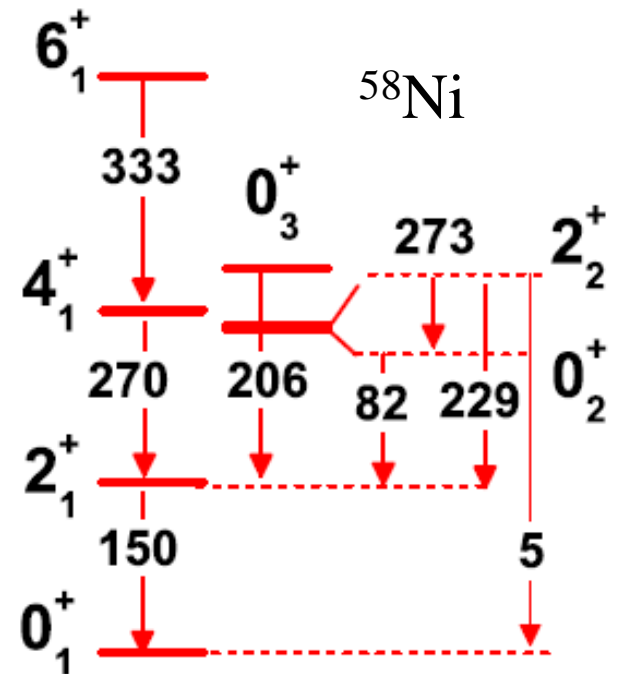
Coupled-channels
+ microscopic nuclear structure
calculations
(GCM, Shell Model, IBM.....)



simple harmonic
oscillator

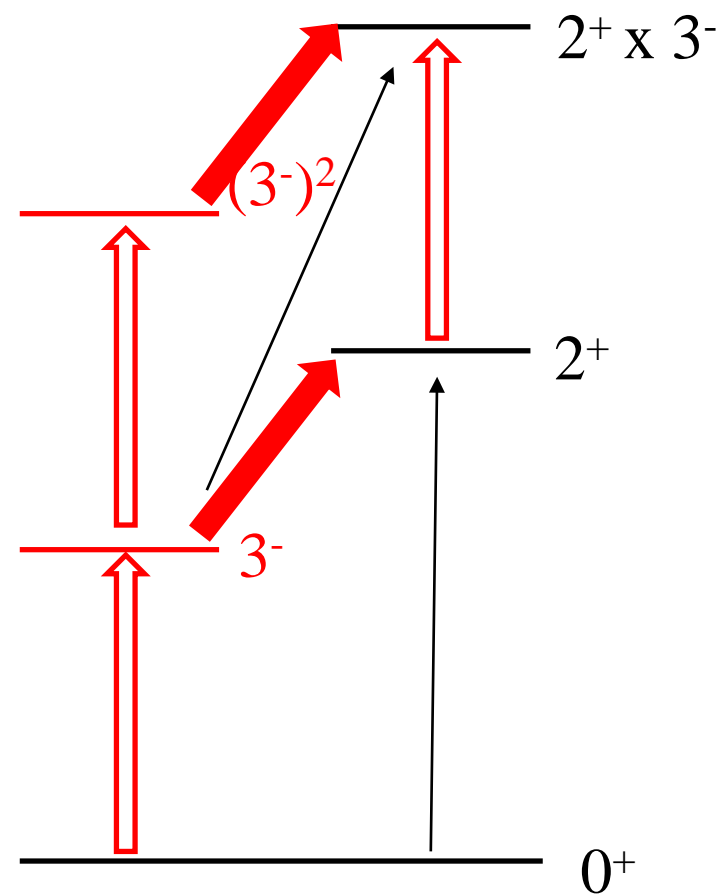
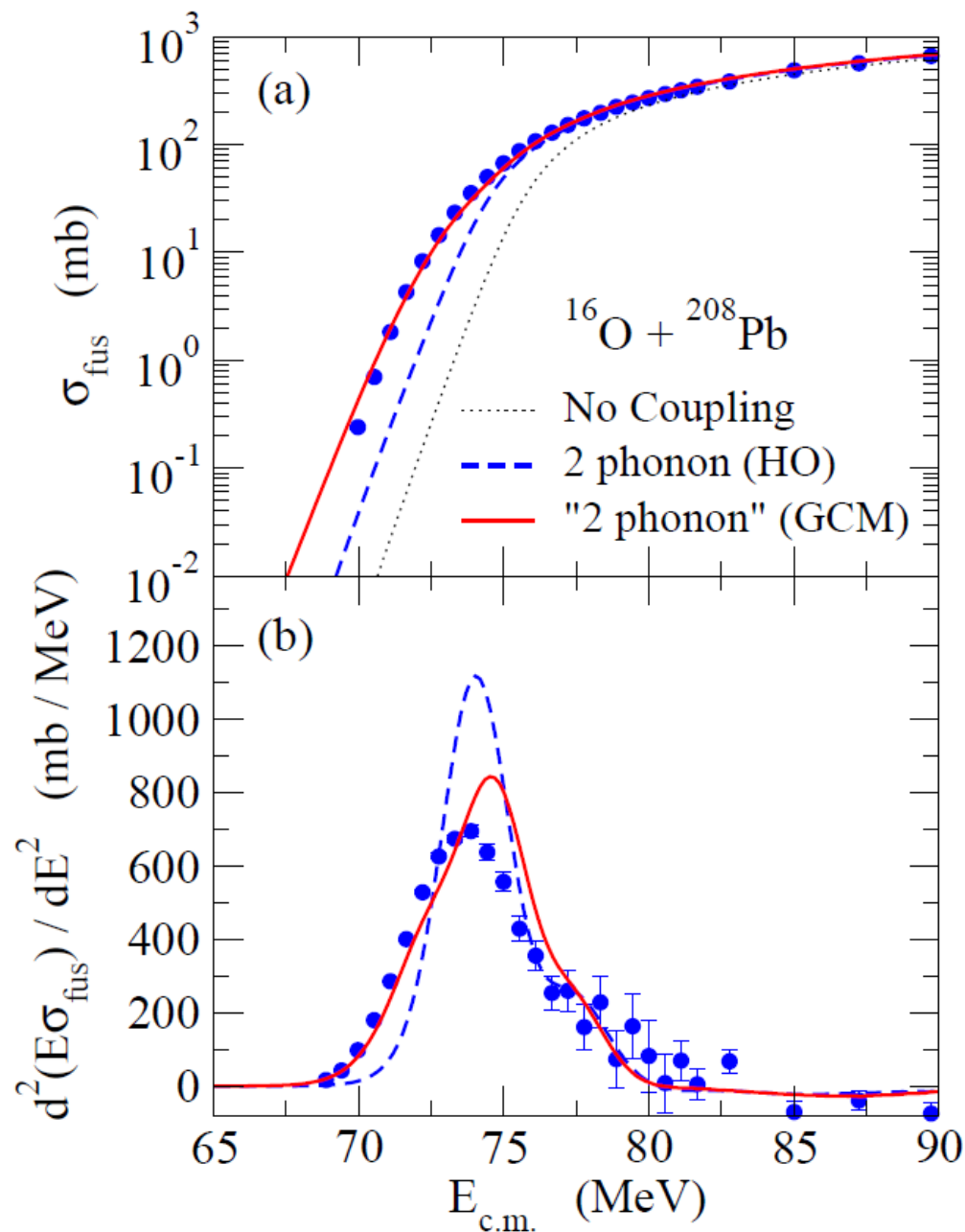


anharmonicity
in phonon spectra



relativistic MF + GCM

Relativistic Mean-Field + Quantum fluctuation + coupled-channels



J.M. Yao and K.H.,
PRC94 ('16) 11303(R)

From phenomenological approach to microscopic approach

Macroscopic (phenomenological)

C.C. with collective model

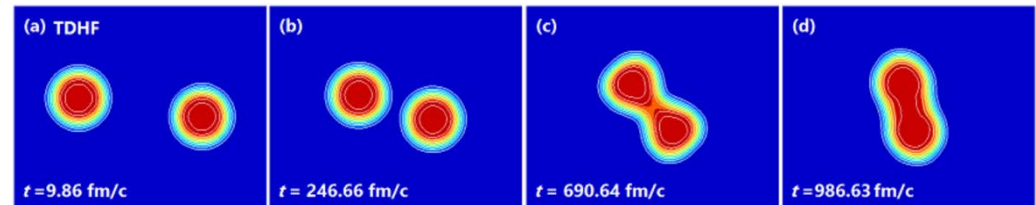
C.C. with inputs from
microscopic nuclear
structure calculations

C.C. with inputs based
on TDHF

TDHF simulations

Microscopic

TDHF = Time Dependent Hartree-Fock



S. Ebata, T. Nakatsukasa, JPC Conf. Proc. 6 ('15)

ab initio, but no tunneling

From phenomenological approach to microscopic approach

TDHF simulations

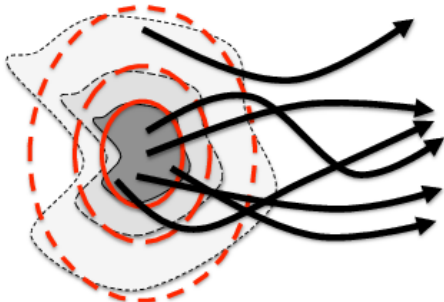
ab initio, but no tunneling

➤ “Beyond mean-field” approximations

✓ Time-dependent GCM?

a single Slater determinant (SD) to multi-SD

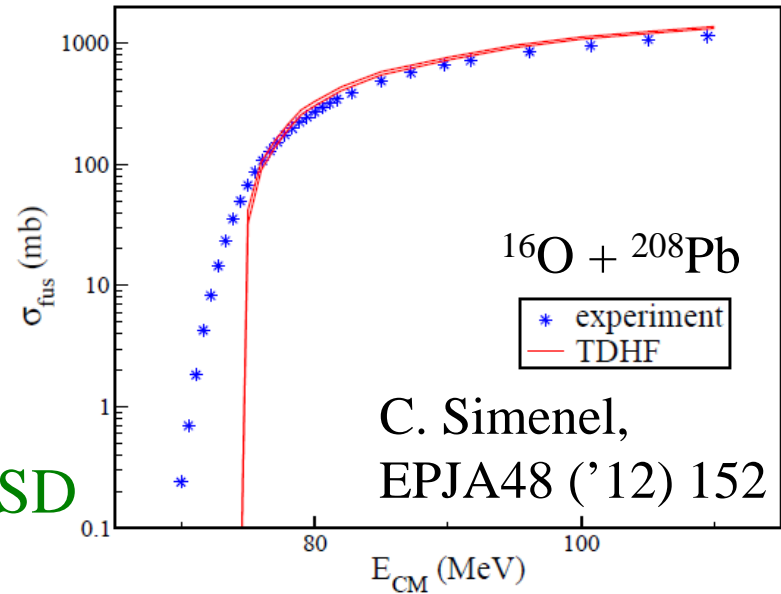
$$|\Psi(t)\rangle = \int dq f(q, t) |\Phi_q(t)\rangle$$



dynamics with a superposition of many
“TDHF trajectories (Slater determinants)”

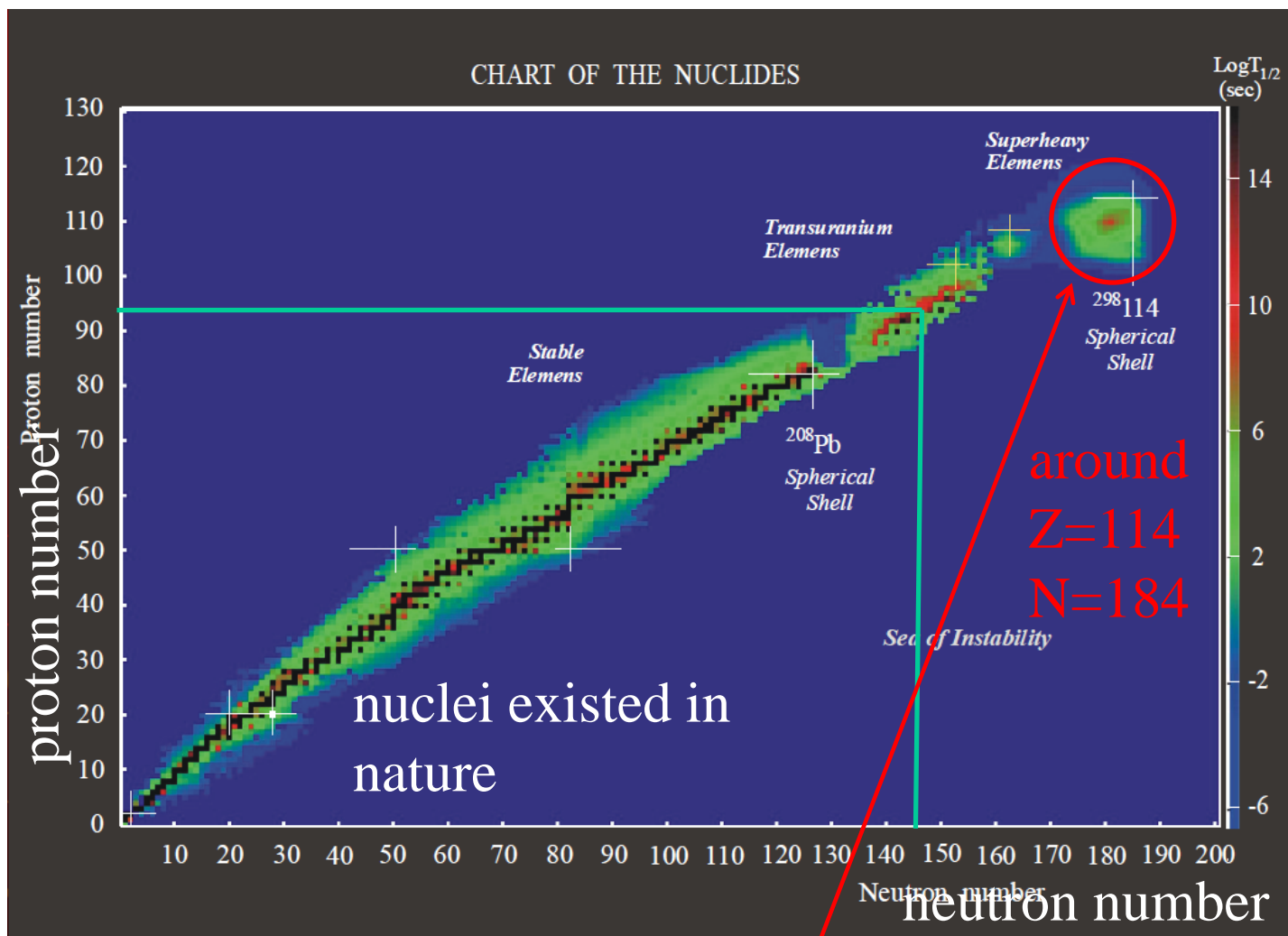
an open problem

K.H., N. Hasegawa, and Y. Tanimura,
a work in progress



cf. Stochastic mean-field method
B. Yilmaz et al.,
PRC90 ('14) 054617

Future perspectives: fusion for superheavy elements



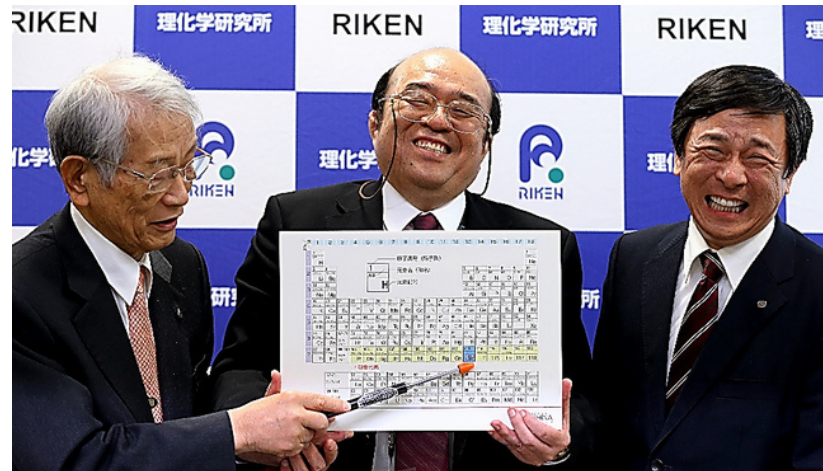
a prediction of island of stability
(Swiatecki et al., 1966)

Yuri Oganessian

Fusion reactions for SHE

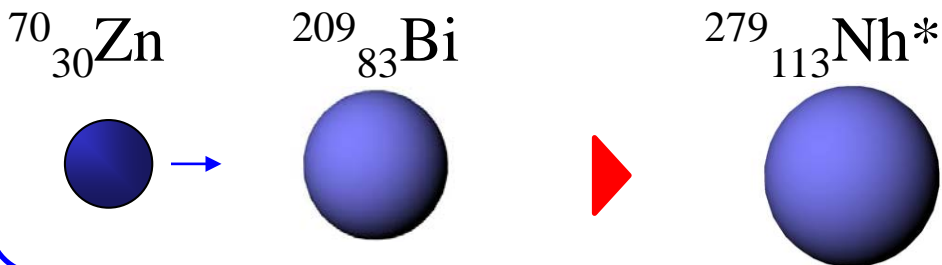
the element 113: Nh

113 Nh nihonium	115 Mc moscovium
117 Ts tennessine	118 Og oganesson



Group	1	2	3											11	12	13	14	15	16	17	18
Period																					
1	1 H																	2 He			
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne			
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar			
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr			
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe			
6	55 Cs	56 Ba	57 La*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn			
7	87 Fr	88 Ra	89 Ac*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og			
* 58 Ce 59 Pr 60 Nd 61 Pm 62 Sm 63 Eu 64 Gd 65 Tb 66 Dy 67 Ho 68 Er 69 Tm 70 Yb 71 Lu																					
* 90 Th 91 Pa 92 U 93 Np 94 Pu 95 Am 96 Cm 97 Bk 98 Cf 99 Es 100 Fm 101 Md 102 No 103 Lr																					

November, 2016

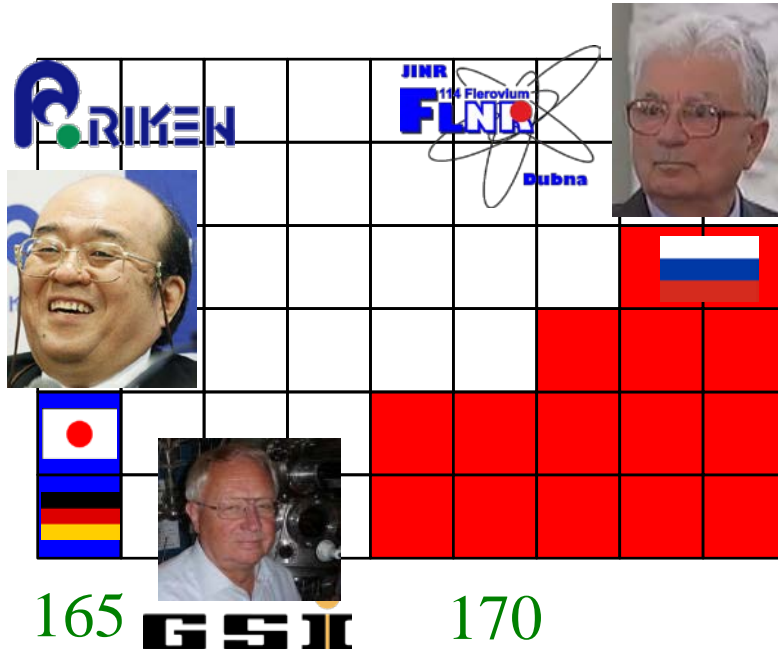


Heavy-ion fusion reaction

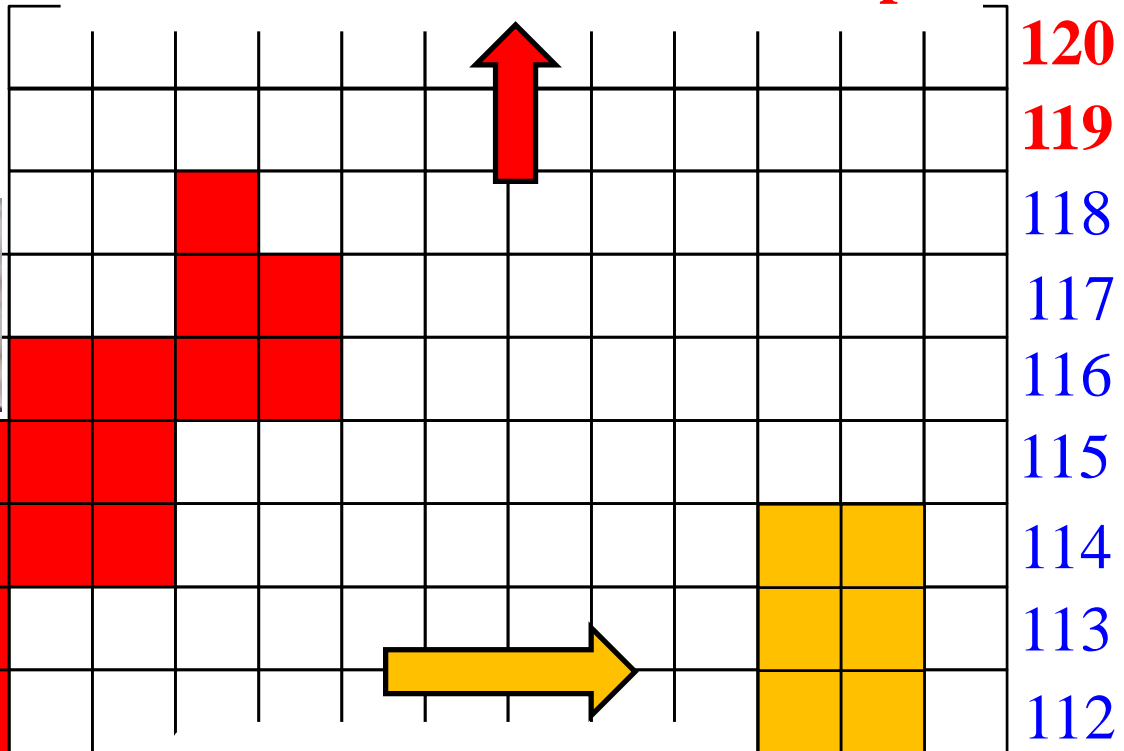
Wikipedia

Future directions of SHE

Superheavy elements synthesized so far



Towards Z=119 and 120 isotopes



Towards the island of stability

↑ 185
the island of stability?

➤ Towards Z=119 and 120 isotopes

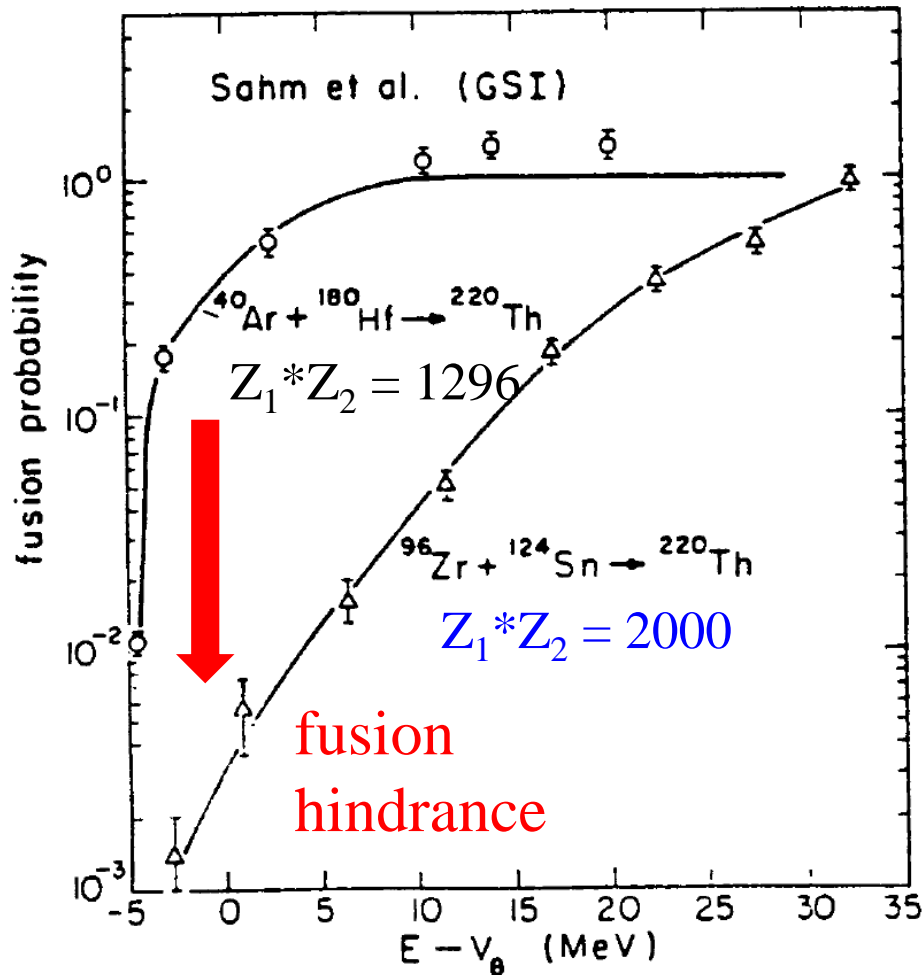
Hot fusion reactions with ^{48}Ca , $^{50}_{22}\text{Ti}$, $^{51}_{23}\text{V}$, $^{54}_{24}\text{Cr}$ etc.

➤ Towards the island of stability

neutron-rich beams: indispensable → reaction dynamics?

Fusion reactions in the SHE region ($Z_P * Z_T > 1600 \sim 1800$)

fusion hindrance



C.-C. Sahm et al., Z. Phys. A319('84)113

superheavy nuclei

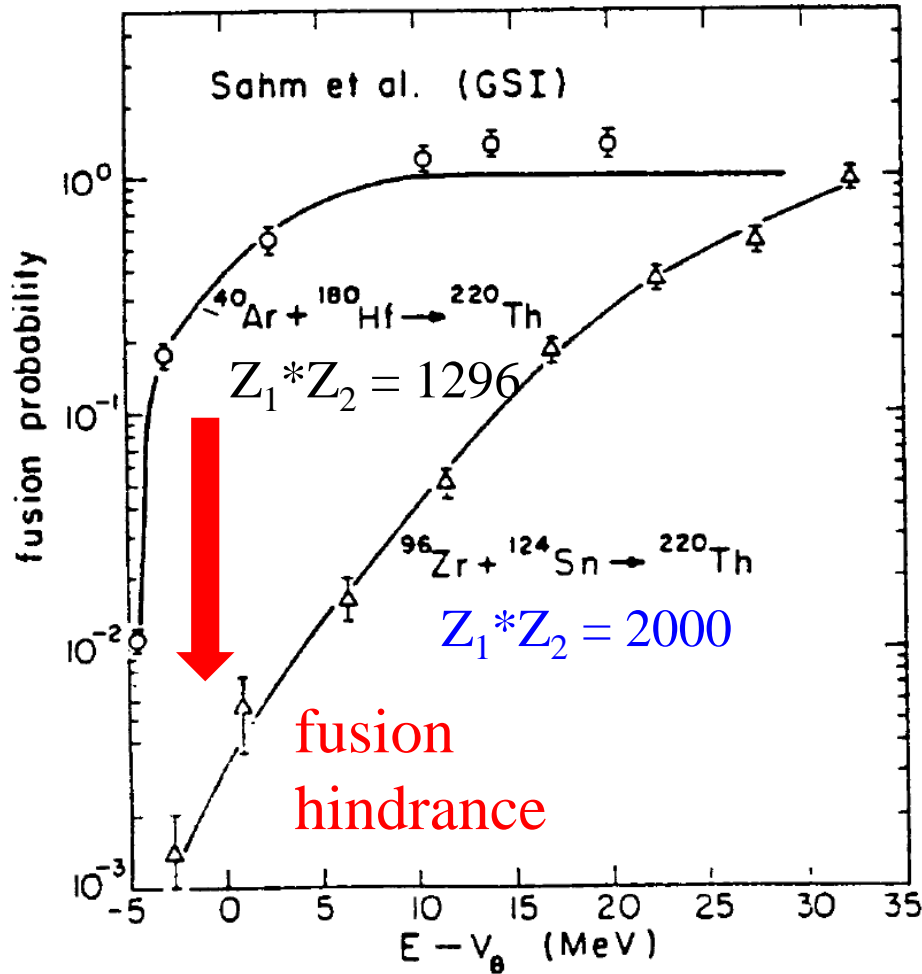
113 Nh nihonium	115 Mc moscovium
117 Ts tennessine	118 Og oganeson

syntheses with heavy-ion fusion reactions

theoretical issues:
understanding the reaction dynamics

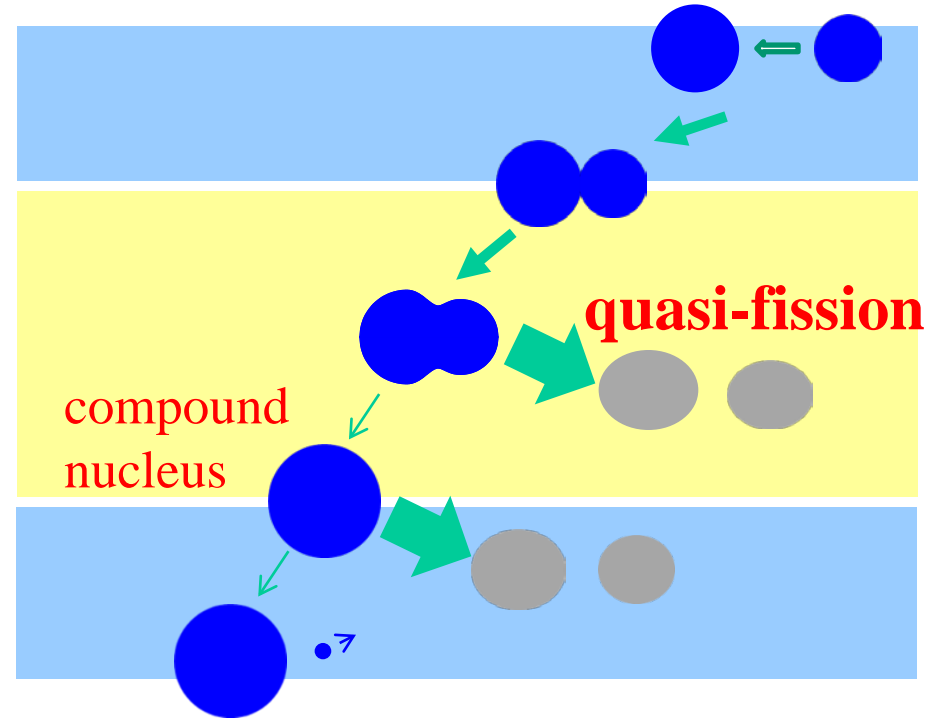
Fusion reactions in the SHE region ($Z_P * Z_T > 1600 \sim 1800$)

fusion hindrance



C.-C. Sahm et al., Z. Phys. A319('84)113

modern interpretation of hindrance



strong Coulomb repulsion
 \rightarrow re-separation before the
 compound nucleus

Fusion reactions in the SHE region ($Z_P * Z_T > 1600 \sim 1800$)

SHE formation: a very rare event

→ large theoretical uncertainties

✓ No data for P_{CN}

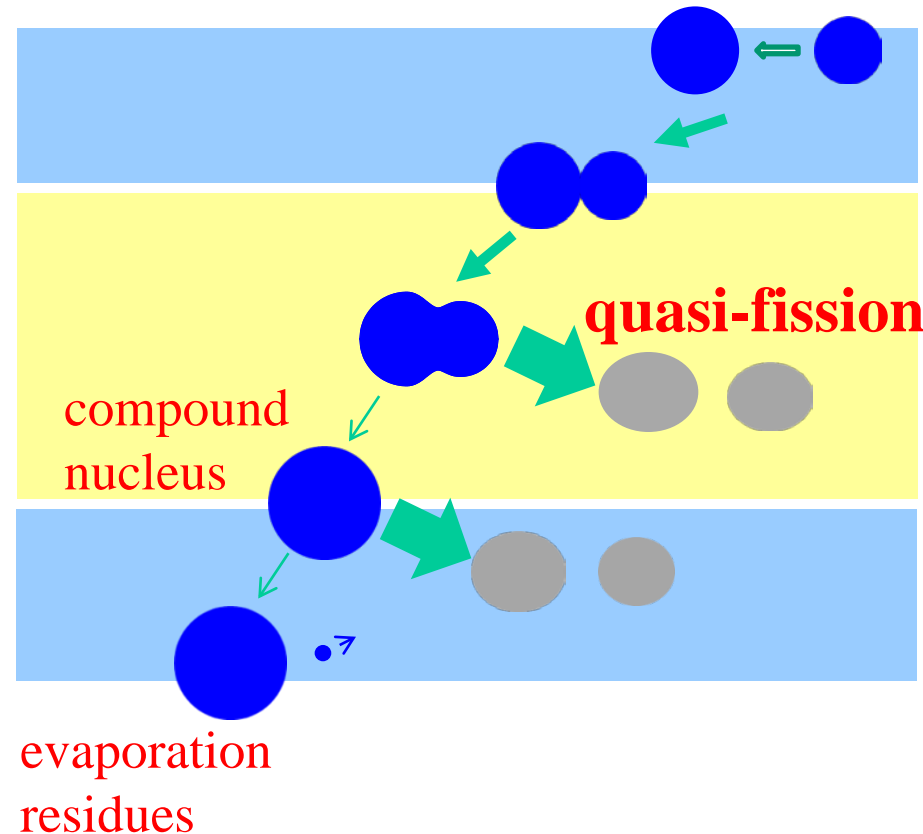
✓ Data: only for P_{ER}

CN = compound nucleus

ER = evaporation residues

**theoretical challenge:
to reduce theoretical uncertainties
and make a reliable prediction**

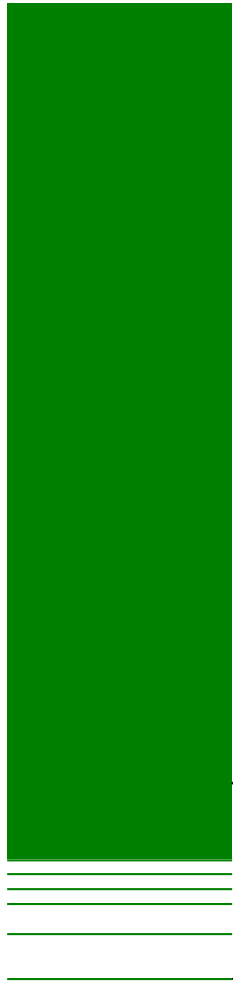
modern interpretation of hindrance



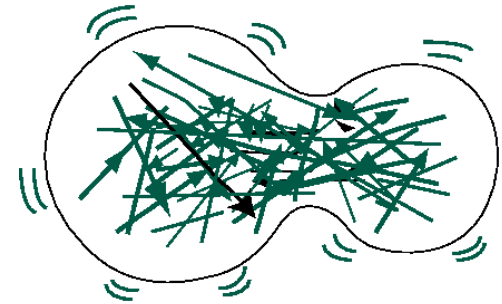
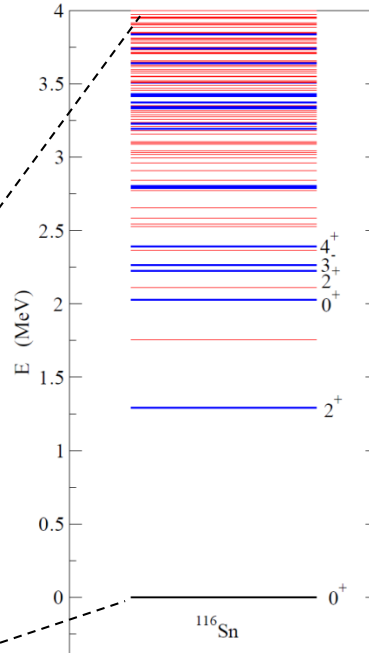
strong Coulomb repulsion
→ re-separation before the
compound nucleus

Nuclear friction and heavy-ion fusion reactions

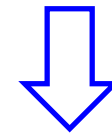
E^*



$$\rho(E) \sim e^{2\sqrt{aE^*}}$$



These states:
are excited in a complicated way.



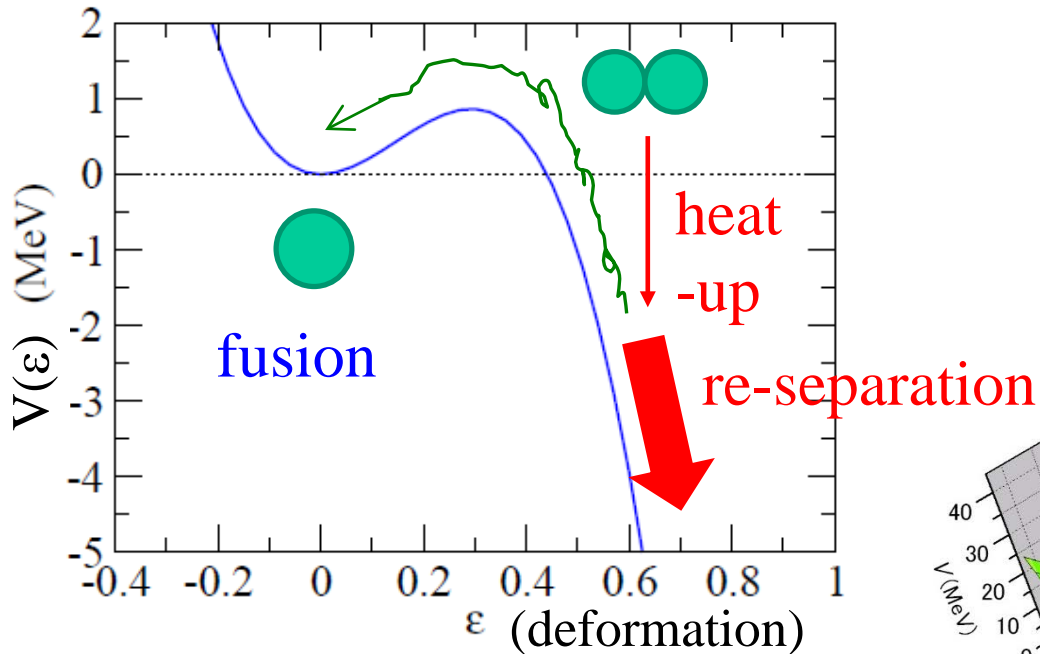
nuclear intrinsic d.o.f.
: act as environment

“intrinsic environment”

→ friction

nuclear spectrum

Langevin approach



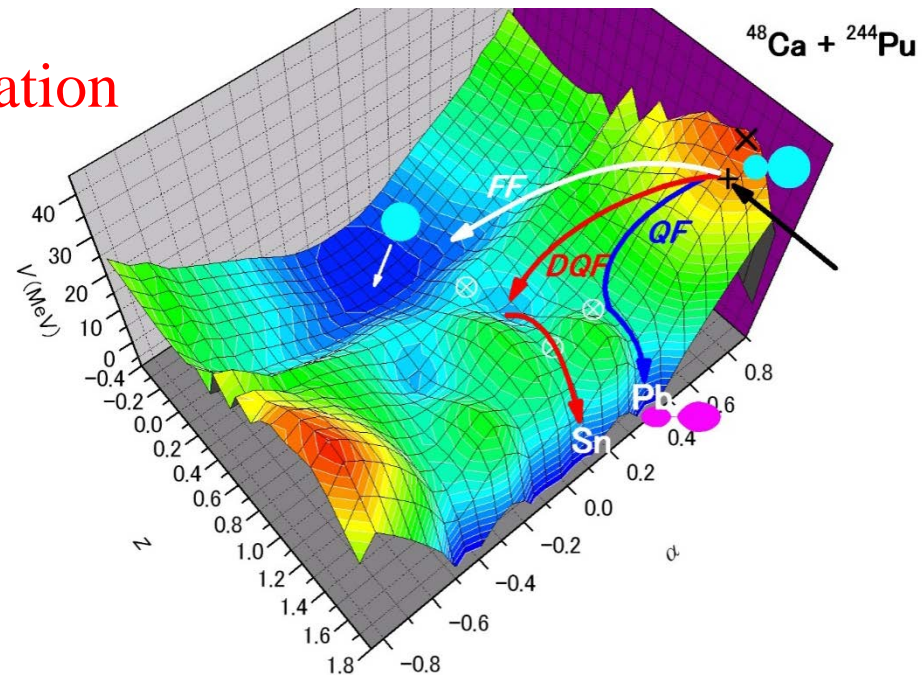
thermal diffusion

→ Langevin approach
(Brownian motion)

$$m \frac{d^2 q}{dt^2} = - \frac{dV(q)}{dq} - \gamma \frac{dq}{dt} + R(t)$$

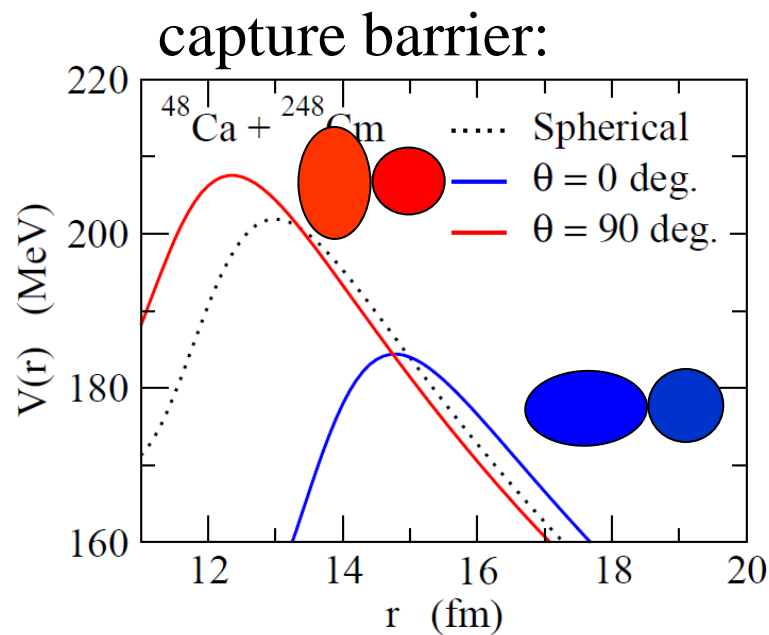
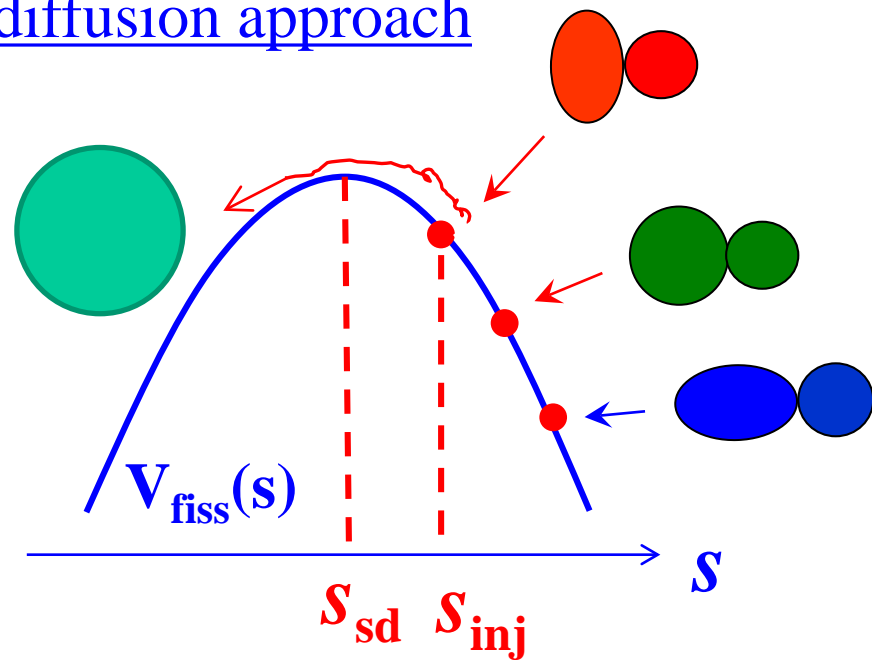
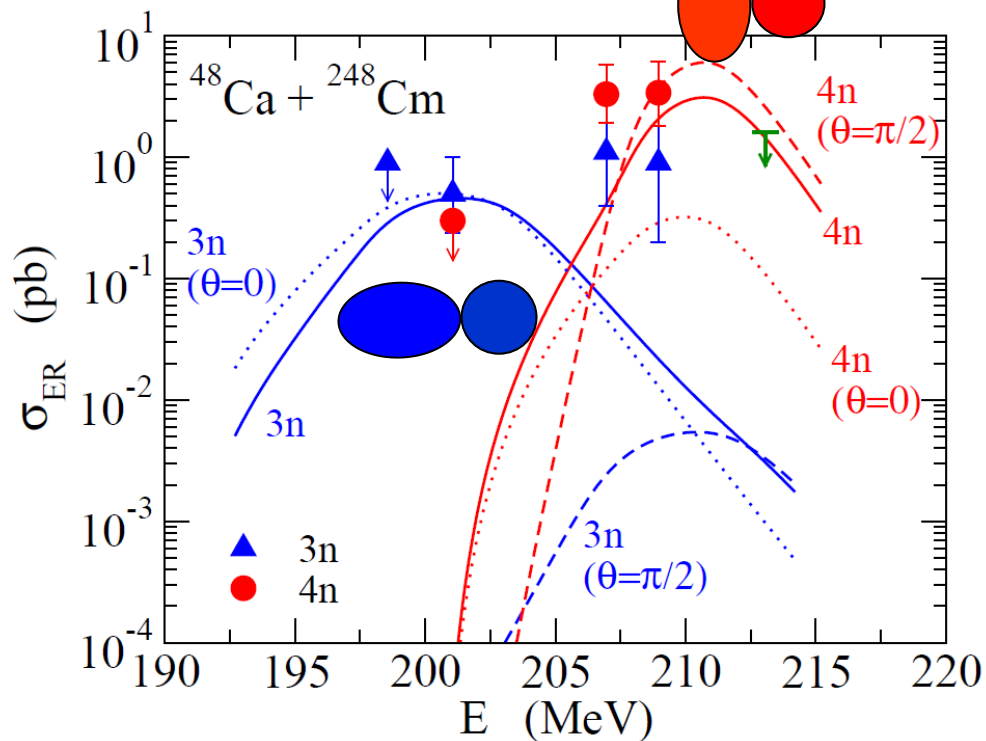
Multi-dimensional space

- internuclear separation
- deformation
- mass asymmetry of the two fragments



Analysis with an extended fusion-by-diffusion approach

K.H., PRC98 ('18) 014607



New hybrid model: TDHF + Langevin approach

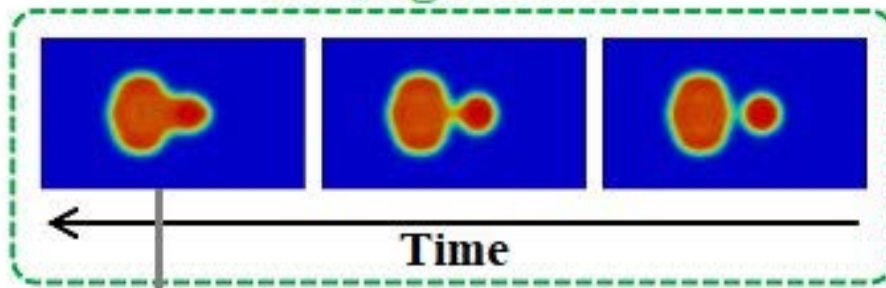
K. Sekizawa and K.H., PRC99 (2019) 051602(R)



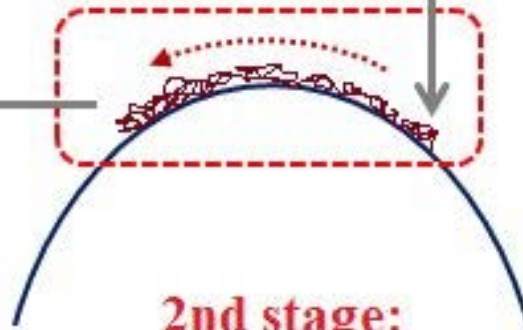
TDHF+Langevin:

a new hybrid model of fusion reactions for superheavy elements

1st stage: TDHF



3rd stage:
statistical model

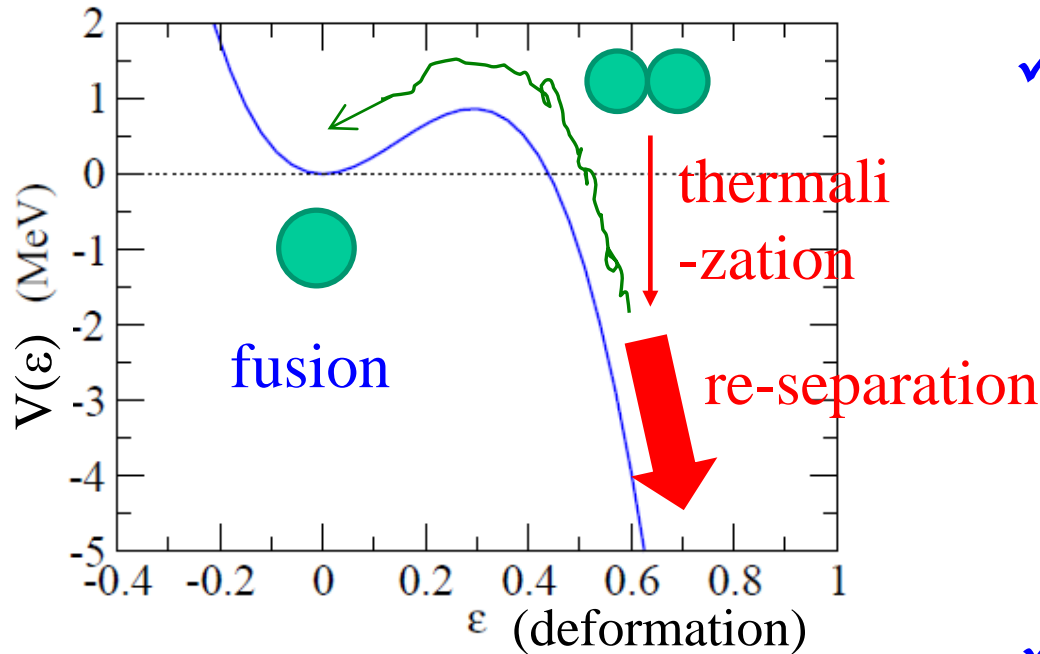
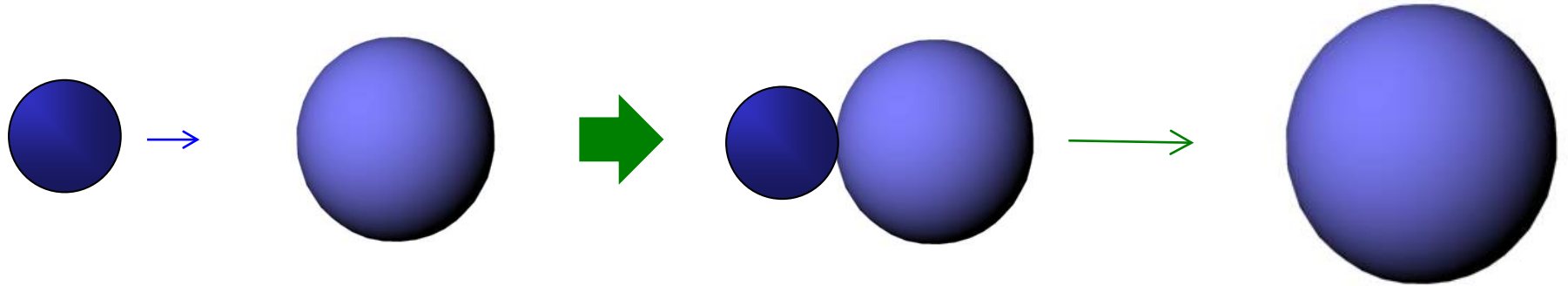


2nd stage:
Langevin model

System	CN (fm)	R_{\min}	P_{fus} ($\times 10^{13}$)
$^{48}\text{Ca} + ^{254}\text{Fm}$	302_{120}	12.93	302
$^{54}\text{Cr} + ^{248}\text{Cm}$	302_{120}	13.09	2.47
$^{51}\text{V} + ^{249}\text{Bk}$	300_{120}	12.94	0.461
$^{48}\text{Ca} + ^{257}\text{Fm}$	305_{120}	12.94	1.82

a special role of ^{48}Ca ?

issues from the theoretical physics point of view



✓ how to thermalize?

Quantum friction theory

c.f. tunneling with quantum friction

M. Tokieda and K.H.,
PRC95 ('17) 054604

✓ non-Markov effect?

✓ quantum correction for diffusion over the barrier?

thermal diffusion

→ Langevin approach

Quantum friction

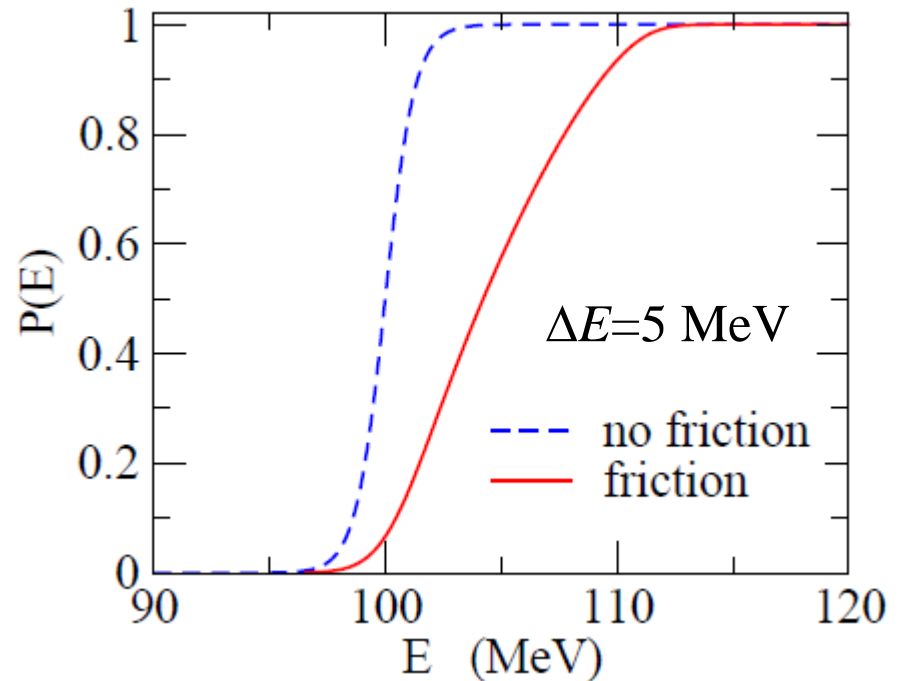
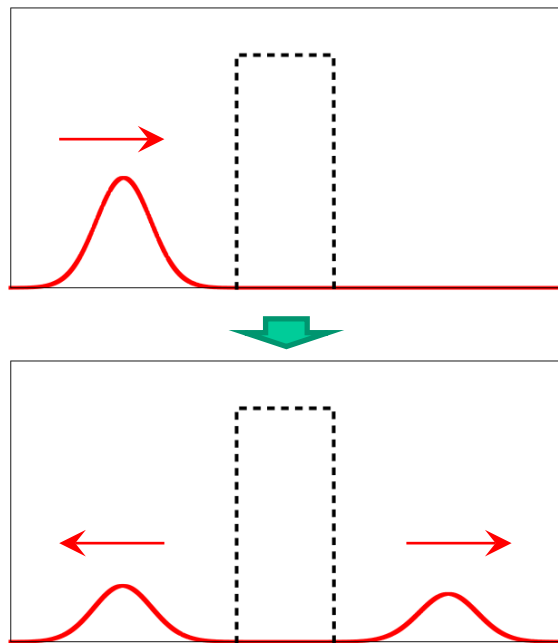
classical eq. of motion $\dot{p} = -V'(x) - \gamma p$

a quantization: Kanai model E. Kanai, PTP 3 (1948) 440

$$H = \frac{p^2}{2m} + V(x) \rightarrow \frac{\pi^2}{2m} e^{-\gamma t} + e^{\gamma t} V(x) \quad (\pi = e^{\gamma t} p)$$

$\longrightarrow \frac{d}{dt} \langle p \rangle = -\langle V'(x) \rangle - \gamma \langle p \rangle$

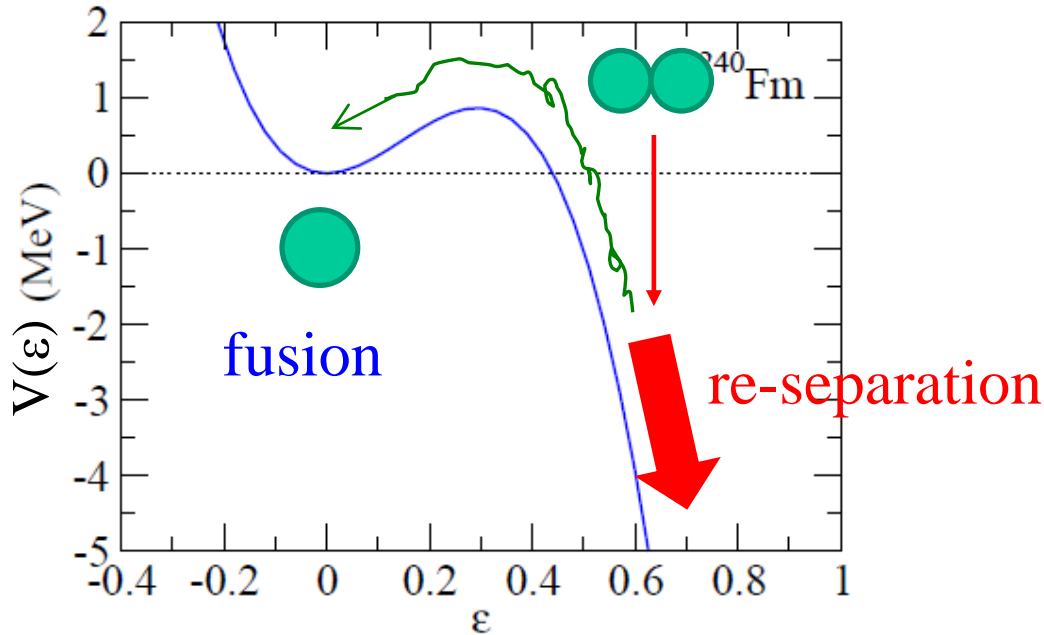
time-dep. wave packet approach



Fusion reactions and non-equilibrium statistical mechanics

:Langevin dynamics under a temperature gradient

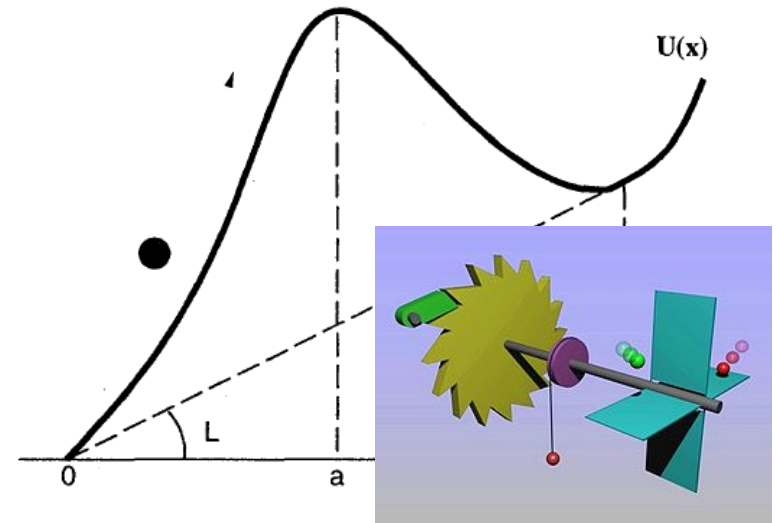
➤ Superheavy elements



$$E_{\text{int}} = E^* - E_{\text{kin}} - V(\epsilon) = aT^2$$

← coordinate dependent temperature

➤ a math model for molecular motors

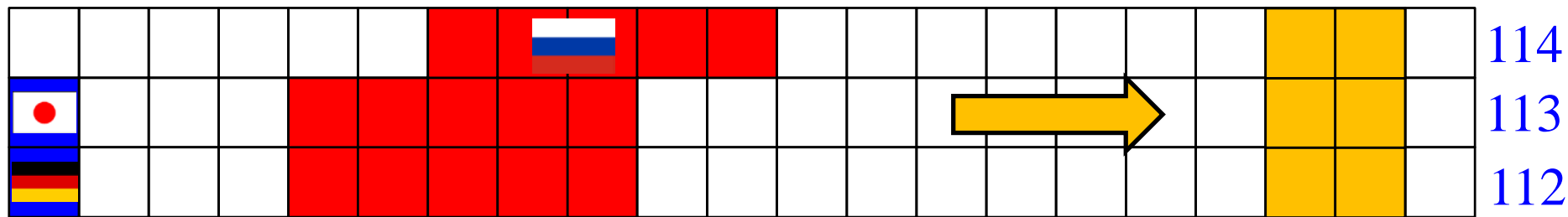


松尾美希、物性研究 73 ('99) 557

temperature gradient
→ one-way dynamics

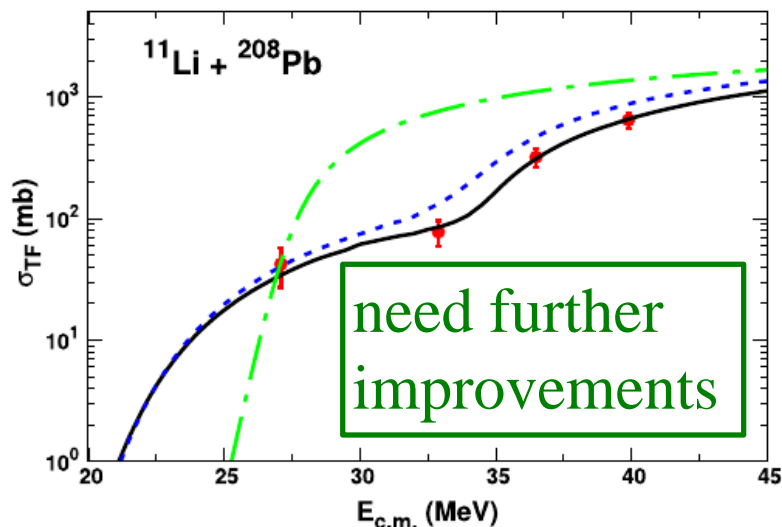
SHE formation reactions as a general problem of non-eq. stat. mechanics?

Fusion of unstable nuclei

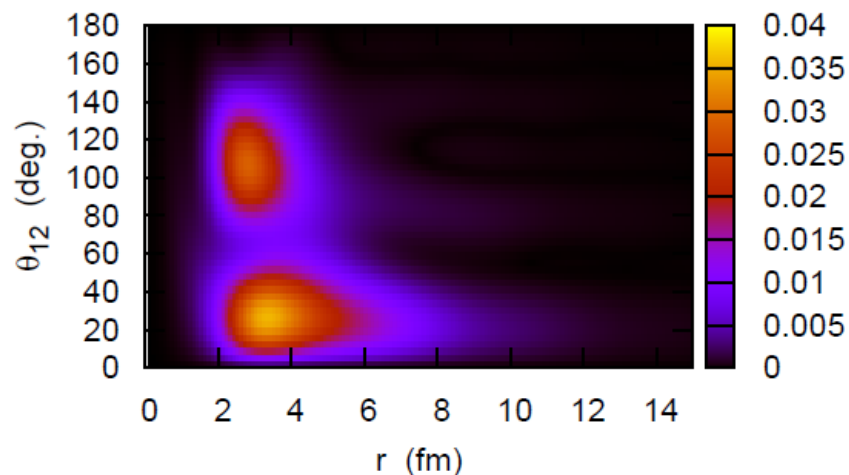


165 170 175 180 the island of stability

neutron-rich beams: indispensable → reaction dynamics?



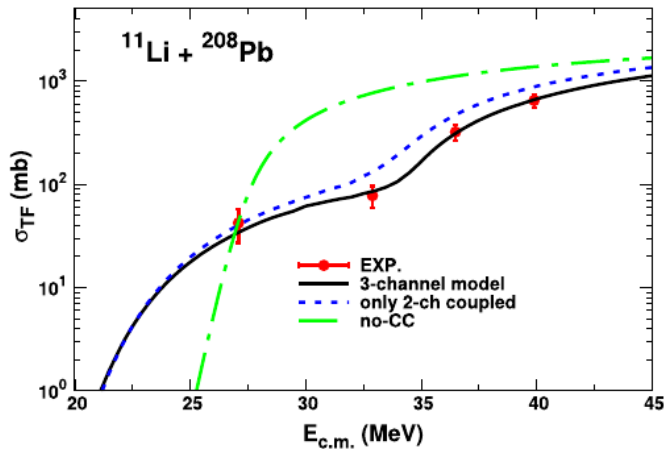
K.-S. Choi, K. Hagino et al.,
Phys. Lett. B780 ('18) 455



K.H. and H. Sagawa, PRC72('05)044321

good understandings of the structure
of neutron-rich nuclei is also important

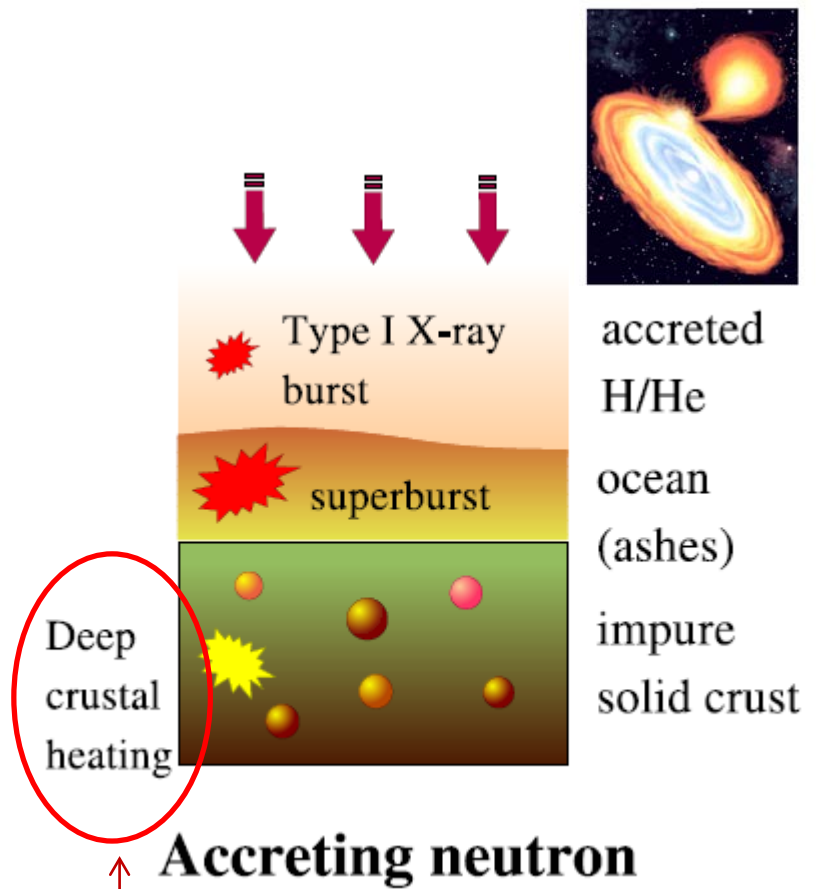
reactions of neutron-rich nuclei



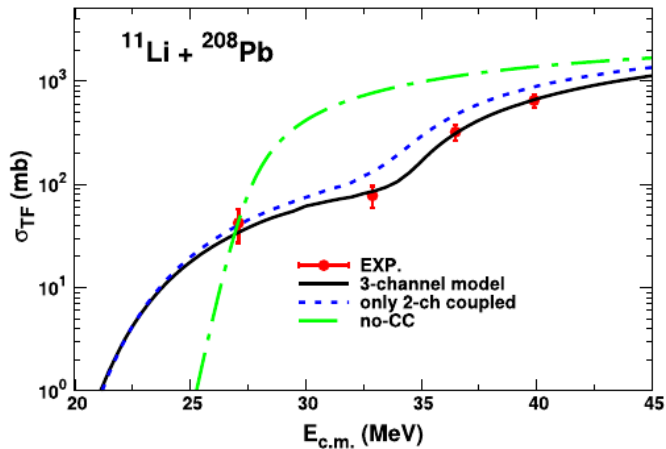
- ✓ fusion
- ✓ transfer



- development of microscopic nuclear reaction theory
- nuclear reactions in neutron stars

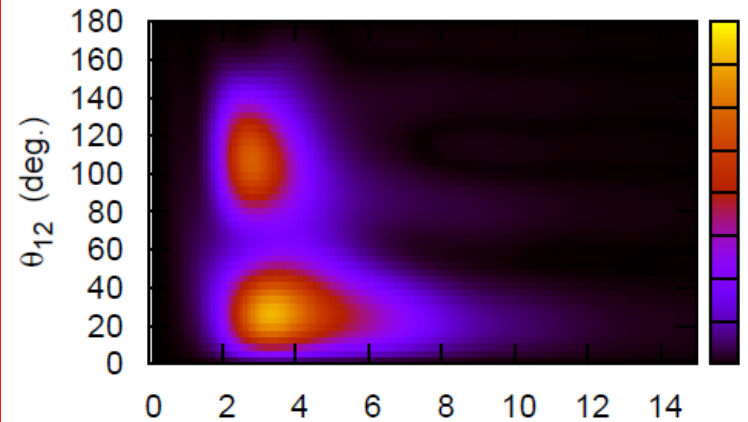


reactions of neutron-rich nuclei



- ✓ fusion
- ✓ transfer

structure of neutron-rich nuclei



- ✓ nucleon correlations
- ✓ collective motions
- ✓ fission

- development of microscopic nuclear reaction theory
- nuclear reactions in neutron stars

from few-body to many-body

Physics of SHE with n-rich nuclei as important ingredient

Summary

SHE: quantum many-body systems with a strong Coulomb field

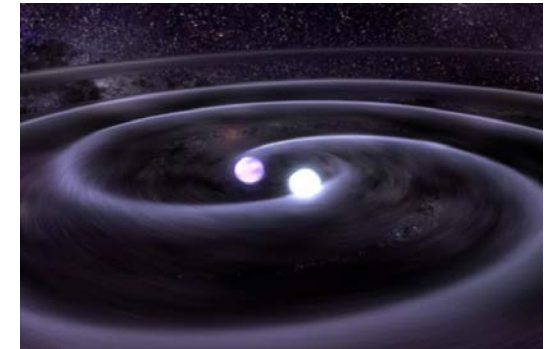
physics

chemistry

astronomy

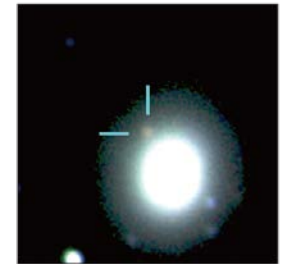
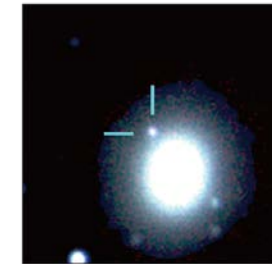
113 Nh nihonium	115 Mc moscovium
117 Ts tennessine	118 Og oganeson

Group	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
1	1 H																	2 He
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
6	55 Cs	56 Ba	57 La*	72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7	87 Fr	88 Ra	89 Ac*	104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
				58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
				90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr	



2017.08.18-19

2017.08.24-25



reaction dynamics

- ✓ quantum friction
- ✓ neutron-rich nuclei



International Year
of the Periodic Table
of Chemical Elements

- ✓ origin of elements
- ✓ r-process
- ✓ kilonova

interdisciplinary SHE science

$$H = -\frac{\hbar^2}{2\mu} \nabla^2 + V_0(r) + \beta f(r) \hat{O}$$

full order treatment

$$V(r, \beta \hat{O})$$

$$\hat{O}|\phi_k\rangle = \lambda_k|\phi_k\rangle \leftarrow \text{diagonalize } \langle n|\hat{O}|m\rangle$$

$$\rightarrow \langle n|V(r, \beta \hat{O})|m\rangle = \sum_k \langle n|\phi_k\rangle \langle \phi_k|m\rangle V(r, \beta \lambda_k)$$

K.H., N. Rowley, and A.T. Kruppa,
Comp. Phys. Comm. 123('99) 143.