



**NEW FRONTIERS IN NUCLEAR PHYSICS
AND NUCLEAR ASTROPHYSICS**

28 May - 1 June 2018 Antalya, TURKEY



NEW FRONTIERS IN NUCLEAR PHYSICS AND ASTROPHYSICS-1

28 May – 1 June 2018

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Asi
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Mes

Sait Umar
Andrea Vitturi



Heavy-ion fusion reactions for superheavy elements

Kouichi Hagino

Tohoku University, Sendai, Japan

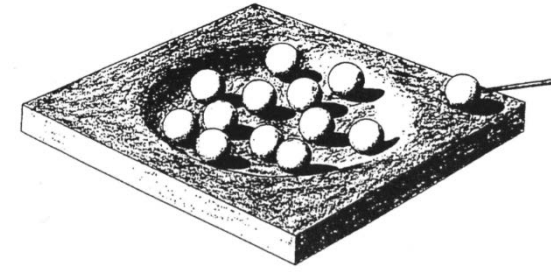
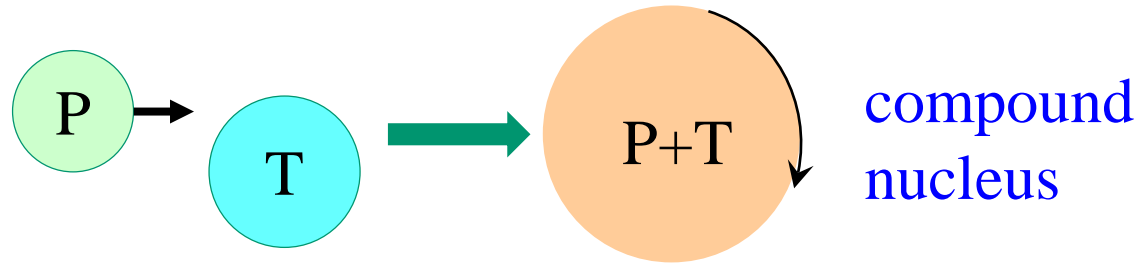


1. H.I. sub-barrier fusion reactions
2. Coupled-channels approach and barrier distributions
3. Application to superheavy elements
4. Hot fusion reactions of a deformed target
5. Summary and discussions

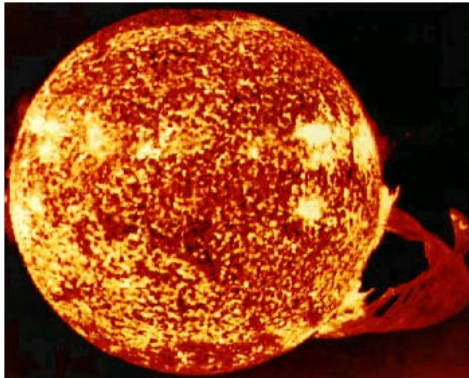
Recent review article:

K. Hagino and N. Takigawa, Prog. Theo. Phys.128 ('12)1061.

Fusion reactions: compound nucleus formation

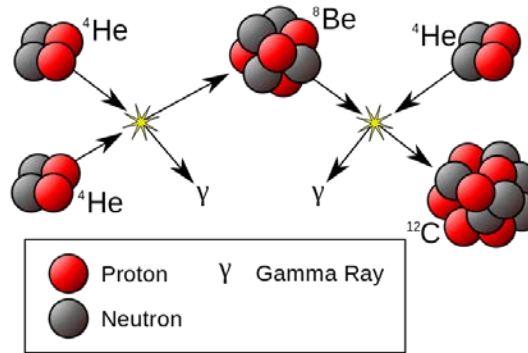


cf. Bohr '36

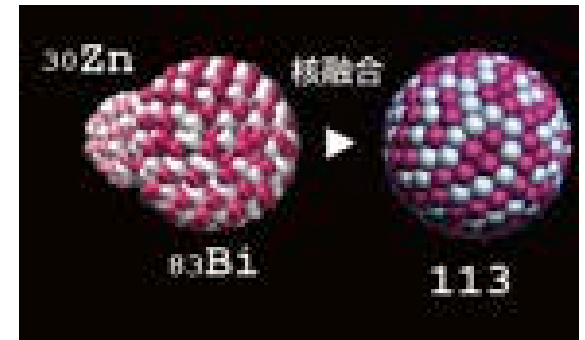


NASA, Skylab space station on December 19, 1973, solar flare reaching 589 000 km off solar surface

energy production
in stars



nucleosynthesis

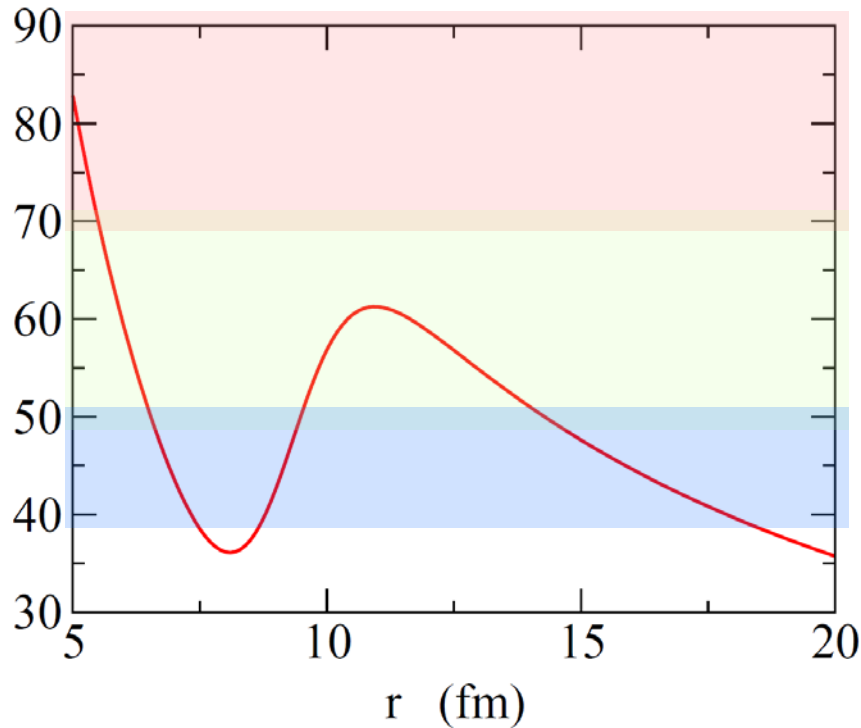


superheavy elements

Fusion and fission: large amplitude motions of quantum many-body systems with strong interaction

← microscopic understanding: an ultimate goal of nuclear physics

Fusion reactions: compound nucleus formation



fusion reactions
in the sub-barrier energy region

$$(|E - V_b| \lesssim 10\text{MeV})$$

- 1. Coulomb force : long range, repulsive
- 2. Nuclear force : short range, attractive



Coulomb barrier

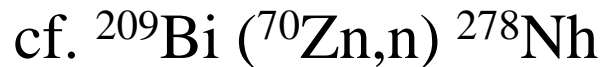
Why sub-barrier fusion?

two obvious reasons:

113 Nh nihonium	115 Mc moscovium
117 Ts tennessine	118 Og oganesson

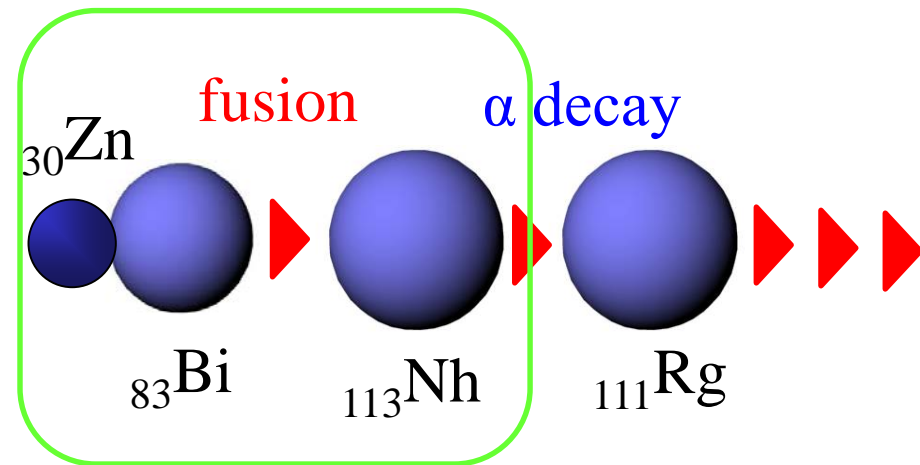
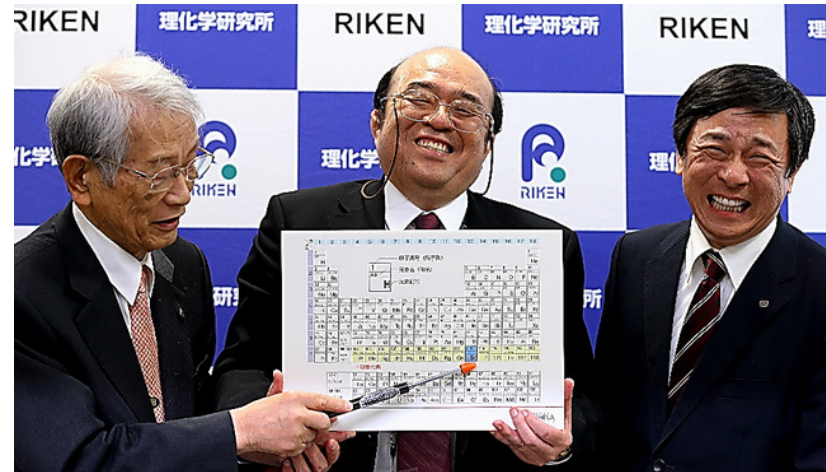
November, 2016

superheavy elements



$$V_B \sim 260 \text{ MeV}$$

$$E_{\text{cm}}^{(\text{exp})} \sim 262 \text{ MeV}$$



Why sub-barrier fusion?

two obvious reasons:

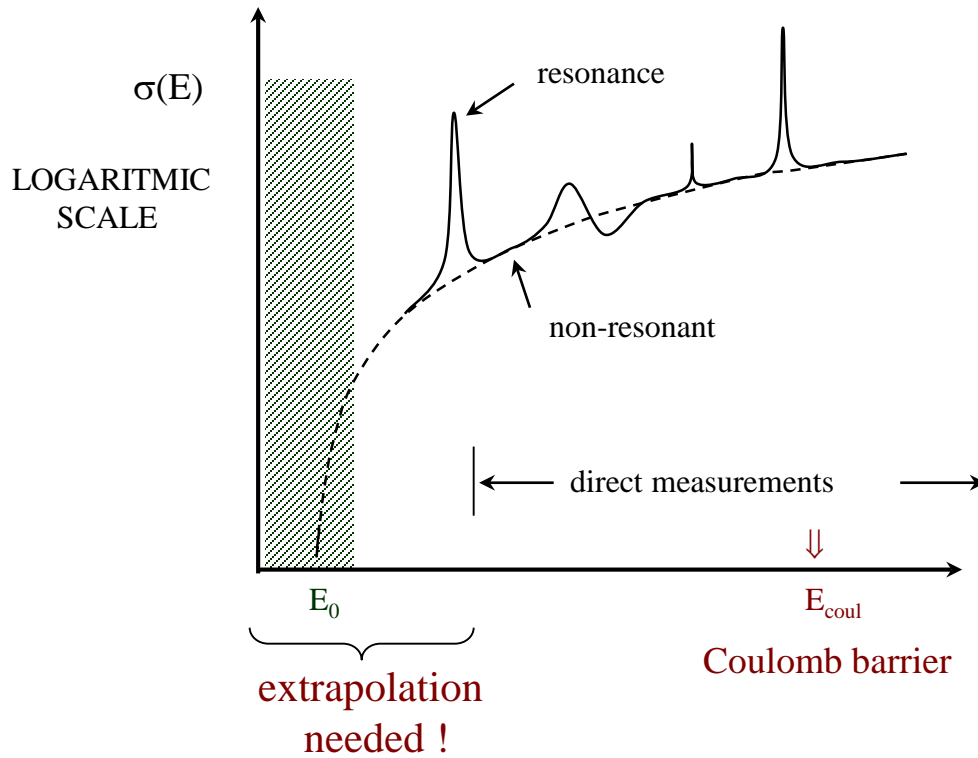
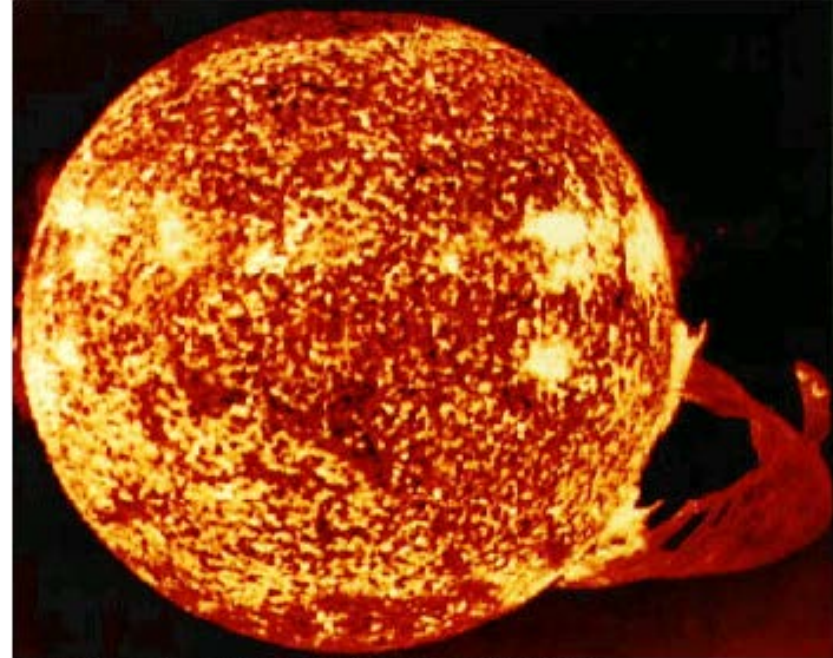


figure: M. Aliotta



NASA, Skylab space station December 19, 1973, solar flare reaching 589 000 km off solar surface

nuclear astrophysics
(nuclear fusion in stars)

cf. extrapolation of data

Why sub-barrier fusion?

Two obvious reasons:

- ✓ discovering new elements (SHE)
- ✓ nuclear astrophysics (fusion in stars)

113 Nh nihonium	115 Mc moscovium
117 Ts tennessine	118 Og oganesson



Other reasons:

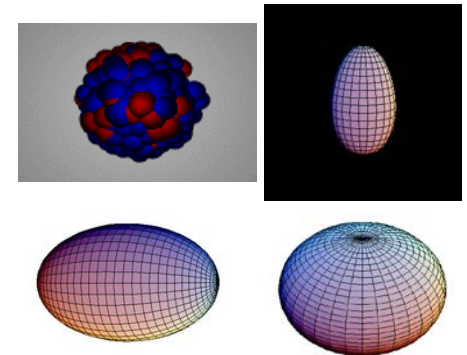
◆ reaction mechanism

strong interplay between reaction and nuclear structure

cf. high E reactions: much simpler reaction mechanism

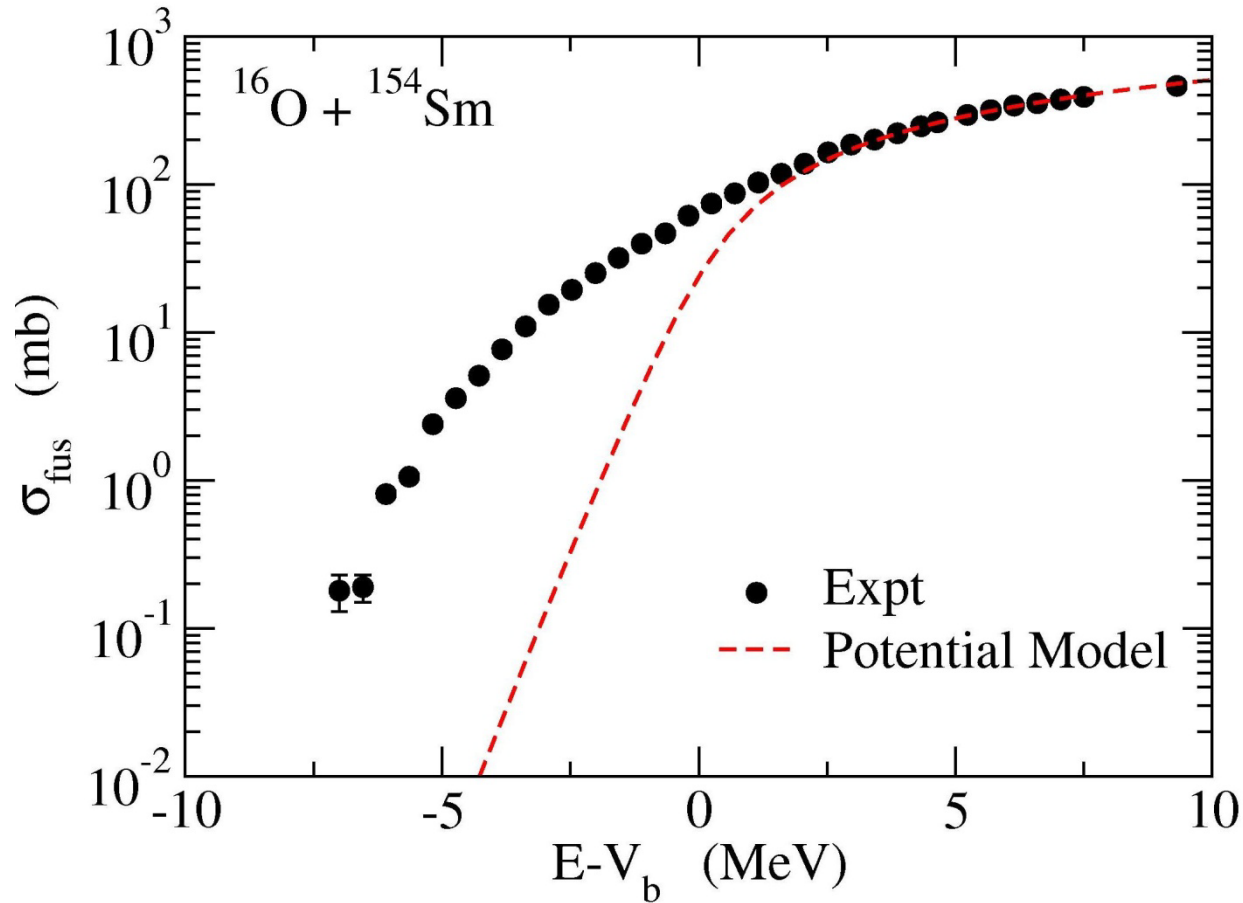
◆ many-particle tunneling

- ✓ many types of intrinsic degrees of freedom
- ✓ energy dependence of tunneling probability
cf. alpha decay: fixed energy



H.I. fusion reaction = an ideal playground to study quantum tunneling with many degrees of freedom

Discovery of large sub-barrier enhancement of σ_{fus}



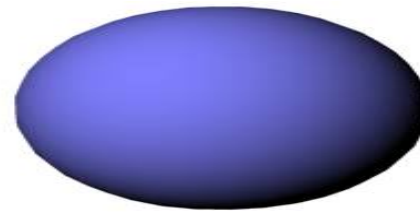
potential model: $V(r) + \text{absorption}$

cf. seminal work:

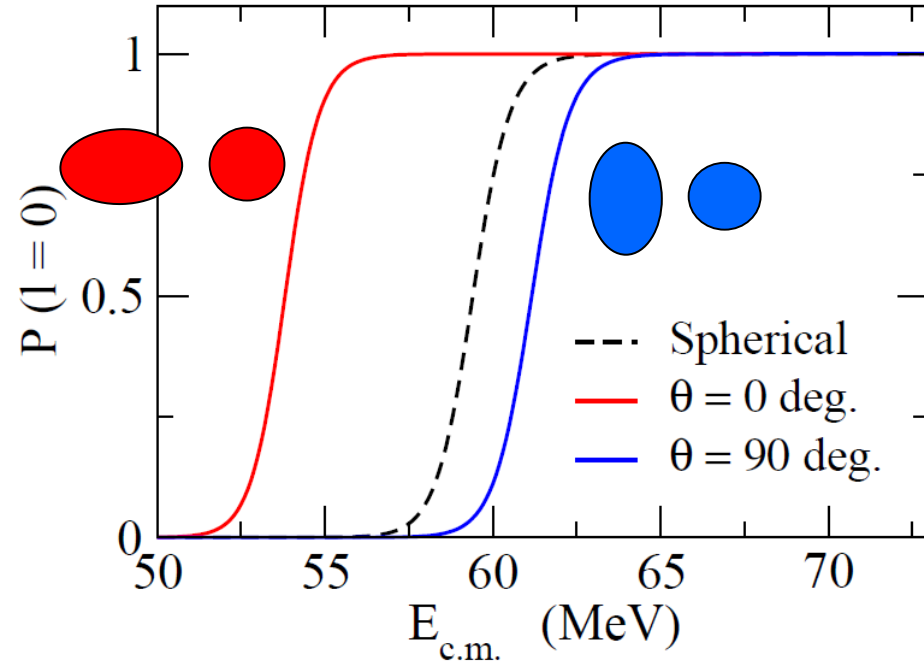
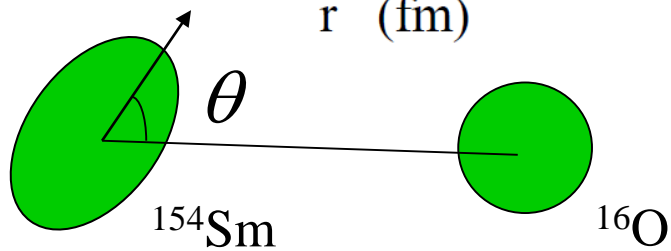
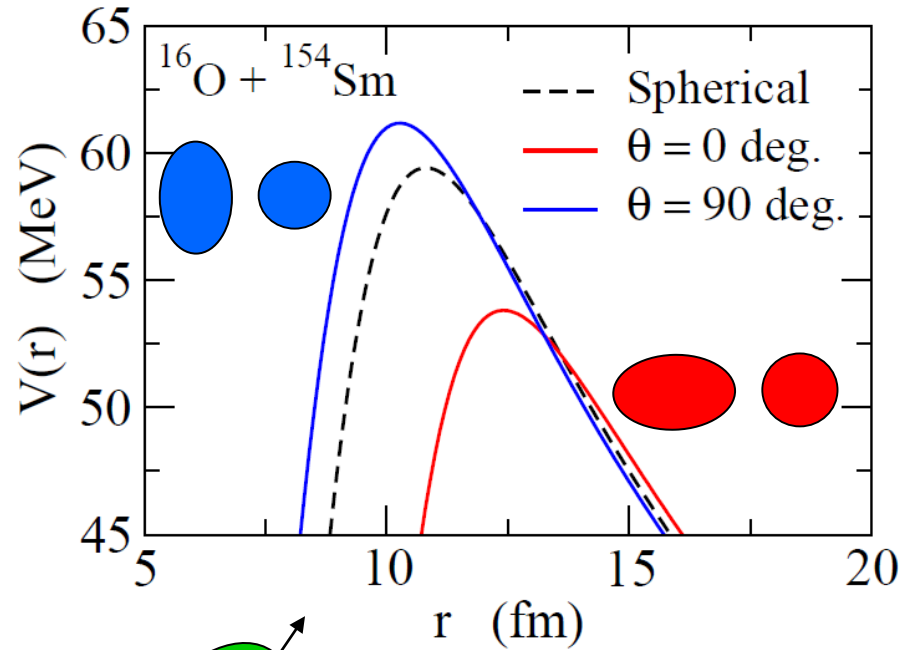
R.G. Stokstad et al., PRL41('78) 465

Effects of nuclear deformation

^{154}Sm : a typical deformed nucleus

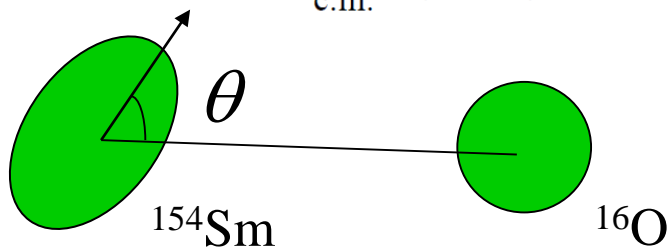
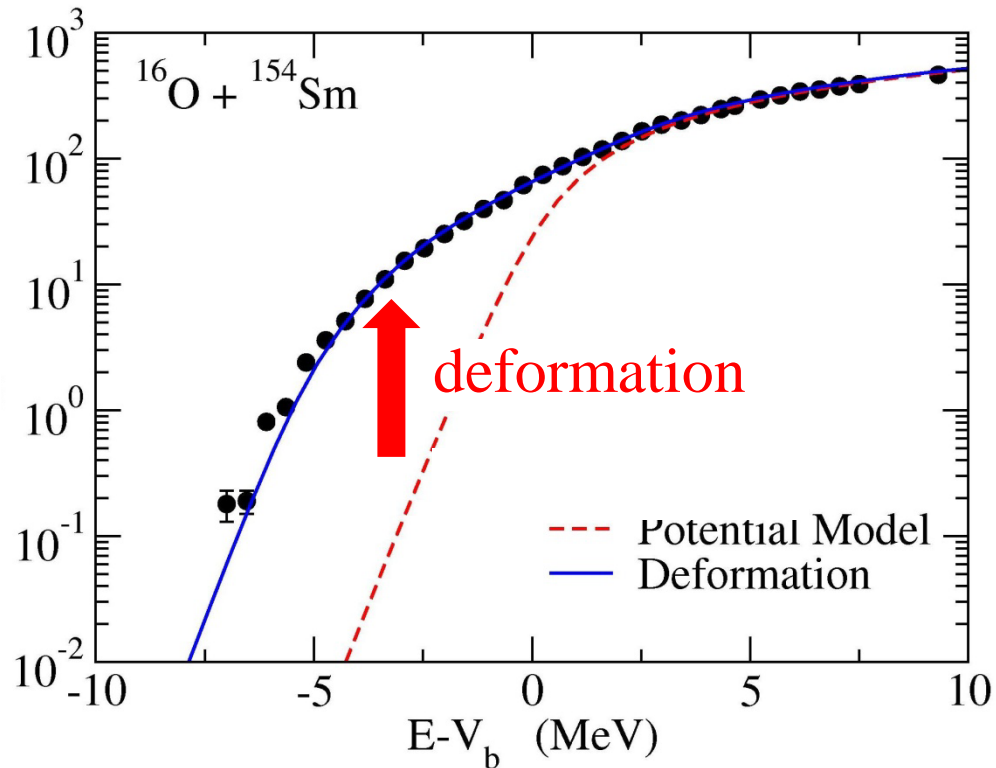
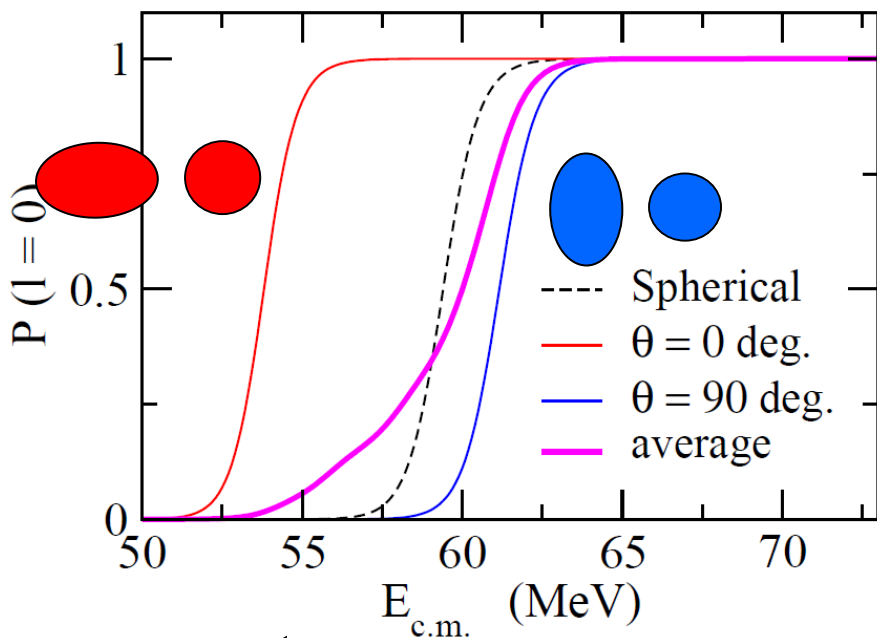
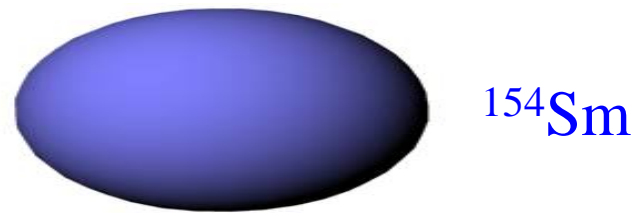


^{154}Sm



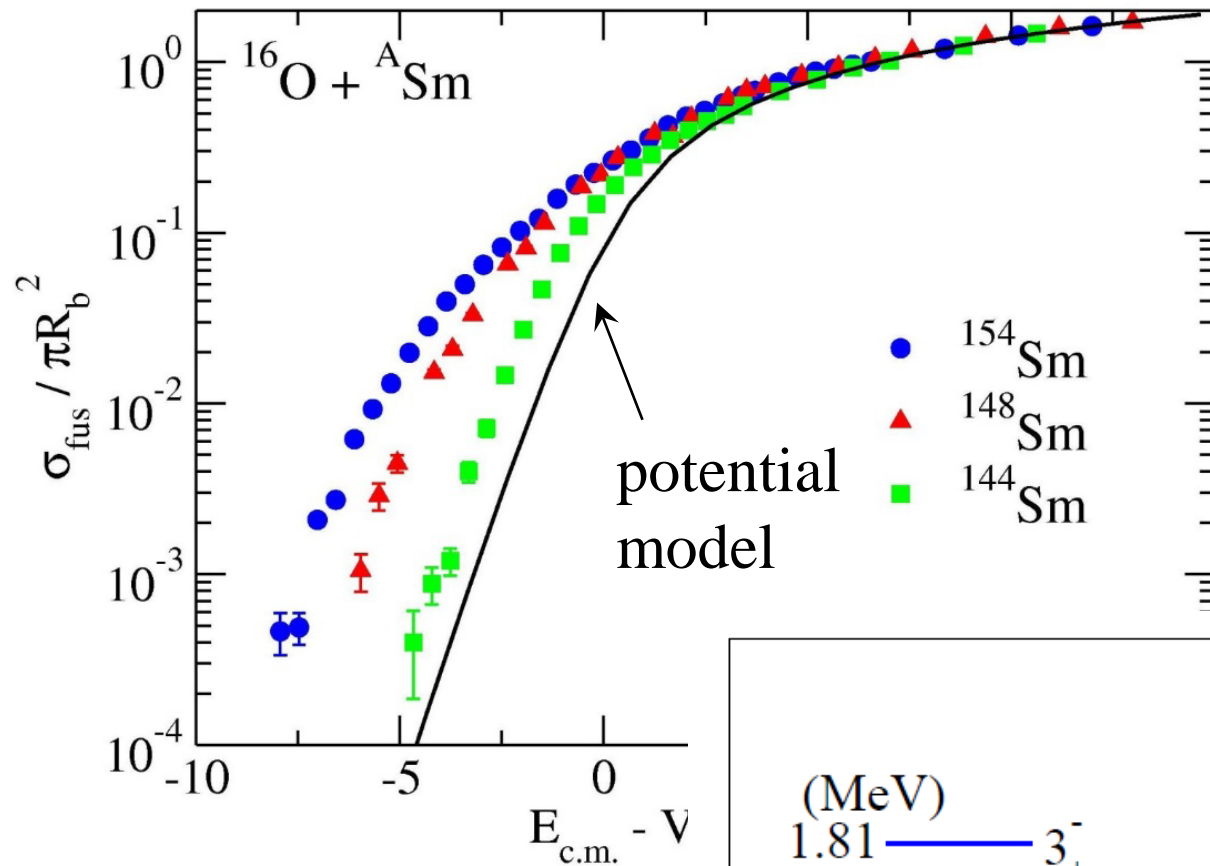
Effects of nuclear deformation

^{154}Sm : a typical deformed nucleus



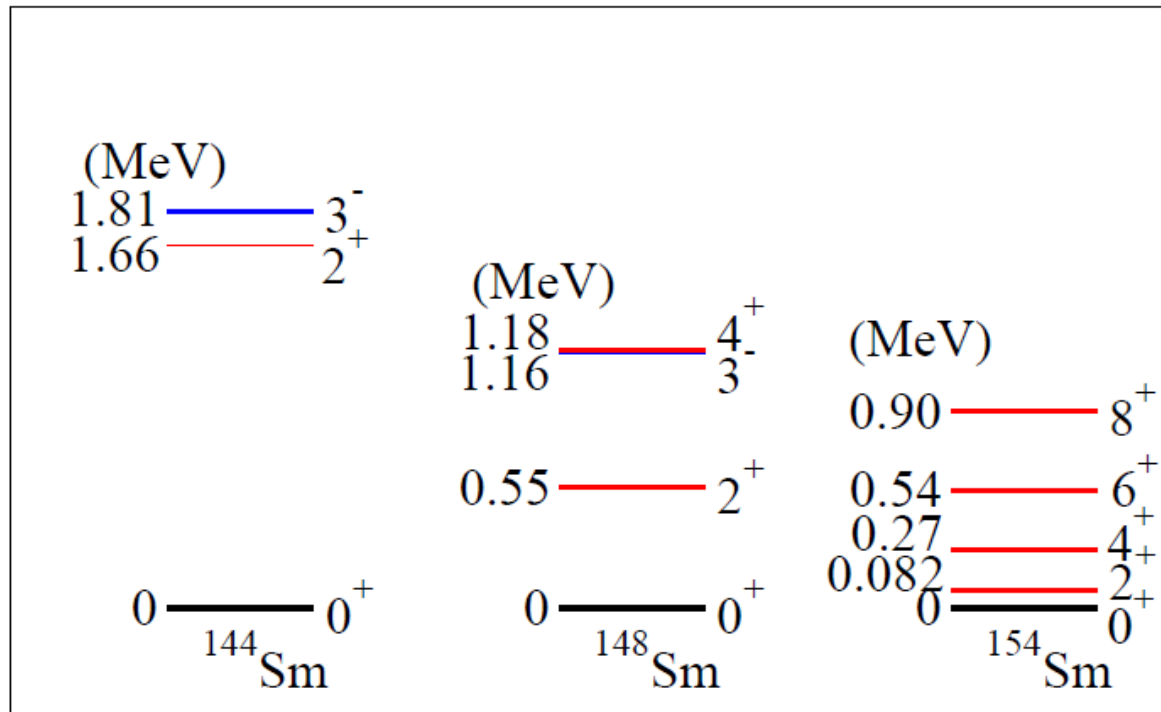
$$\sigma_{\text{fus}}(E) = \int_0^1 d(\cos \theta) \sigma_{\text{fus}}(E; \theta)$$

Fusion: strong interplay between nuclear structure and reaction



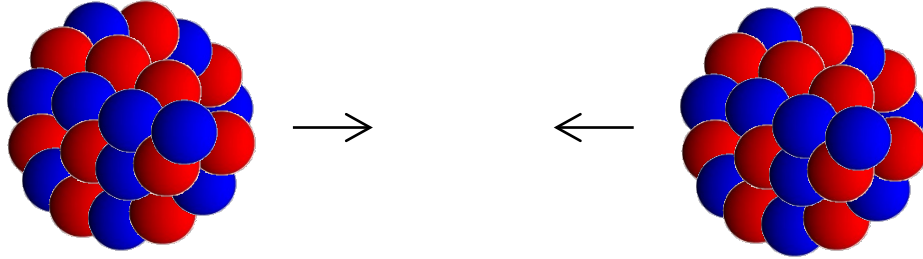
enhancement of fusion cross sections
: a general phenomenon

strong correlation with nuclear spectrum
→ coupling assisted tunneling



Coupled-channels method: a quantal scattering theory with excitations

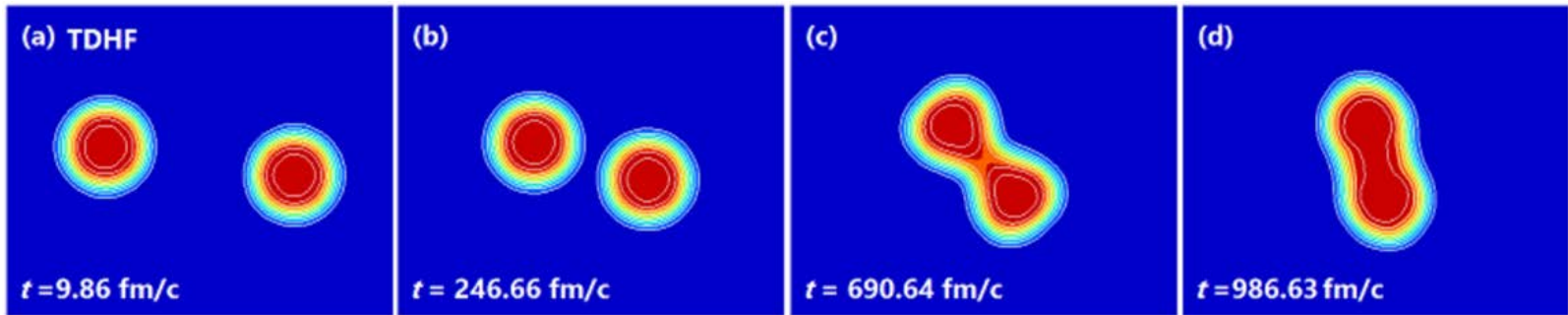
many-body problem



still very challenging

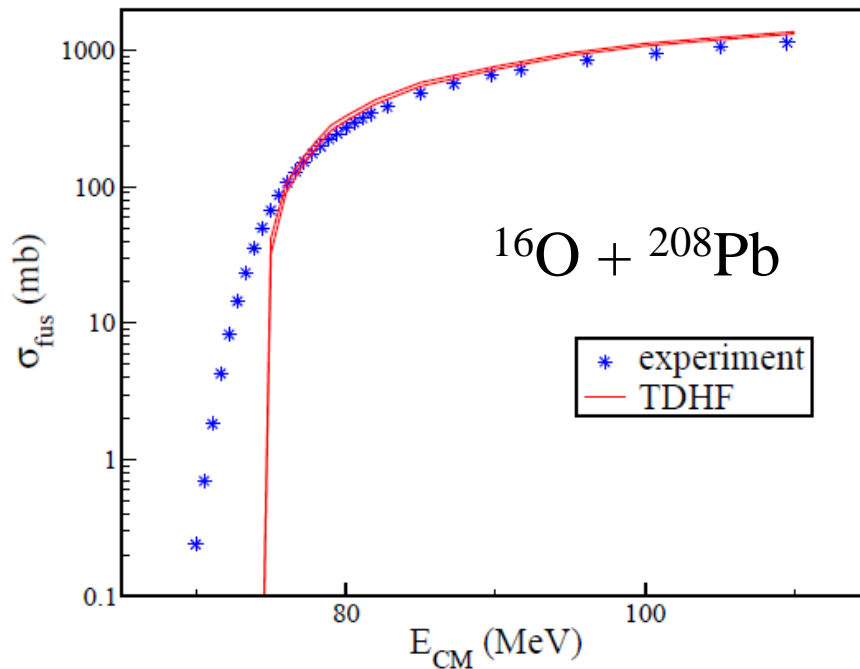
TDHF simulation

TDHF = Time Dependent Hartree-Fock



S. Ebata, T. Nakatsukasa, JPC Conf. Proc. 6 ('15) 020056

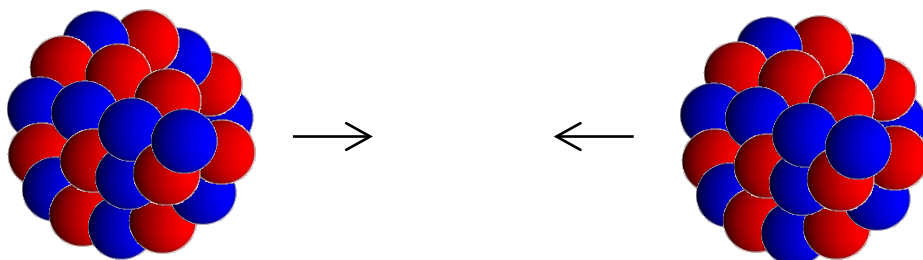
“ab-initio”, but no tunneling



C. Simenel,
EPJA48 ('12) 152

Coupled-channels method: a quantal scattering theory with excitations

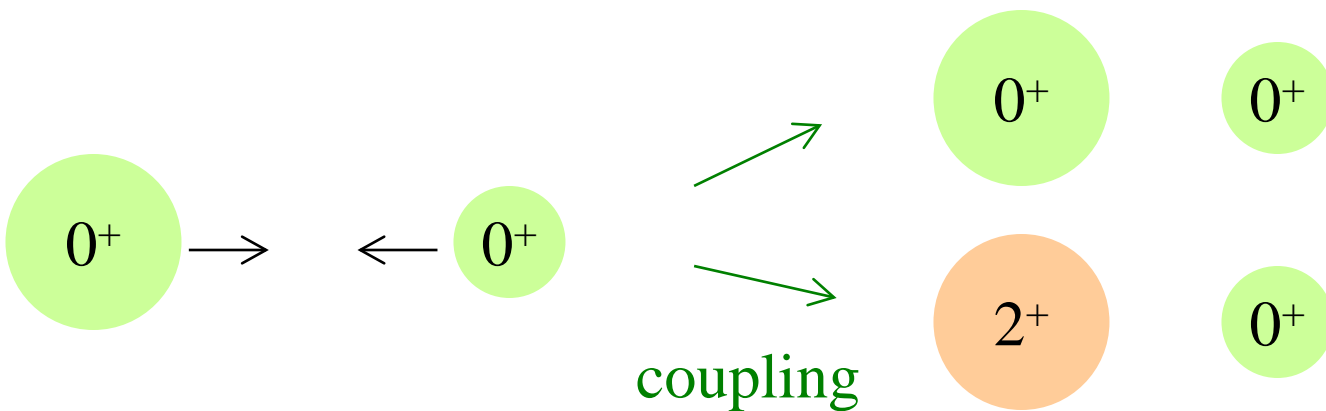
many-body problem



still very challenging



two-body problem, but with excitations
(coupled-channels approach)



scattering theory with excitations

C.C. approach: a standard tool for sub-barrier fusion reactions

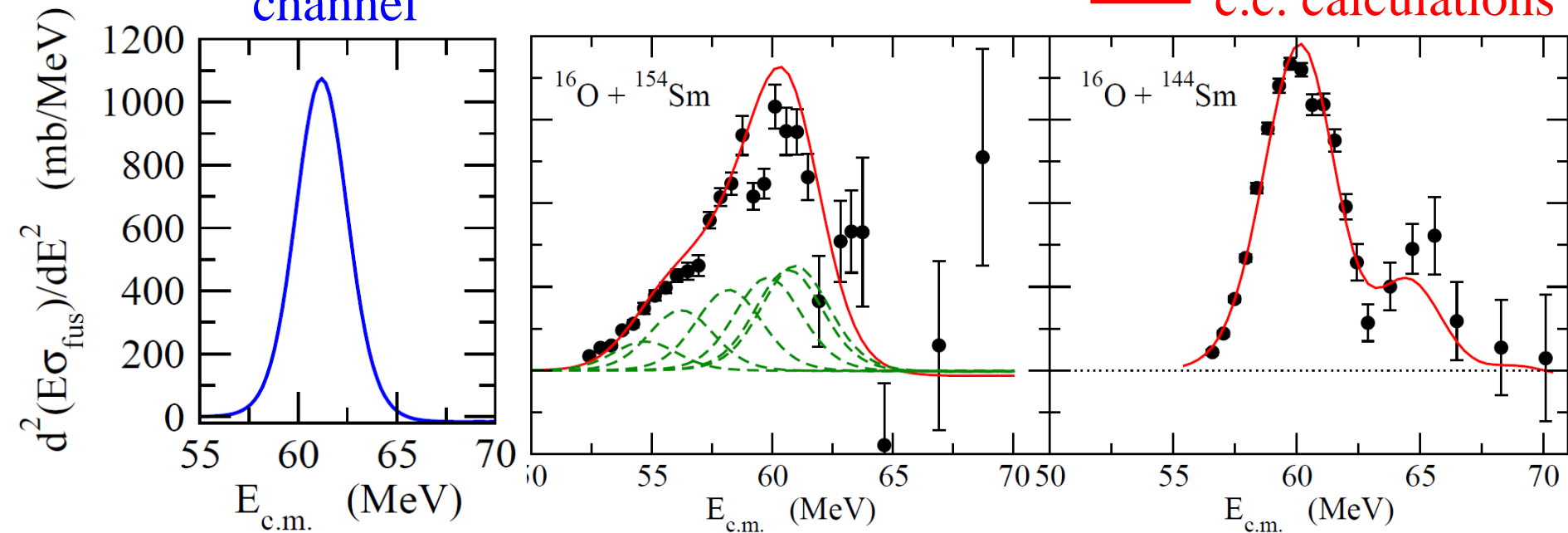
cf. CCFULL (K.H., N. Rowley, A.T. Kruppa, CPC123 ('99) 143)

- ✓ Fusion barrier distribution [Rowley, Satchler, Stelson, PLB254('91)]

$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2}$$

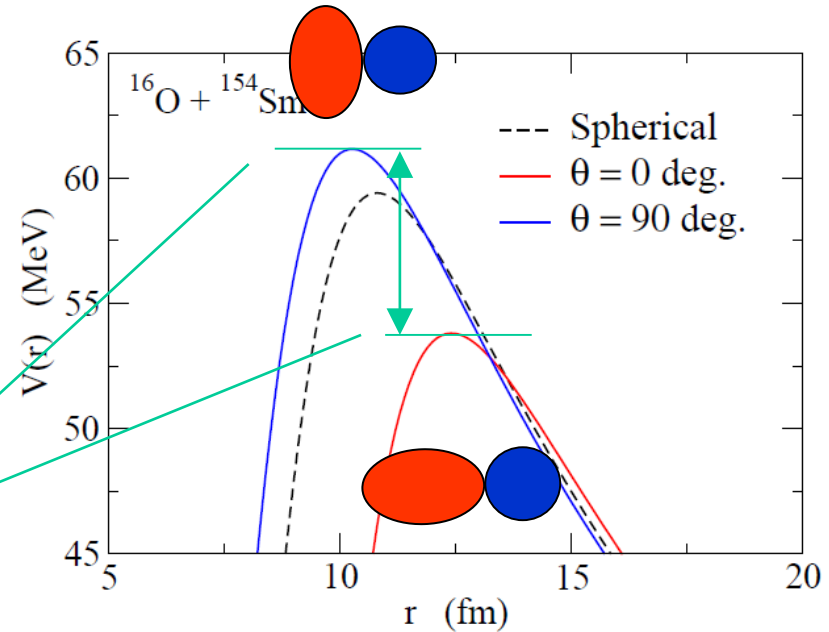
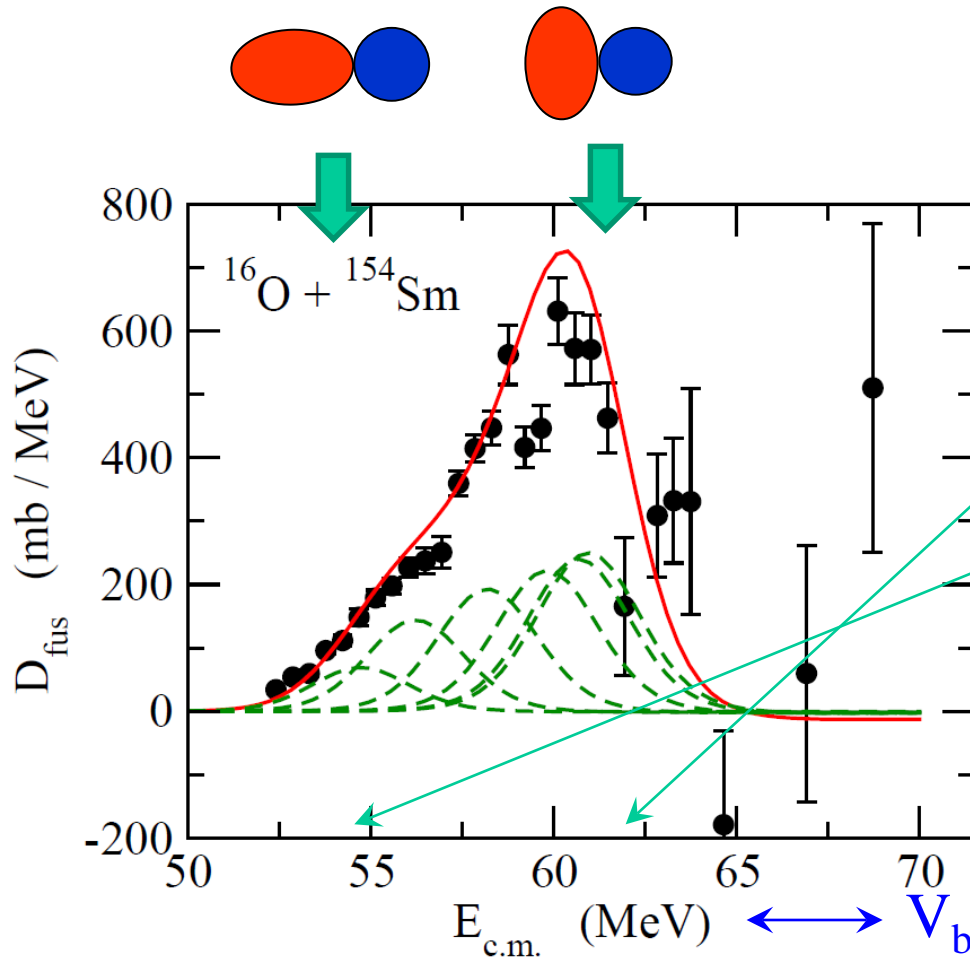
single
channel

— c.c. calculations



✓ Fusion barrier distribution (Rowley, Satchler, Stelson, PLB254('91))

$$D_{\text{fus}}(E) = \frac{d^2(E\sigma_{\text{fus}})}{dE^2}$$



Data: J.R. Leigh et al.,
PRC52 ('95) 3151

a nice tool to understand the reaction dynamics

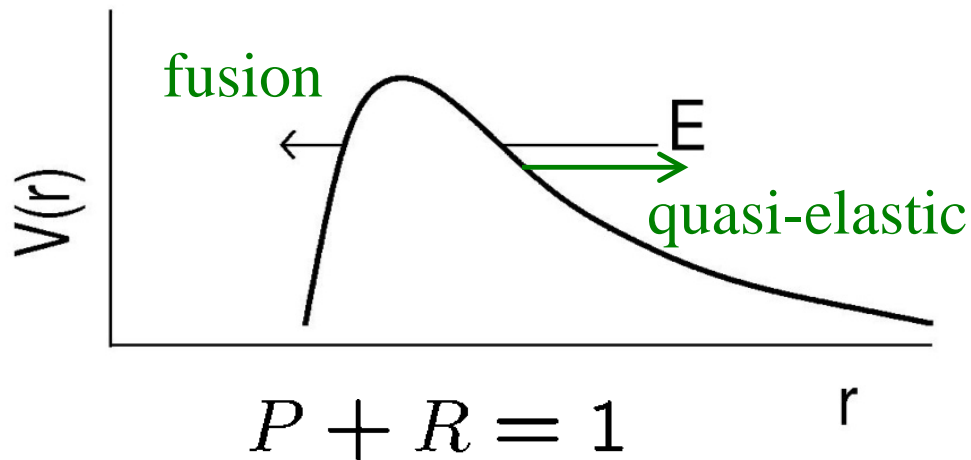
Recent application to SHE : Quasi-elastic B.D.

hot fusion reactions



= deformation \rightarrow

reaction dynamics with
barrier distributions?



Quasi-elastic scattering
: reflected flux at the barrier

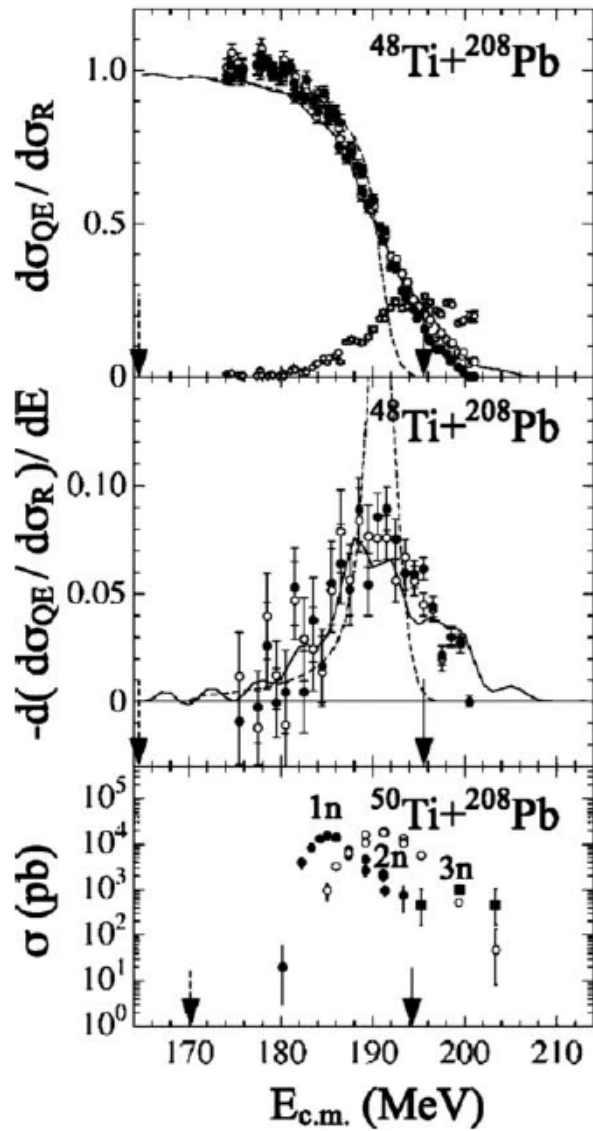
- a sum of elastic, inelastic, and transfer
- easier to measure than capture

Quasi-elastic barrier distribution

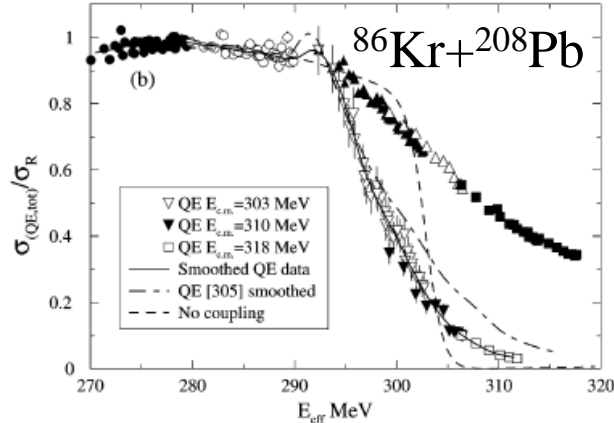
$$D_{\text{qel}}(E) = -\frac{d}{dE} \left(\frac{\sigma_{\text{qel}}(E, \pi)}{\sigma_R(E, \pi)} \right)$$

H. Timmers et al., NPA584('95)190
K.H. and N. Rowley, PRC69('04)054610

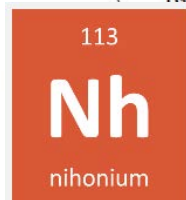
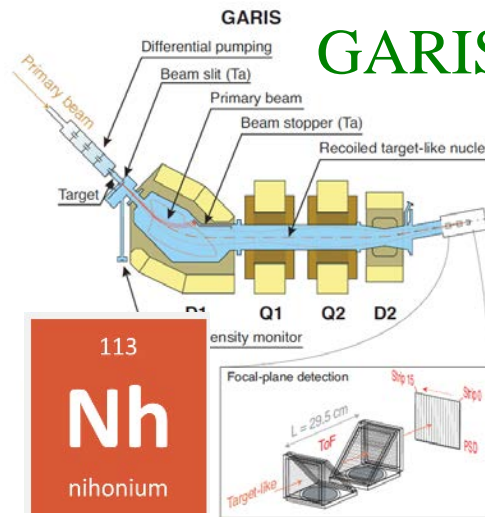
previous attempts



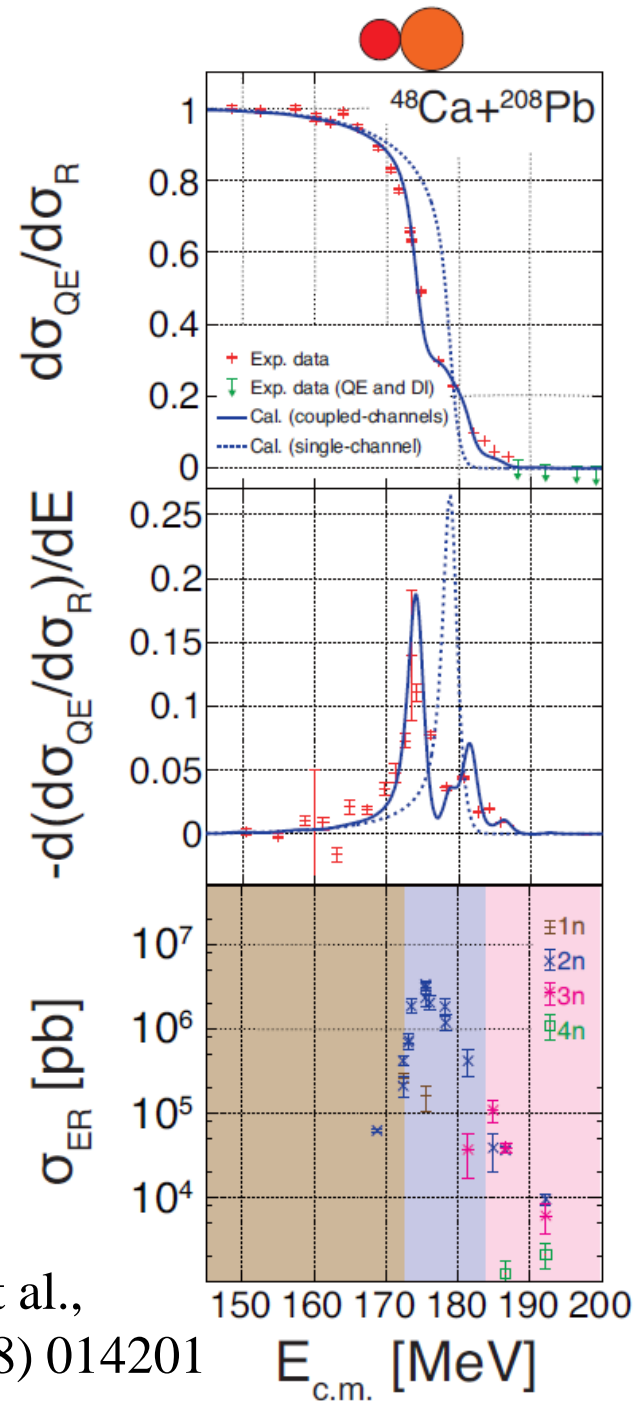
S. Mitsuoka et al.,
PRL99 ('07) 182701



S.S. Ntshangase et al.,
PLB651 ('07) 27



T. Tanaka et al.,
JPSJ 87 ('18) 014201

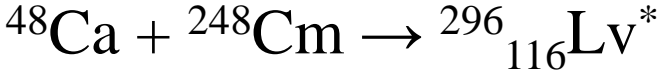


T. Tanaka et al.,
JPSJ 87 ('18) 014201

Analysis for a hot fusion reaction $^{48}\text{Ca} + ^{248}\text{Cm}$

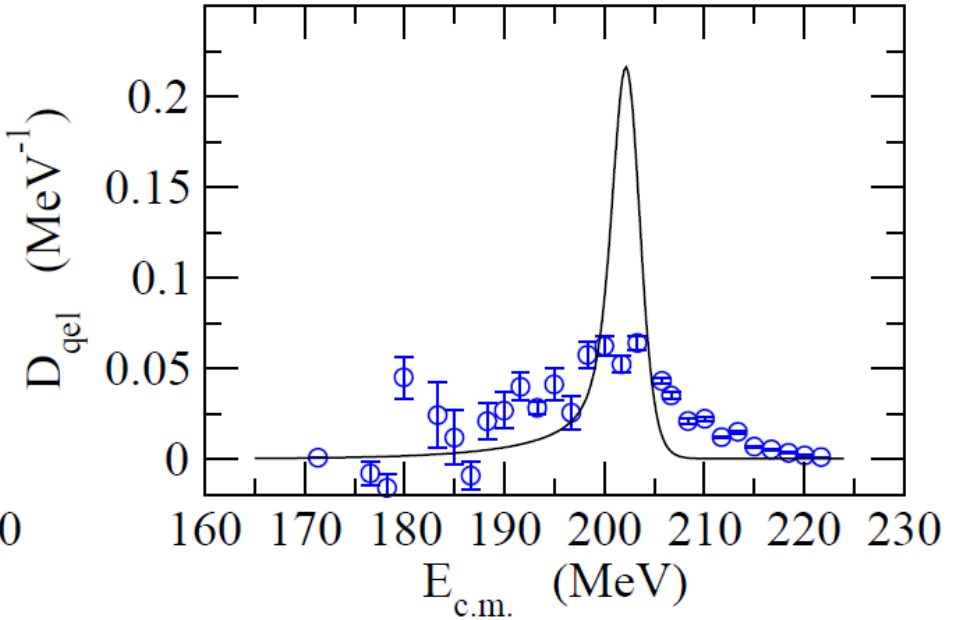
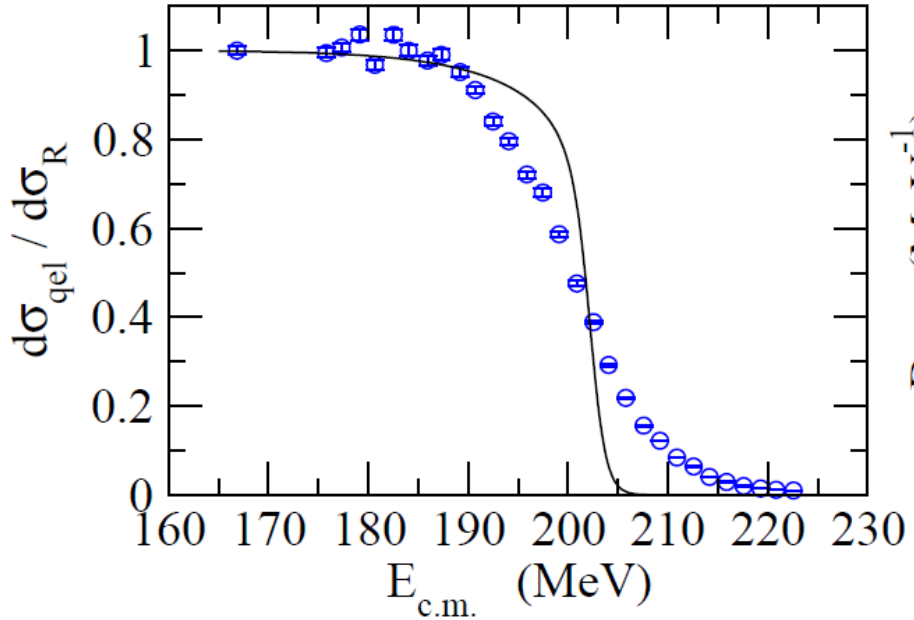
K.H. and T. Tanaka (2017)

(T. Tanaka et al., JPSJ 87 ('18) 014201)



single-channel calculation
(spherical ^{248}Cm)

$$D_{\text{qel}}(E) = -\frac{d}{dE} \left(\frac{\sigma_{\text{qel}}(E, \pi)}{\sigma_R(E, \pi)} \right)$$



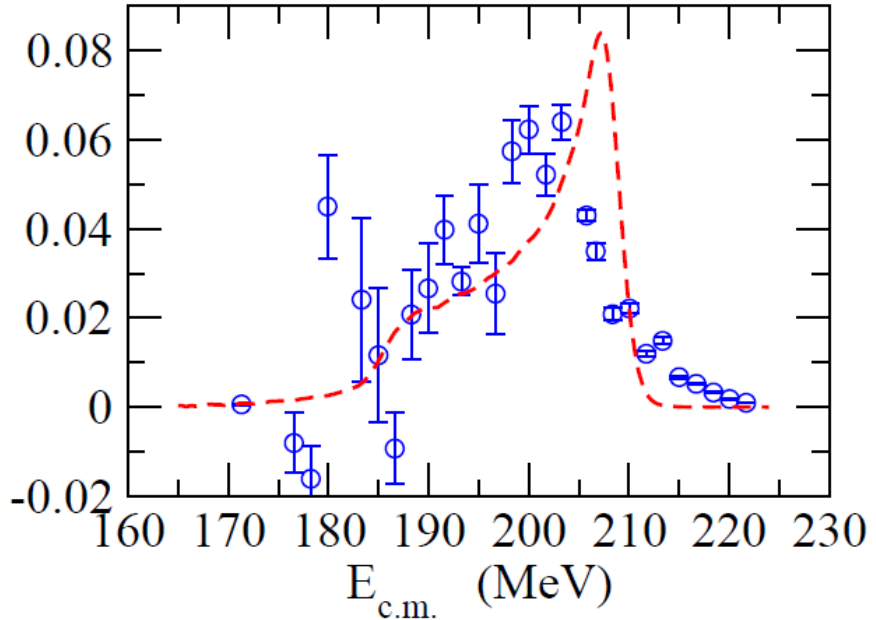
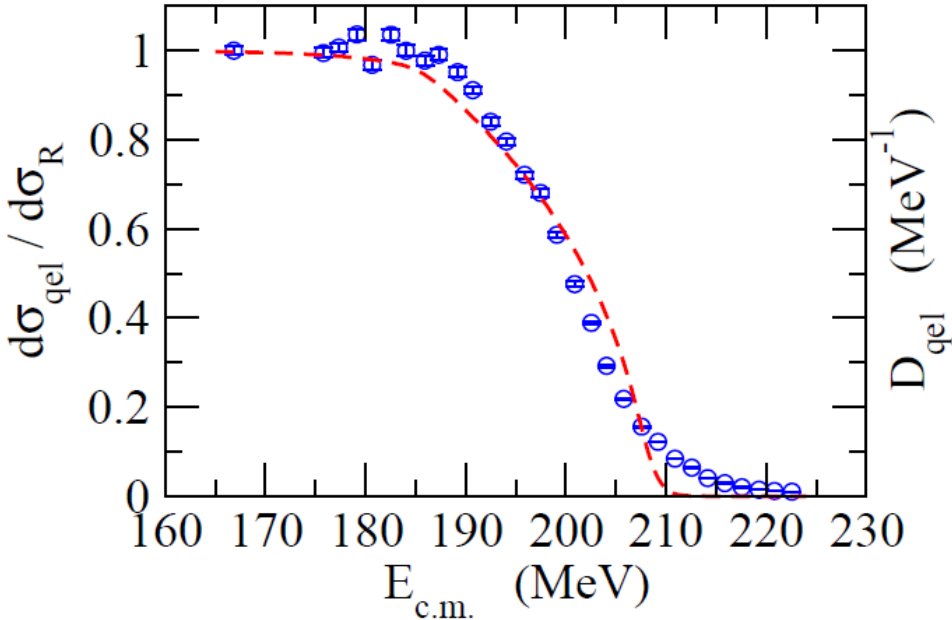
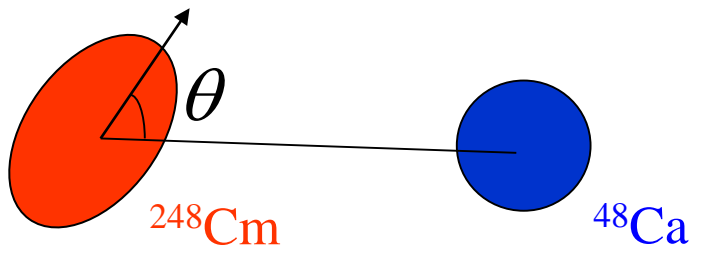
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K.H. and T. Tanaka (2017)



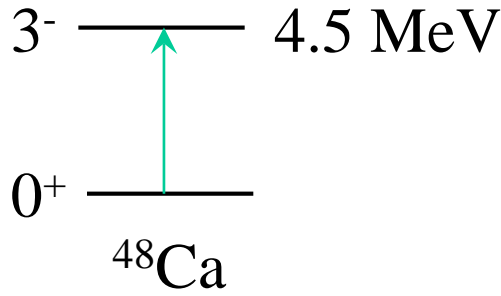
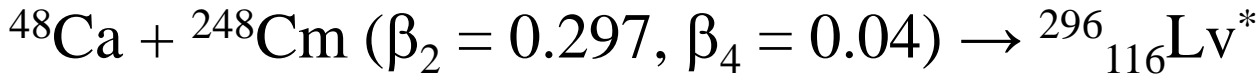
[β_2 and β_4 from P. Moller]

$$\frac{d\sigma_{\text{qel}}}{d\Omega} = \int_0^1 d(\cos\theta) \left(\frac{d\sigma_{\text{el}}}{d\Omega} \right)_\theta$$



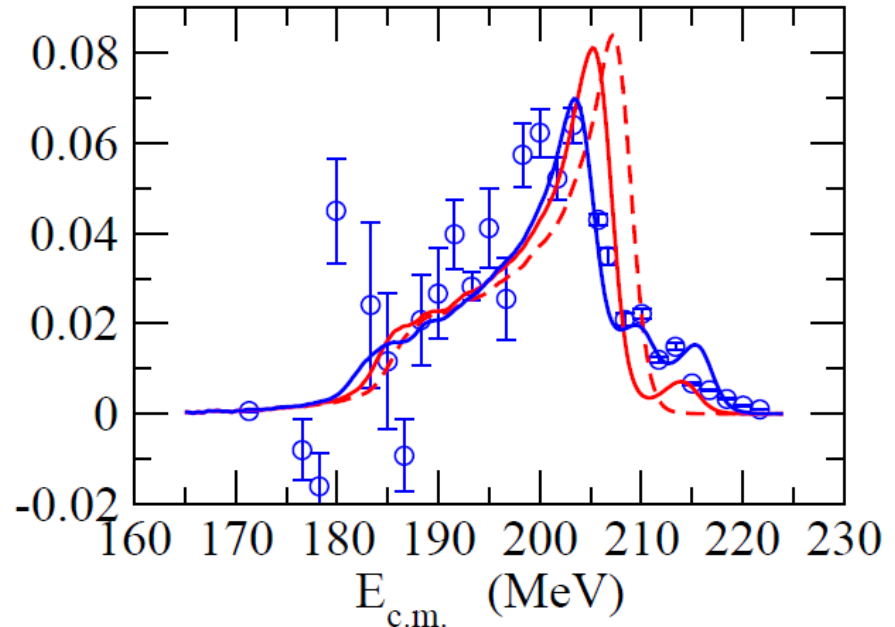
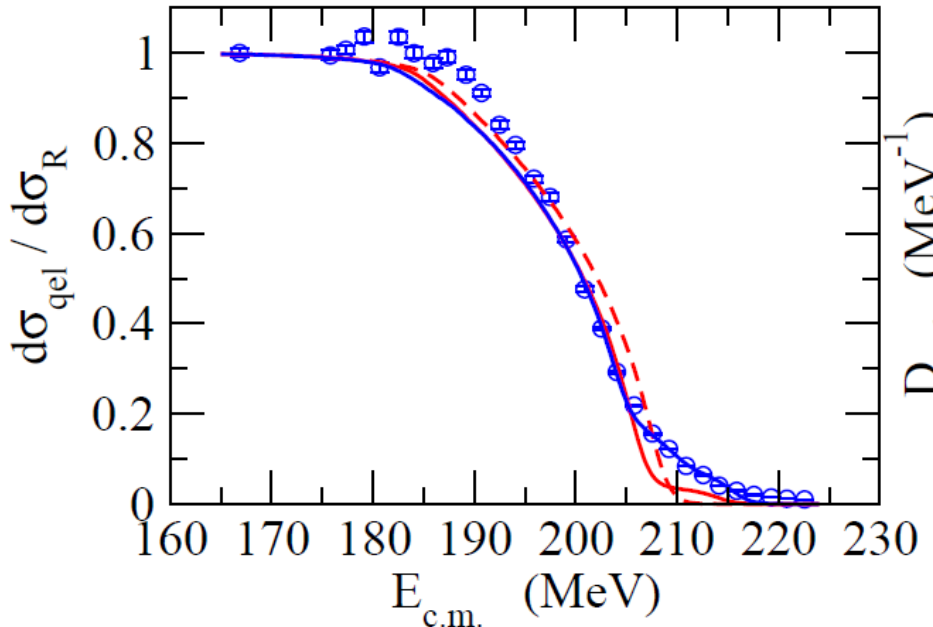
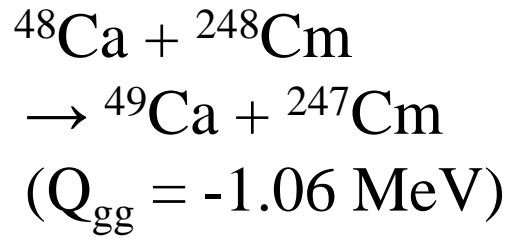
Analysis for a hot fusion reaction $^{48}\text{Ca} + ^{248}\text{Cm}$

K.H. and T. Tanaka (2017)

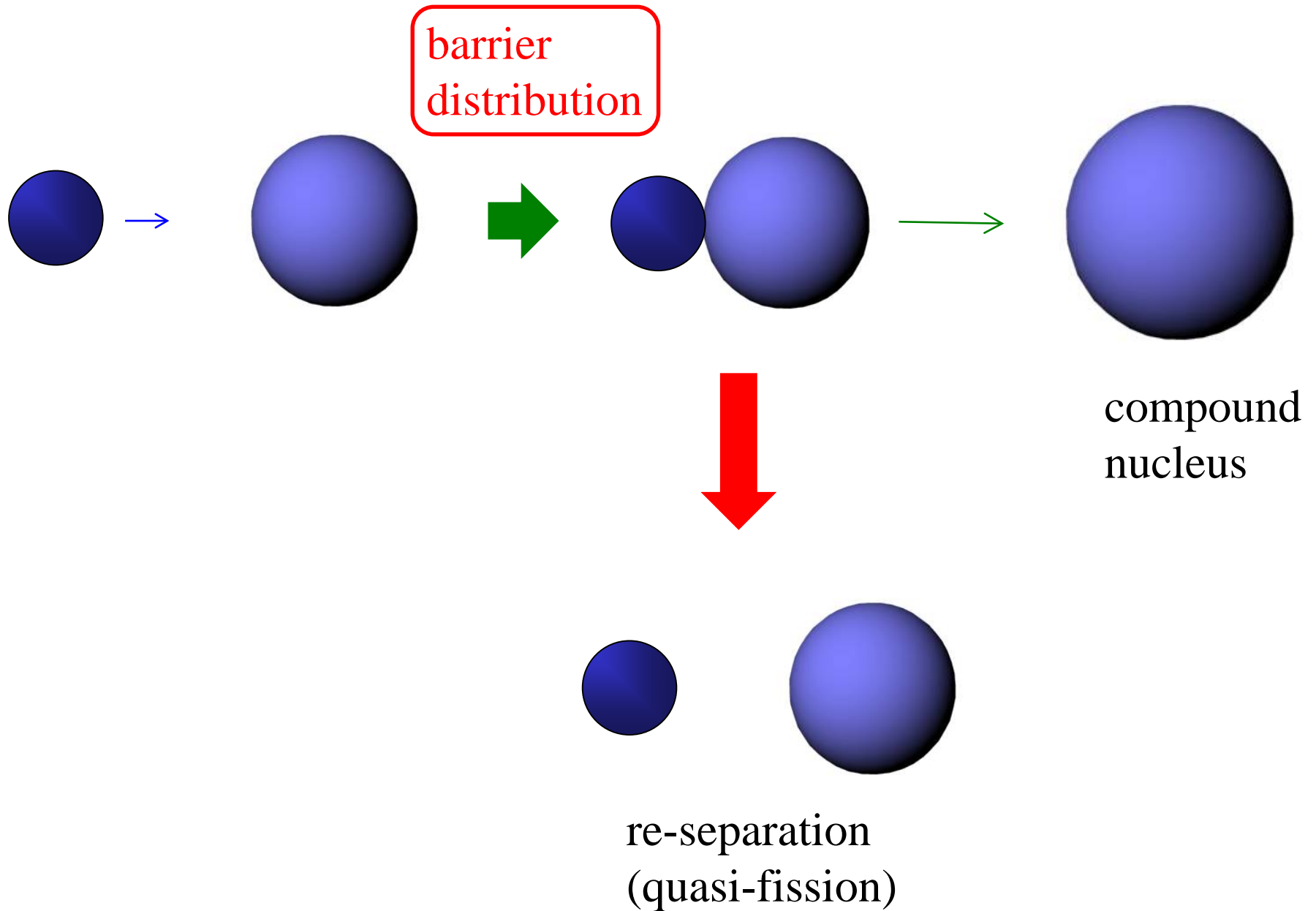


- def. of ^{248}Cm
- + $^{48}\text{Ca} (3^-)$
- + In transfer

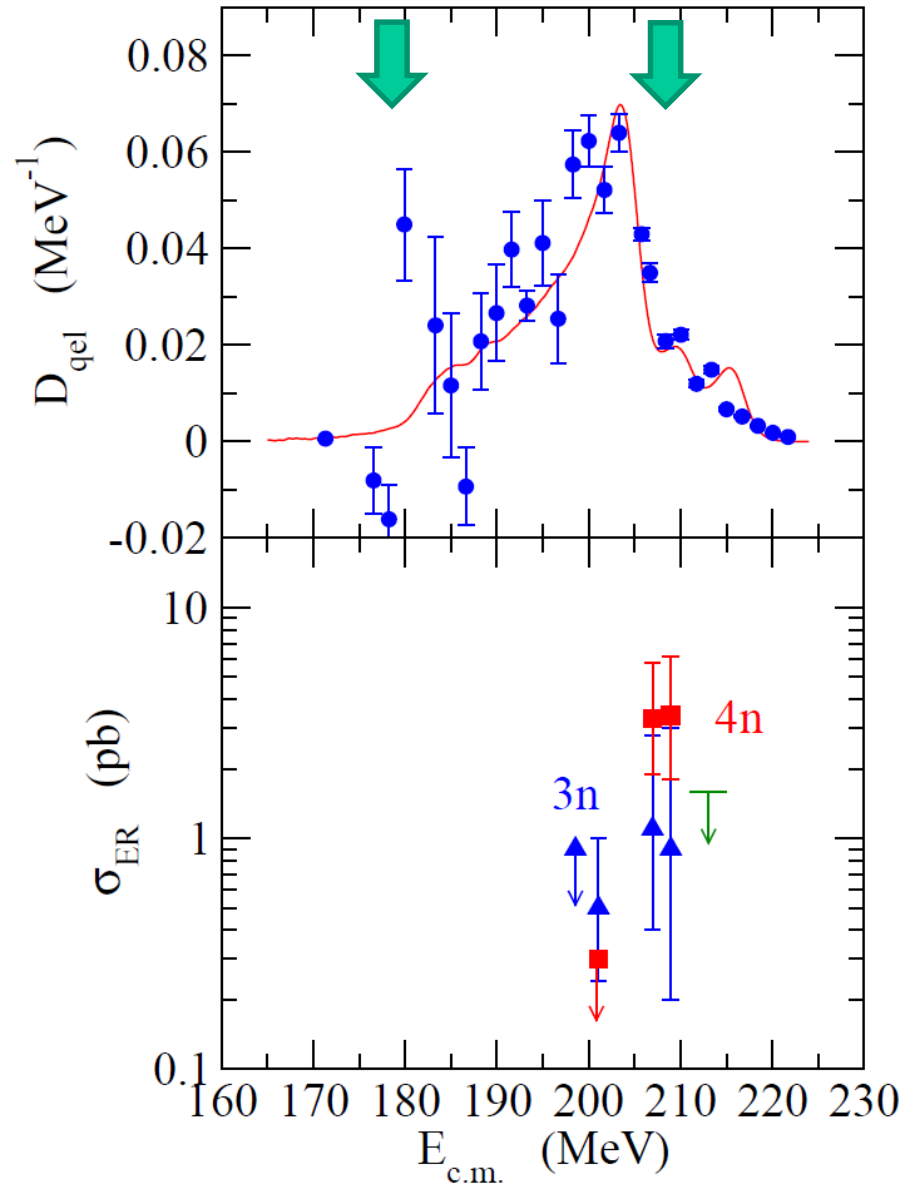
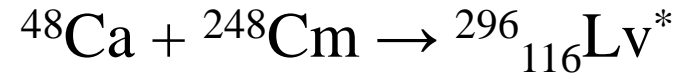
In transfer



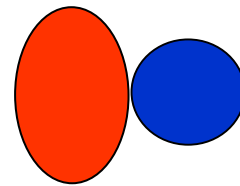
Connection to the ER cross sections



Connection to the ER cross sections



notion of compactness:
D.J. Hinde et al., PRL74 ('95) 1295



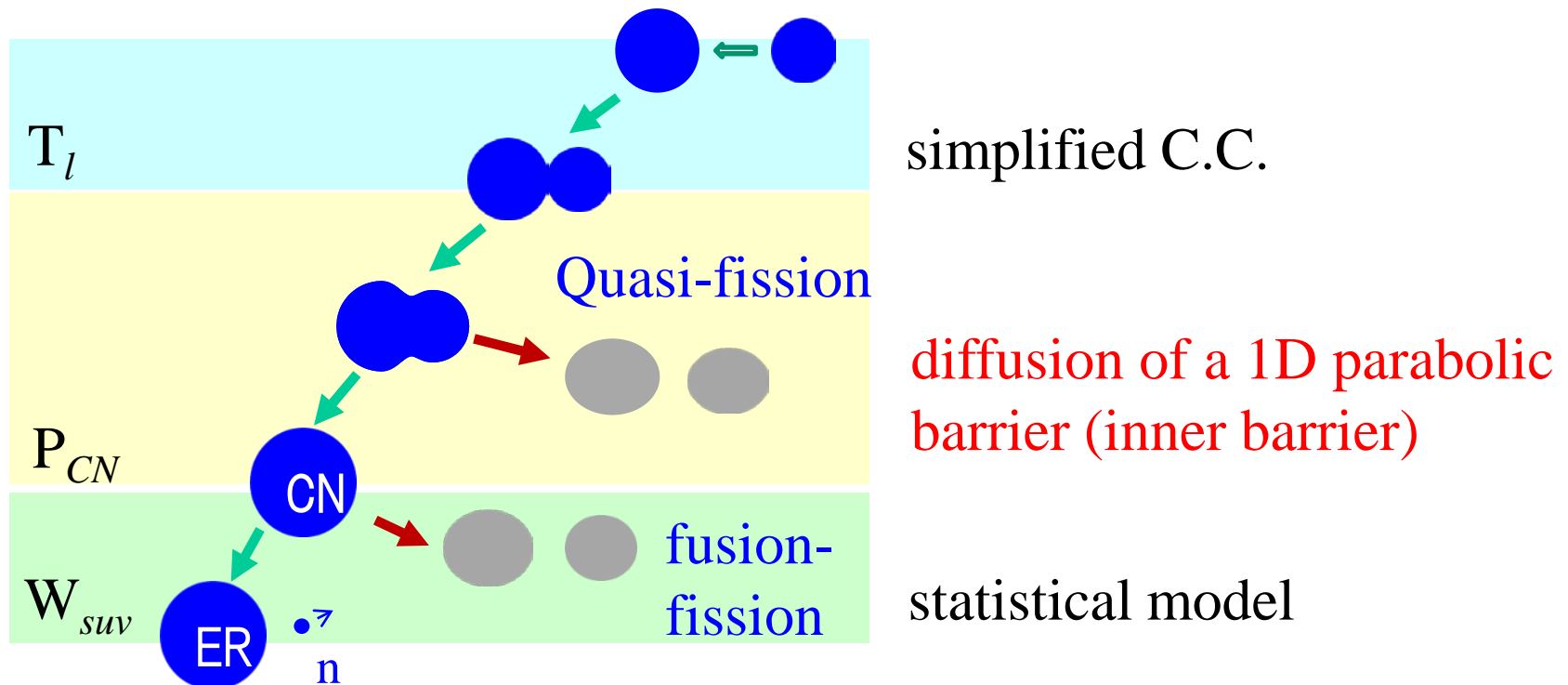
= more compact at the touching
→ favorable for CN

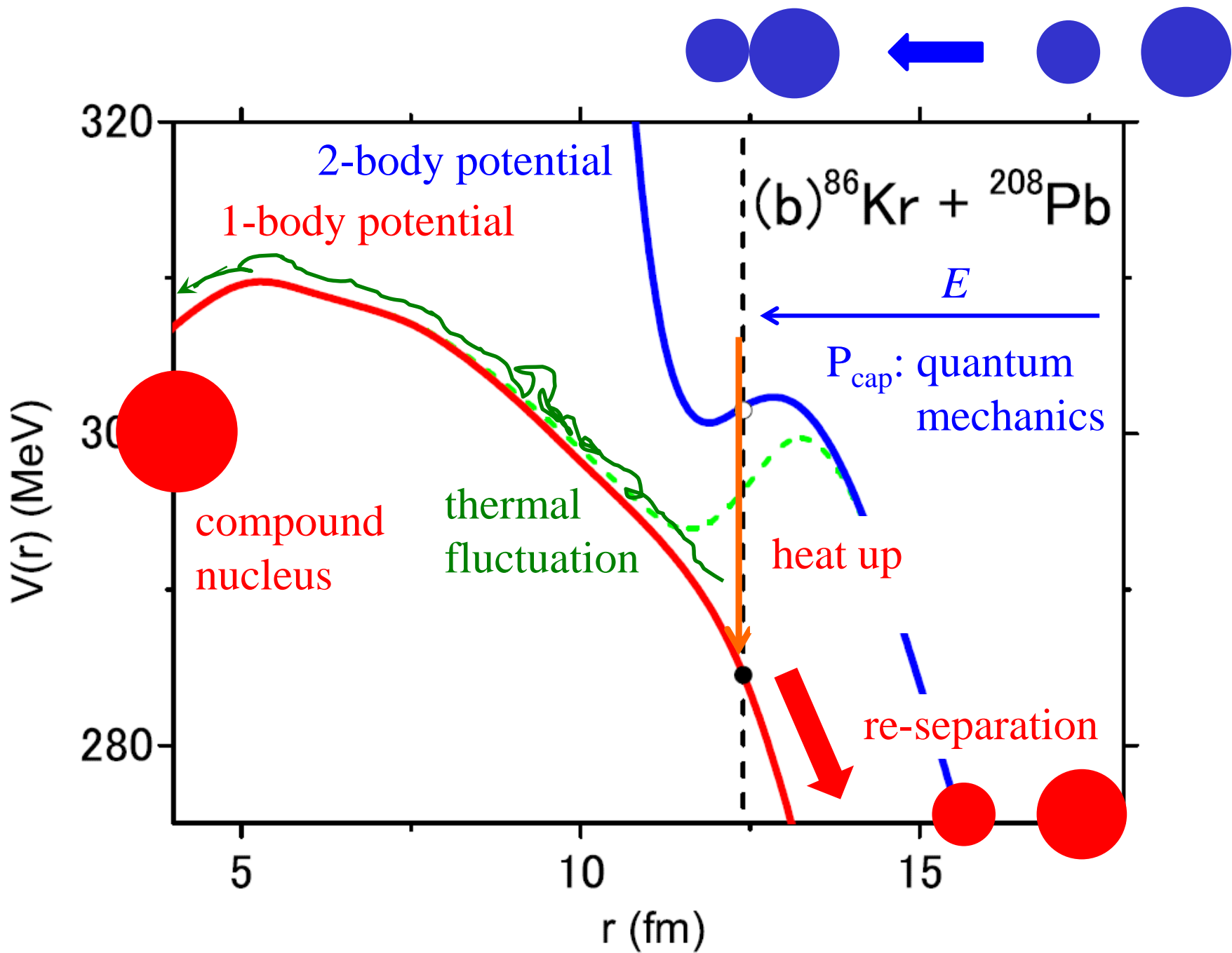
Extension of the fusion-by-diffusion model

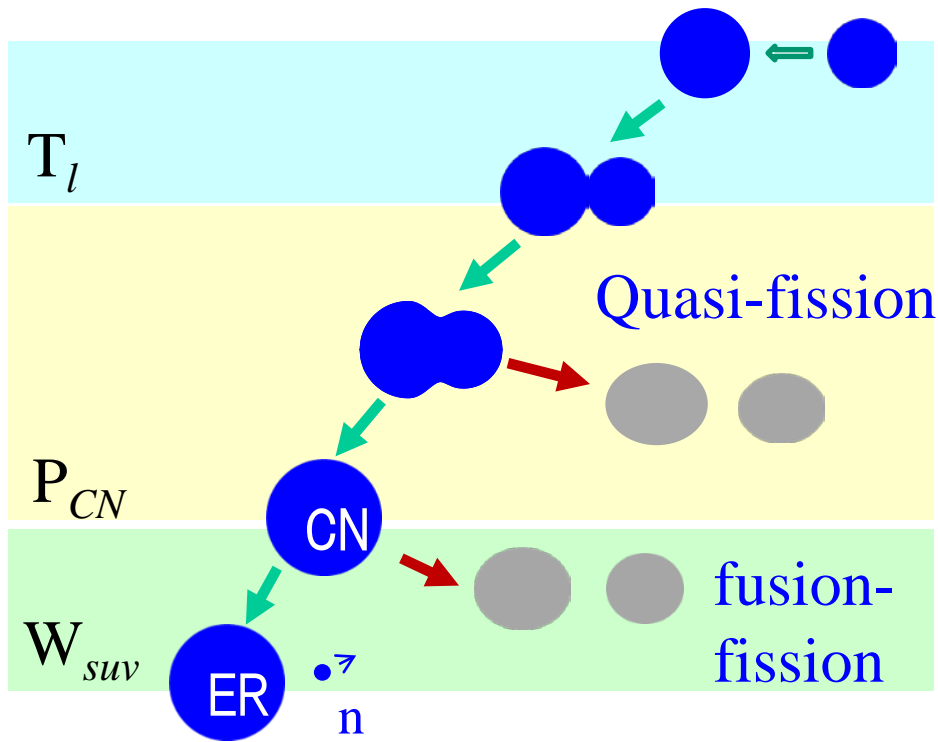
K.H., arXiv: 1803.02036

Fusion-by-diffusion model

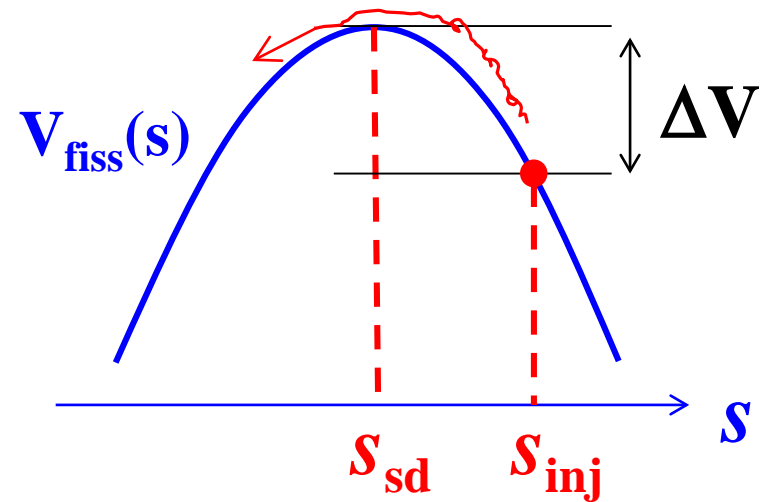
W.J. Swiatecki et al., Acta Phys. Pol. B34 ('03) 2049
PRC71 ('05) 014602





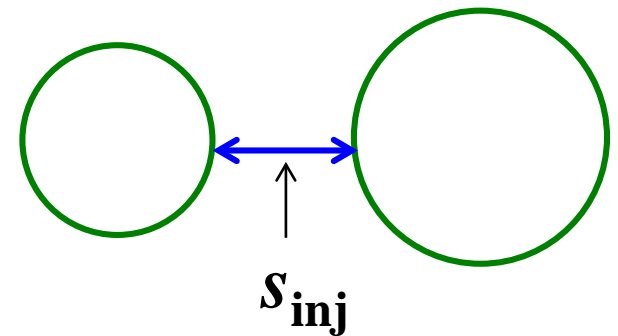


diffusion of a 1D parabolic barrier



Langevin in the overdamped limit:

$$P_{CN}(E) = \frac{1}{2} \left[1 - \operatorname{erf} \left(\frac{\Delta V}{T} \right) \right]$$



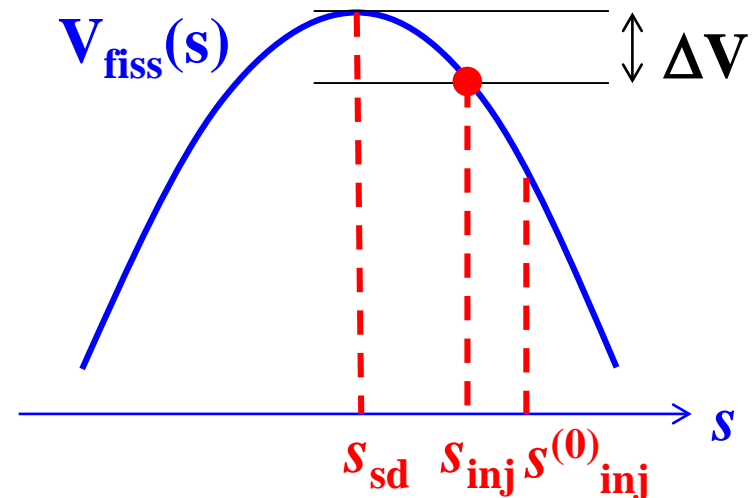
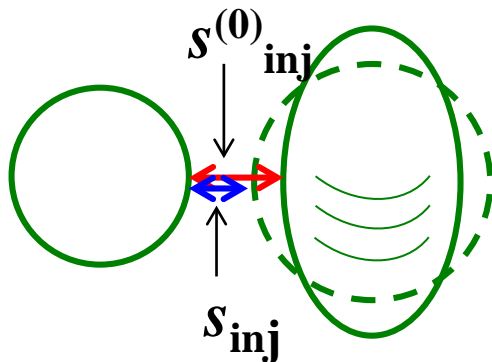
Extension of the fusion-by-diffusion model

K.H., arXiv: 1803.02036

$$s_{\text{inj}}(\theta) = s_{\text{inj}}^{(0)} + R_T \sum_{\lambda} \beta_{\lambda T} Y_{\lambda 0}(\theta)$$

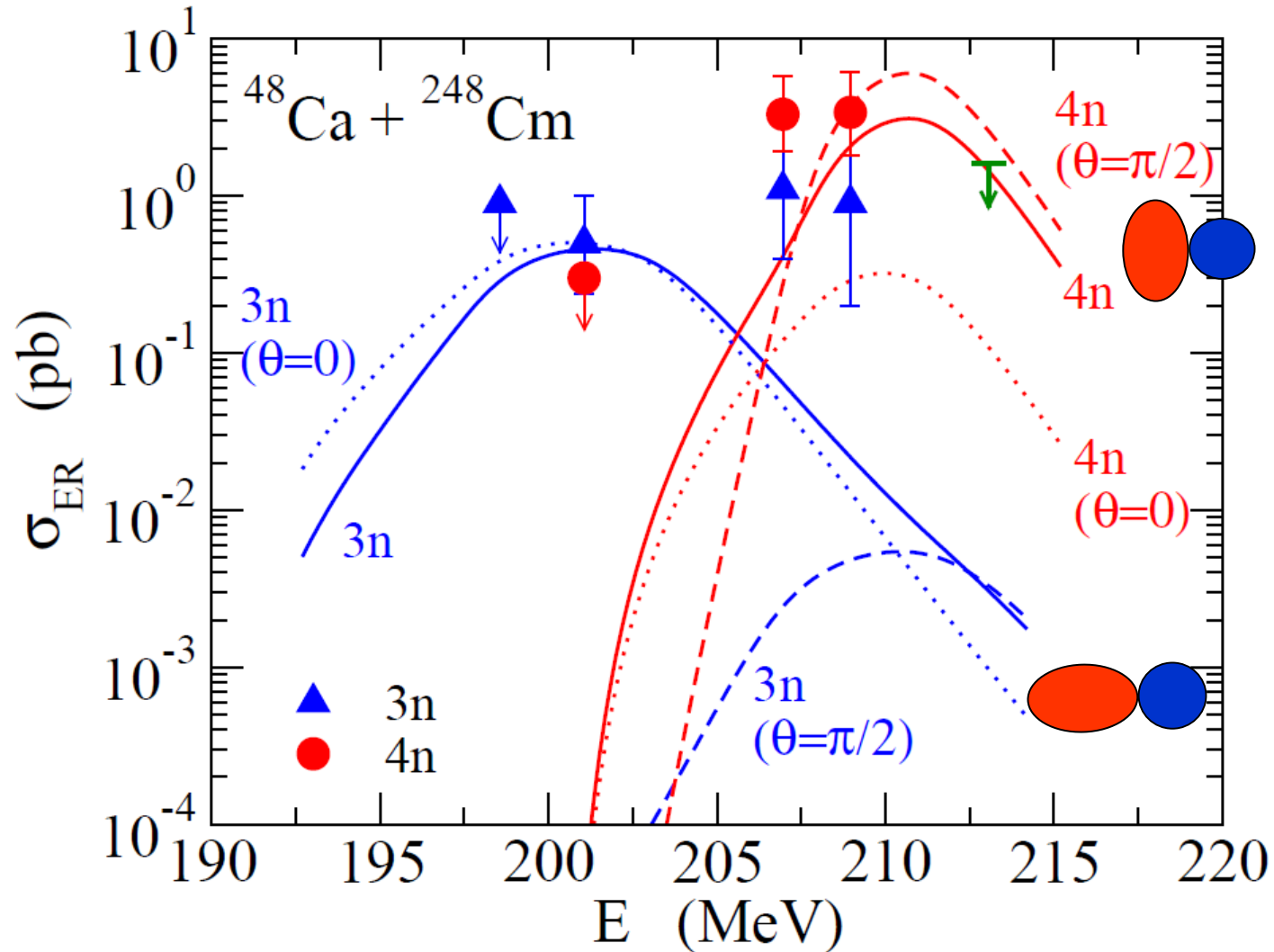
$$P_{\text{CN}}(E, \theta) = \frac{1}{2} \left[1 - \text{erf} \left(\frac{\Delta V(\theta)}{T(\theta)} \right) \right]$$

$\theta = \pi/2$ (side collision)

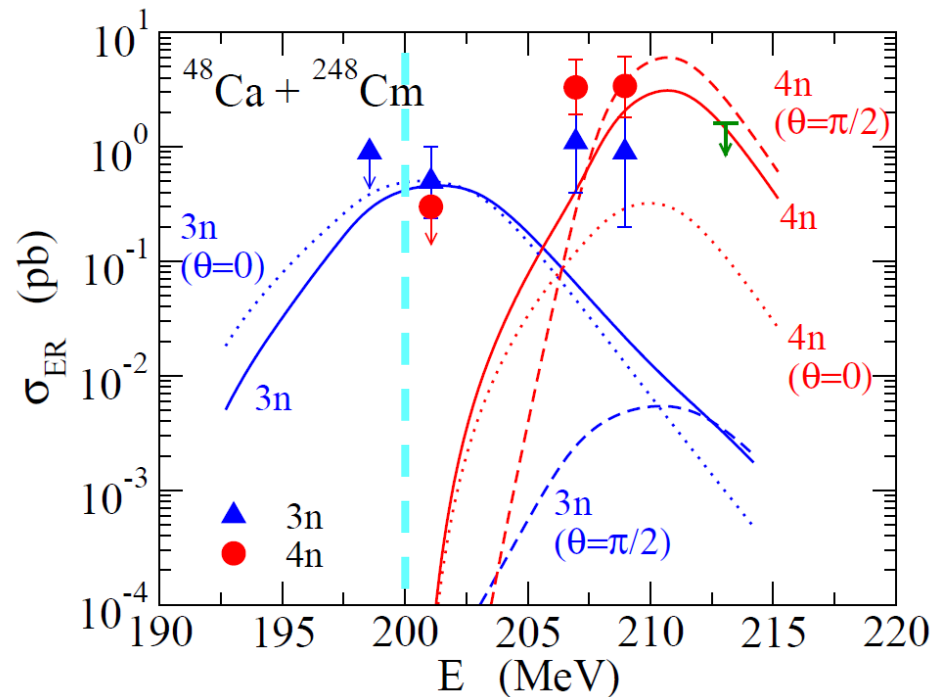
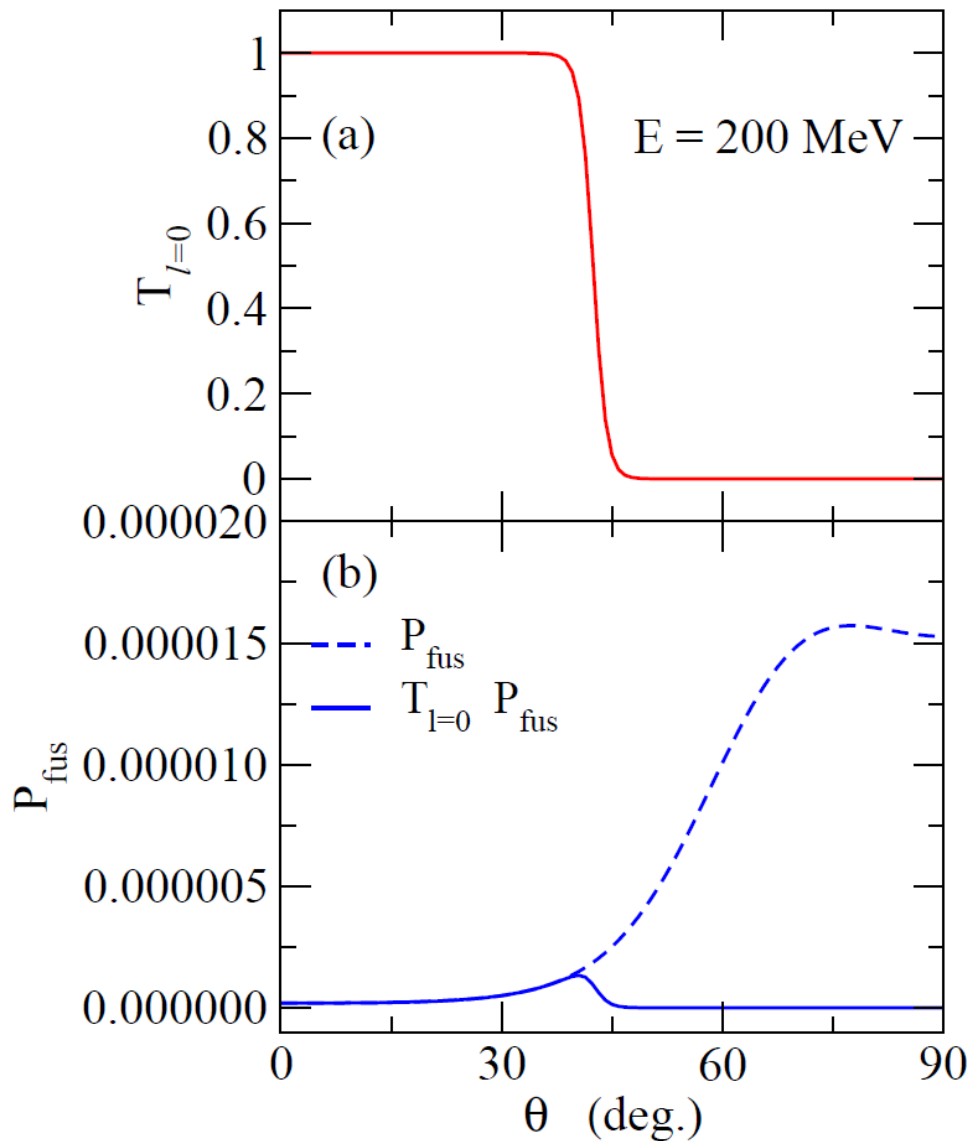


Extension of the fusion-by-diffusion model

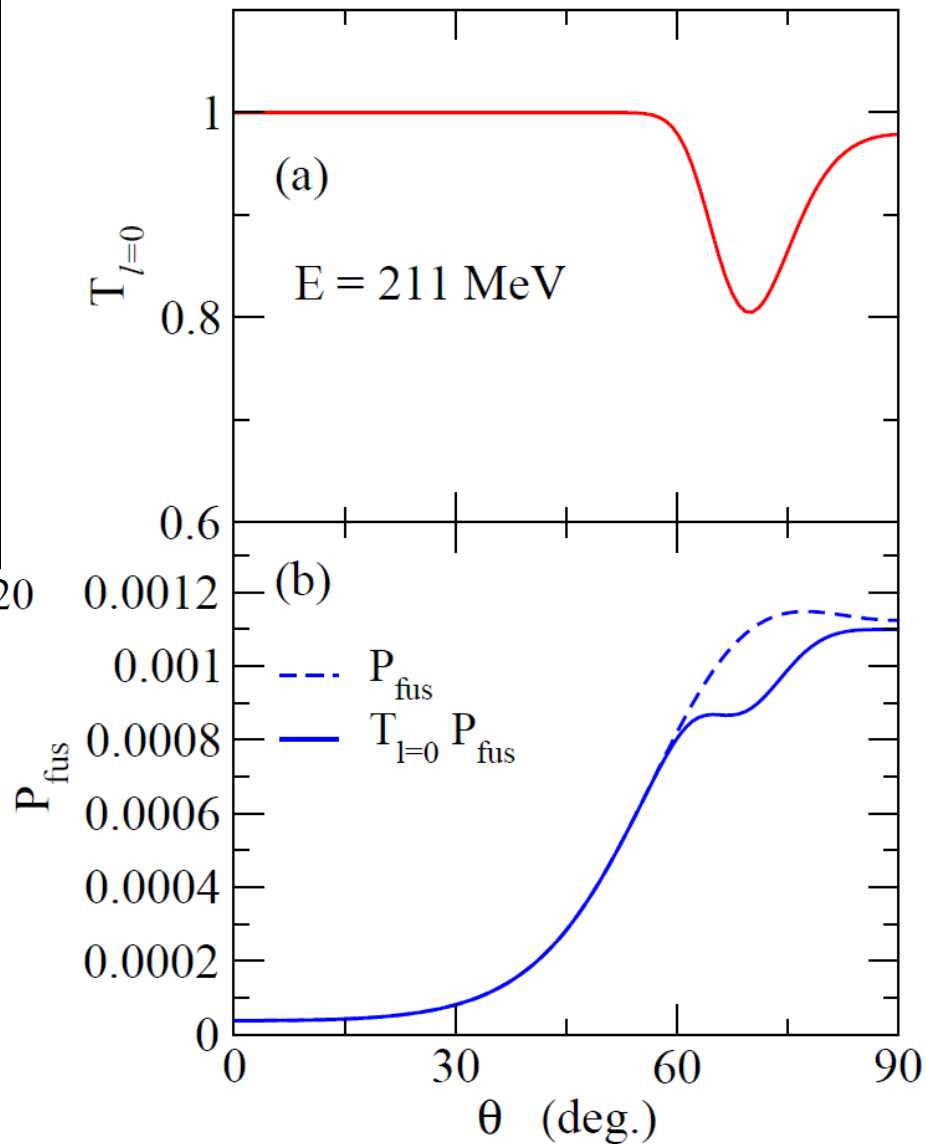
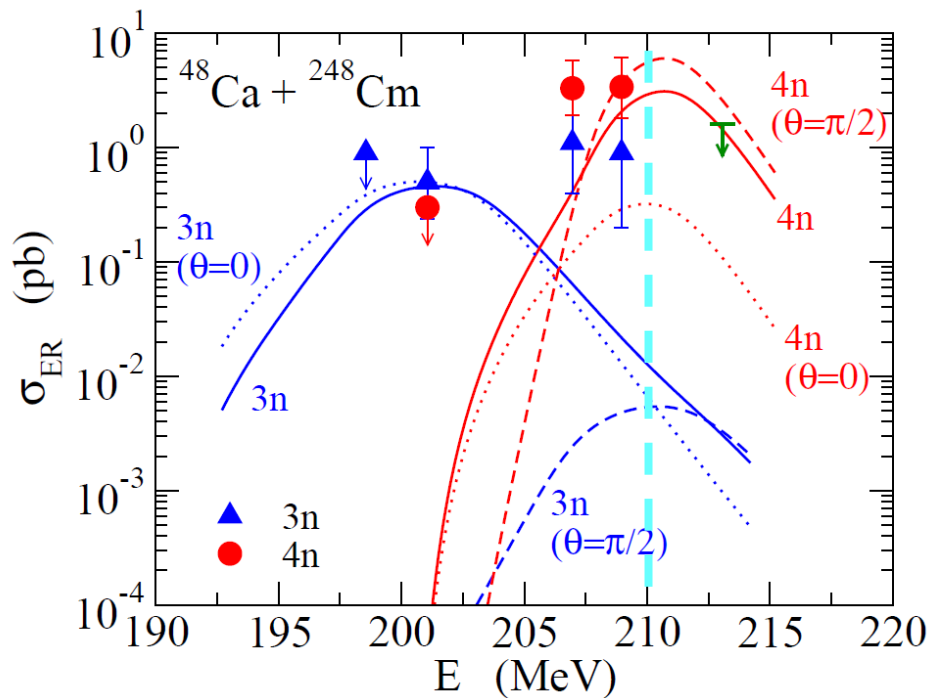
K.H., arXiv: 1803.02036



$$\sigma_{\text{ER}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) \int_0^1 d \cos \theta T_l(E, \theta) P_{\text{fus}}(E, l, \theta) W_{\text{sur}}(E^*, l)$$



$$\sigma_{\text{ER}}(E) = \frac{\pi}{k^2} \sum_l (2l + 1) \int_0^1 d \cos \theta T_l(E, \theta) P_{\text{fus}}(E, l, \theta) W_{\text{sur}}(E^*, l)$$



Summary and discussions

Reaction dynamics for SHE formation reactions

➤ Recent measurement of barrier distributions with GARIS

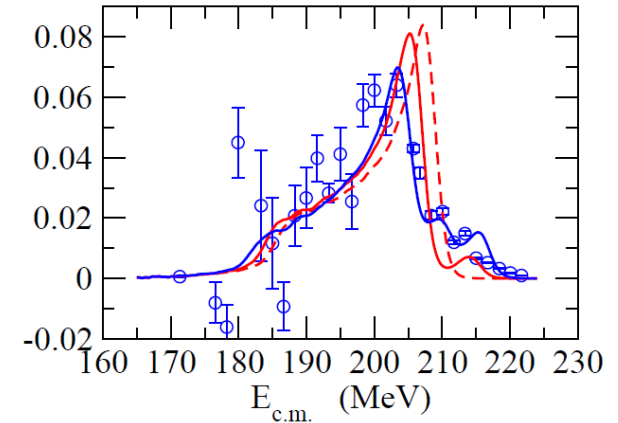
- ✓ $^{48}\text{Ca} + ^{248}\text{Cm}$
- ✓ coupled-channels analysis
- ✓ notion of compactness: ER formation with side collisions

more data coming soon

➤ Open problems

- ✓ reaction dynamics?

quantum theory for friction




cf. M. Tokieda and K.H.,
PRC95 ('17) 054604

Quantum friction

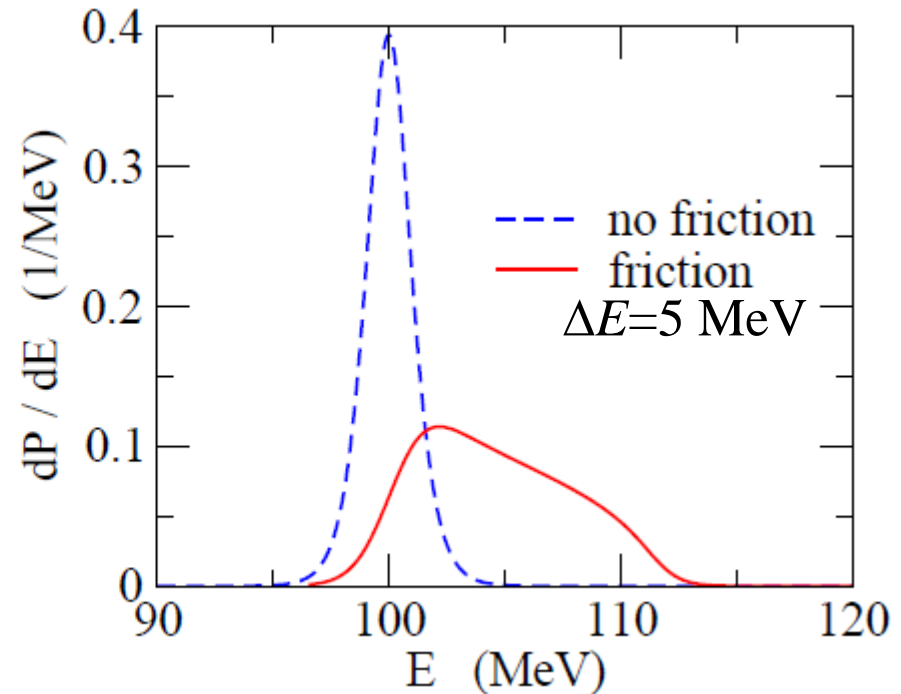
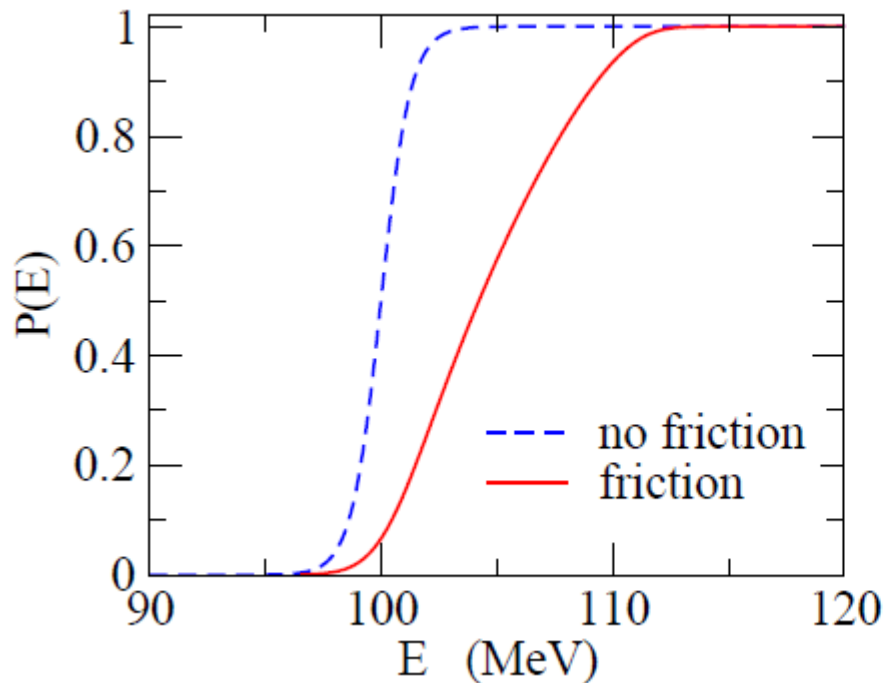
classical eq. of motion $\dot{p} = -V'(x) - \gamma p$

a quantization: Kanai model E. Kanai, PTP 3 (1948) 440

$$H = \frac{p^2}{2m} + V(x) \rightarrow \frac{\pi^2}{2m} e^{-\gamma t} + e^{\gamma t} V(x) \quad (\pi = e^{\gamma t} p)$$

 $\frac{d}{dt} \langle p \rangle = -\langle V'(x) \rangle - \gamma \langle p \rangle$

time-dep. wave packet approach



Summary and discussions

Reaction dynamics for SHE formation reactions

➤ Recent measurement of barrier distributions with GARIS

- ✓ $^{48}\text{Ca} + ^{248}\text{Cm}$
- ✓ coupled-channels analysis
- ✓ notion of compactness: ER formation with side collisions

more data coming soon

➤ Open problems

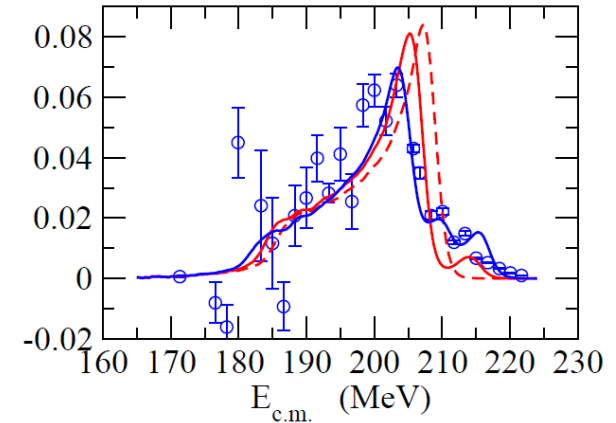
- ✓ reaction dynamics?
- ✓ shape evolution with a deformed target?

quantum theory for friction

how does the deformation disappear during heat-up?

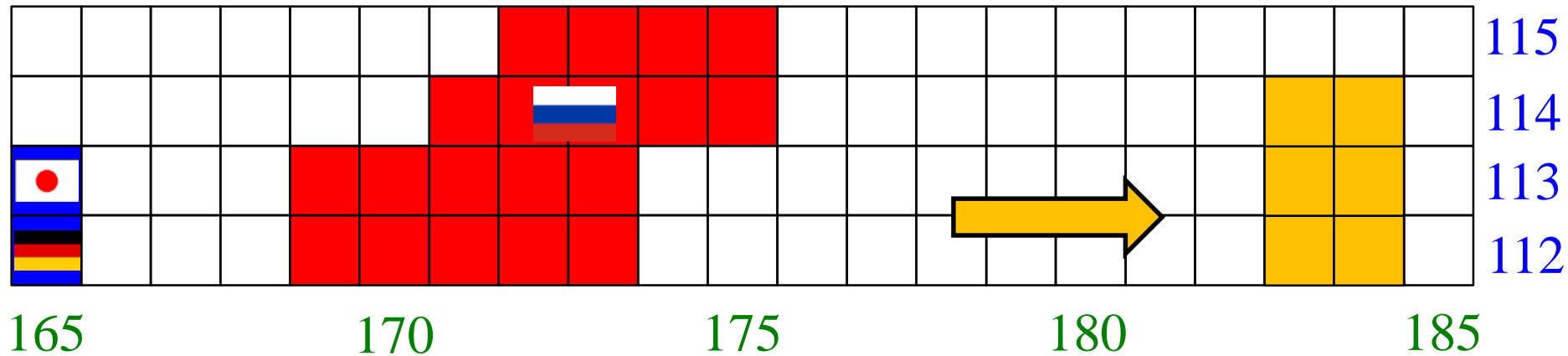
- ✓ towards island of stability

reaction dynamics with neutron-rich beams?



cf. M. Tokieda and K.H.,
PRC95 ('17) 054604

Towards the island of stability



neutron-rich beams: indispensable

- how to deal with low beam intensity?
- reaction dynamics of neutron-rich beams?
 - ✓ capture: role of breakup and (multi-neutron) transfer?
 - ✓ diffusion: neutron emission during a shape evolution?
 - ✓ survival: validity of the statistical model?

structure of exotic nuclei

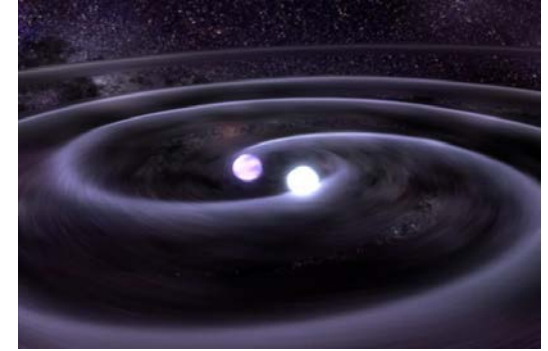
more studies are required

formation of SHE

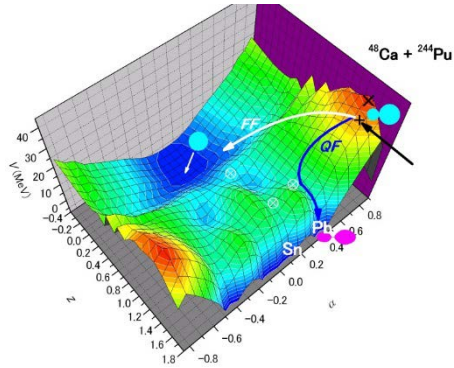


chemistry of SHE

Group →	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
Period ↓	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18
1	1 H																	2 He
2	3 Li	4 Be											5 B	6 C	7 N	8 O	9 F	10 Ne
3	11 Na	12 Mg											13 Al	14 Si	15 P	16 S	17 Cl	18 Ar
4	19 K	20 Ca	21 Sc	22 Ti	23 V	24 Cr	25 Mn	26 Fe	27 Co	28 Ni	29 Cu	30 Zn	31 Ga	32 Ge	33 As	34 Se	35 Br	36 Kr
5	37 Rb	38 Sr	39 Y	40 Zr	41 Nb	42 Mo	43 Tc	44 Ru	45 Rh	46 Pd	47 Ag	48 Cd	49 In	50 Sn	51 Sb	52 Te	53 I	54 Xe
6	55 Cs	56 Ba	57 La	* 72 Hf	73 Ta	74 W	75 Re	76 Os	77 Ir	78 Pt	79 Au	80 Hg	81 Tl	82 Pb	83 Bi	84 Po	85 At	86 Rn
7	87 Fr	88 Ra	89 Ac	* 104 Rf	105 Db	106 Sg	107 Bh	108 Hs	109 Mt	110 Ds	111 Rg	112 Cn	113 Nh	114 Fl	115 Mc	116 Lv	117 Ts	118 Og
				* 58 Ce	59 Pr	60 Nd	61 Pm	62 Sm	63 Eu	64 Gd	65 Tb	66 Dy	67 Ho	68 Er	69 Tm	70 Yb	71 Lu	
				* 90 Th	91 Pa	92 U	93 Np	94 Pu	95 Am	96 Cm	97 Bk	98 Cf	99 Es	100 Fm	101 Md	102 No	103 Lr	



reaction dynamics



Nuclear Physics
(RIBF/FRIB)
Astrophysics

structure of SHE

interdisciplinary SHE science

with physics, chemistry, and astronomy

FUSION20

November 16-20, 2020

Shizuoka, Japan

Kouichi Hagino (co-chair) Tohoku University

Katsuhisa Nishio (co-chair) JAEA



